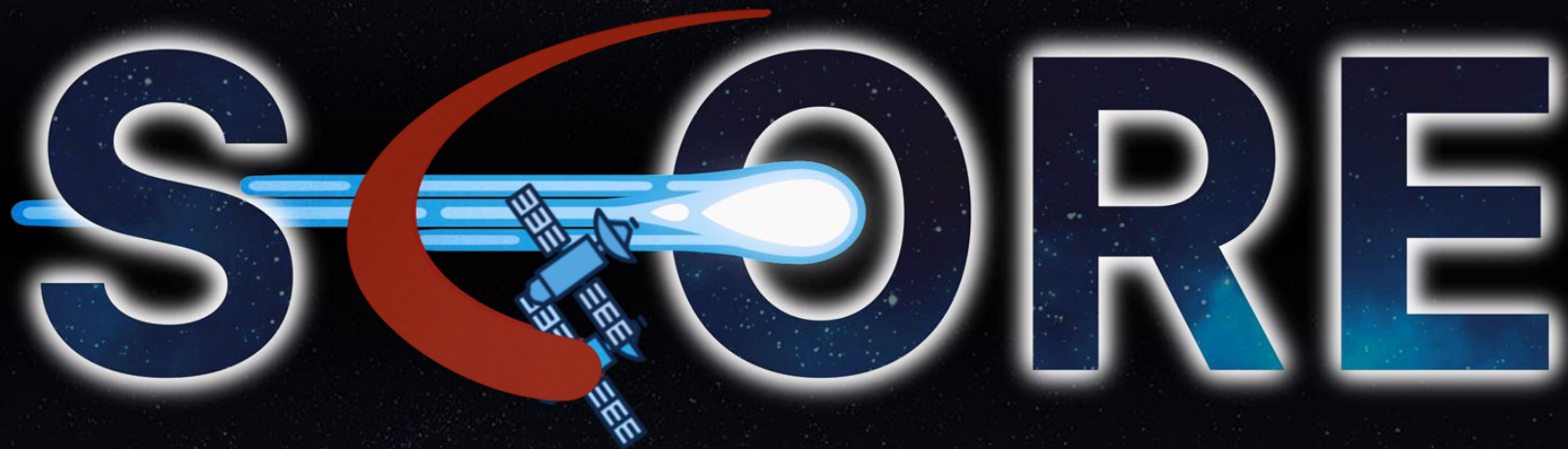
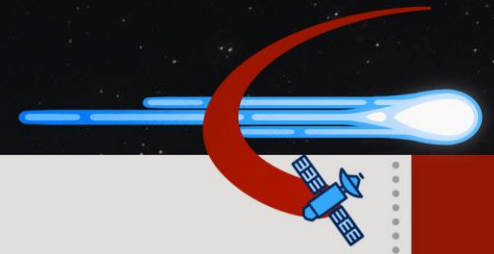


SCORE

The logo for the SCORE mission is centered on a dark blue, star-speckled background. The word "SCORE" is rendered in large, bold, blue letters with a white glow effect. The letter 'C' is stylized as a red comet tail. The letter 'O' is stylized as a blue comet tail. A blue satellite with solar panels is positioned between the 'C' and 'O', appearing to fly through the comet tails.

Scientific mission for Comet Observation, Research and Exploration

WHY LONG PERIOD COMETS?



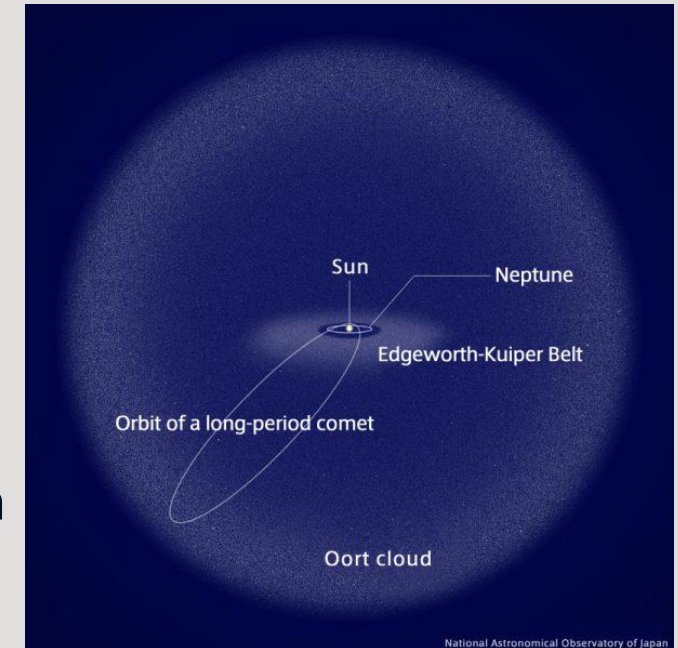
- Long Period Comets are celestial bodies with an orbital period greater than 200 years.
- Long Period Comets have **not been contaminated** by the various passages in the Inner Solar System.
- Long Period Comets have never been deeply explored by previous missions.



- They can bring new important results regarding the **origin and the evolution** of the Solar System and of life on Earth
- Deeper knowledge about the **cometary environment**
- Deeper knowledge about **comets**

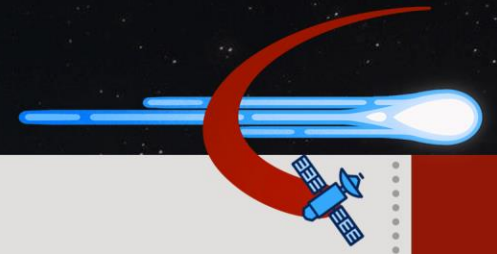


SCORE has been designed with the purpose of intercepting a Long Period Comet in a **flyby scenario** with the use of a **micro-satellite**



Credits: National Astronomical Observatory of Japan (NAOJ)

WHY SCORE?



SCORE makes use of a **micro-satellite** with limited mass and size.

- Deep space exploration missions have been dominated by large satellites.
- **Micro-satellites** have been popular in the last decades, but mostly for *Earth Orbiting systems*.

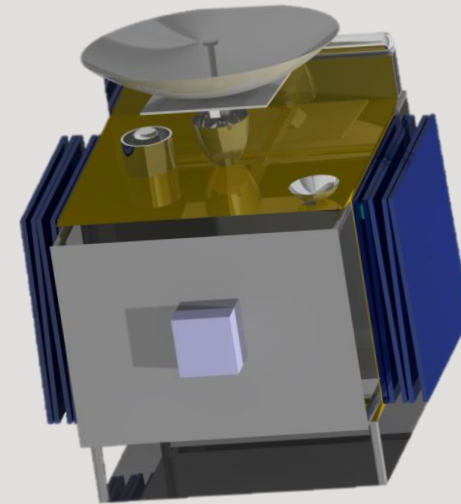


SCORE mission will be an **important test bench for new technologies and subsystems** for micro/nano-satellites:

- Only few others deep space exploration missions used micro-satellites (e.g. MarCO mission)
- **Electric Propulsion** is an innovative technology that can highly increase the maximum distances reachable by micro-satellites.
- The **Shielding against micrometeoroids** provides reliable protection with a limited mass.



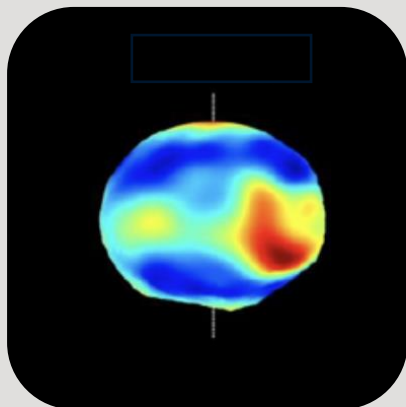
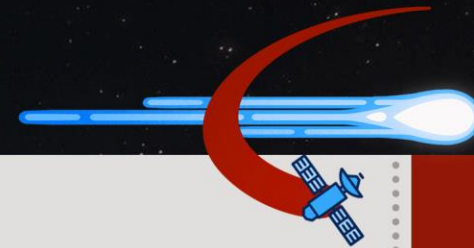
HT400 Electric Thruster





MISSION OBJECTIVES

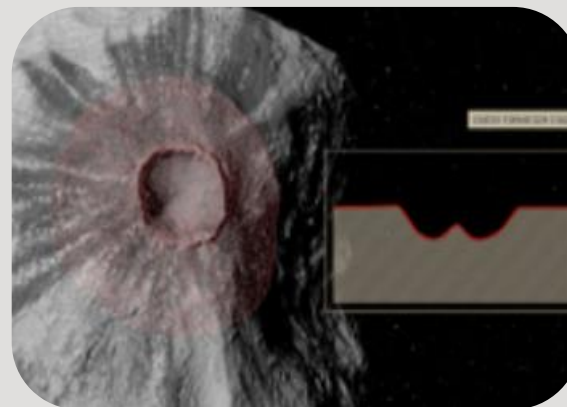
MISSION OBJECTIVES



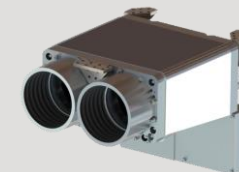
Nucleus Characteristic



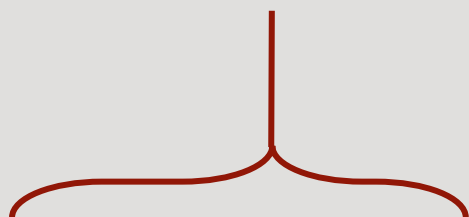
Coma Composition



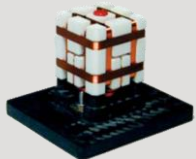
Morphology



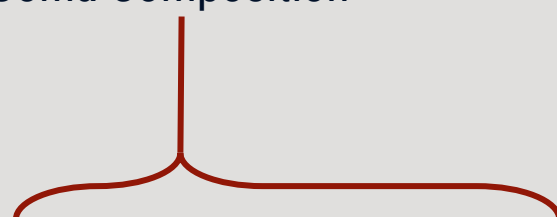
MULTISPECTRAL CAMERA



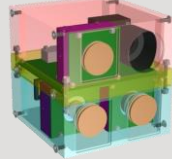
THERMAL CAMERA



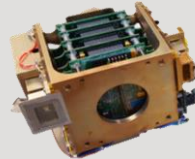
FLUXGATE MAGNETOMETER



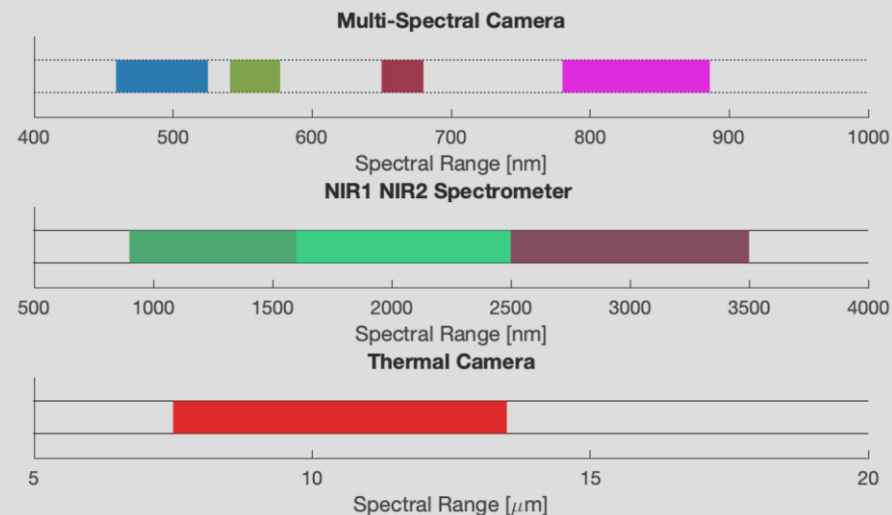
IMPACT SENSOR



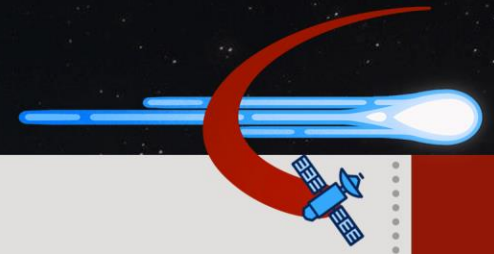
IR SPECTROMETER



MASS SPECTROMETER



MISSION OBJECTIVES



One of main difficulties:

- The **target** Long Period Comet is **uncertain**
- It can be discovered only some years prior to its passage at the **perihelion**



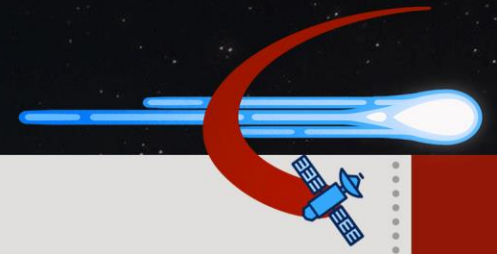
Need of **back-up targets** to guarantee a meaningful scientific return of the mission

Back-up target	Orbital period (years)	q (AU)	e	i (deg)	Flyby date (mm/yyyy)	Encountered by:
67P/Churyumov-Gerasimenko	6.45	1.24	0.641	7.04	09/2034	Rosetta (2014)
8P/Tuttle	13.61	1.03	0.820	54.98	03/2035	/
26P/Grigg-Skjellerup	5.31	1.12	0.633	22.36	06/2034	Giotto (1992)
9P/Tempel	5.58	1.54	0.510	10.47	04/2033	Deep Impact (2005) Stardust (2011)
7P/Pons-Winnecke	6.37	1.26	0.634	22.29	09/2033	/

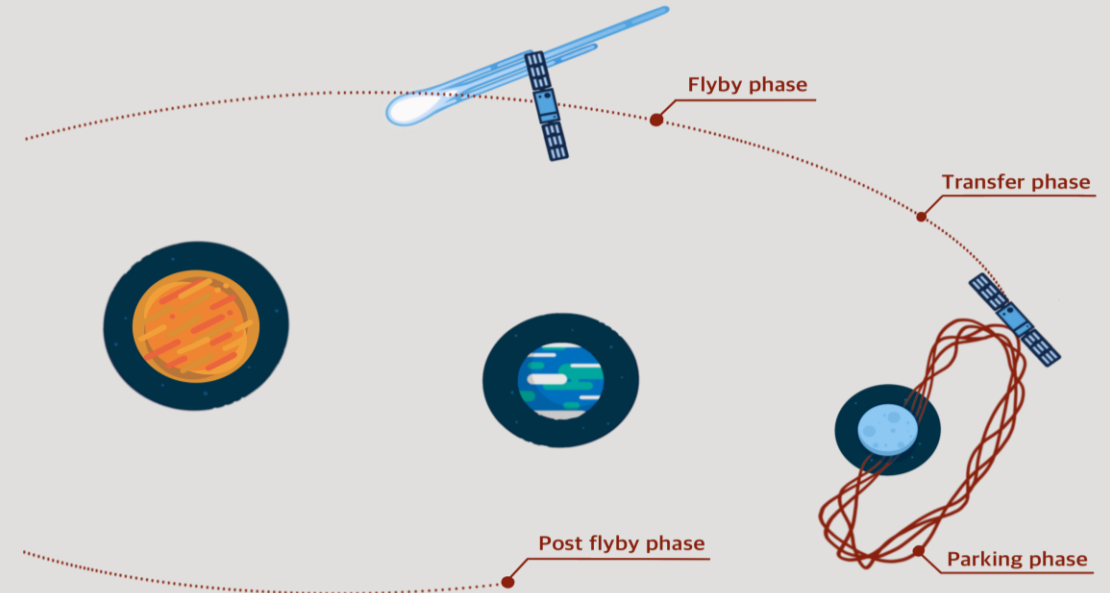


MISSION PHASES

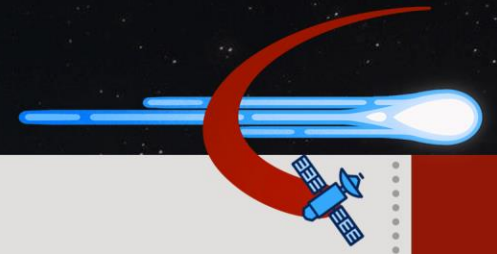
MISSION ARCHITECTURE



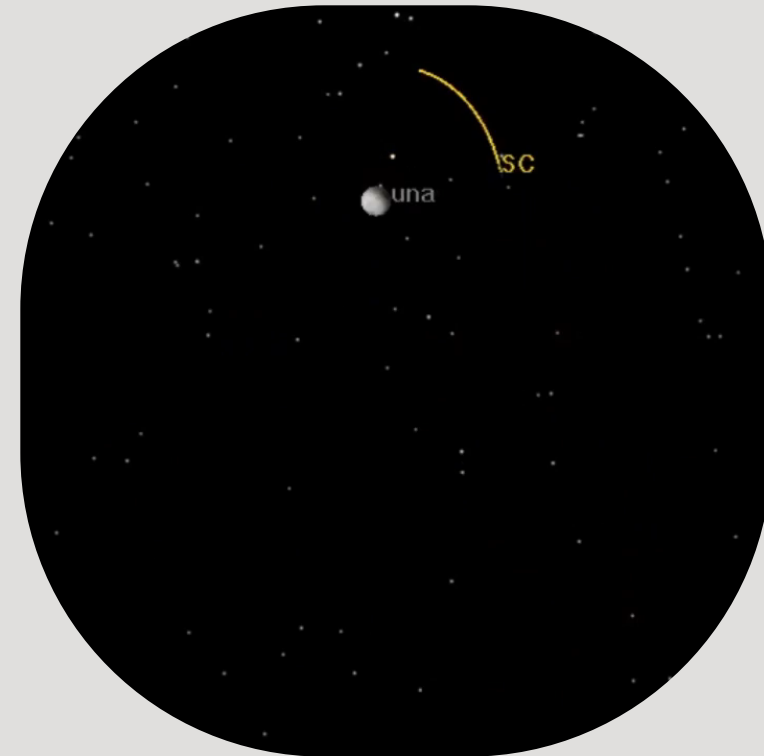
- Mission deployed **during 2029** in the cis-lunar space
- Single spacecraft with a wet mass of 100 kg
- **Region of interest** for the flyby = Sun-centred annulus with radius in between 1 and 1.58 AU
- Mission deployed on an **L2 southern Near Rectilinear Halo Orbit (NRHO)**
- The **lifetime** of the mission shall not exceed **6 years**



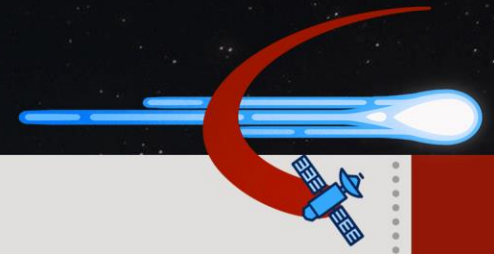
PARKING PHASE



- 16000 km **perilune radius** is chosen for the starting orbit with a **period** of 9.65 days.
- The **NRHO** is stable for **more than 1000 days** without station keeping maneuvers.
- The starting orbit exhibits a **3:1 resonance**.
- Eclipses occur every **1-2 lunar synodic periods** and their duration is short (~ 160 min).
- The **magnitude** of the **Station Keeping maneuvers** is very low (1 m/s per year).



PROBABILISTIC ANALYSIS



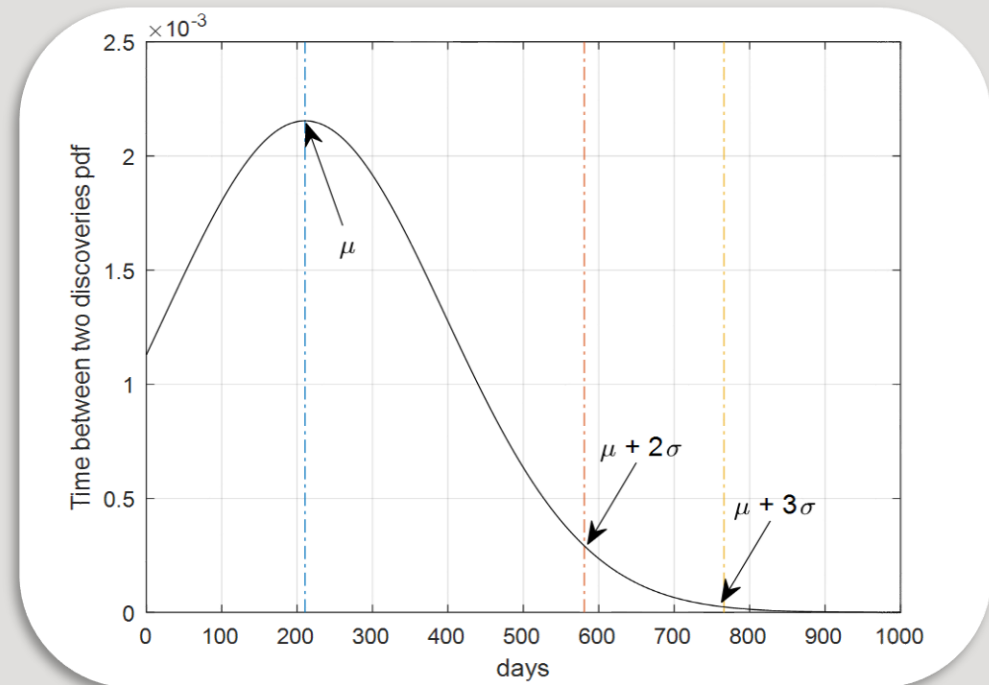
Probabilistic analysis starting from the orbital parameters of **already known** comets necessary to derive a possible orbit of a **generic LPC** that will pass through the plane from 2030.

The **time between two discoveries** has been evaluated from this probabilistic analysis.

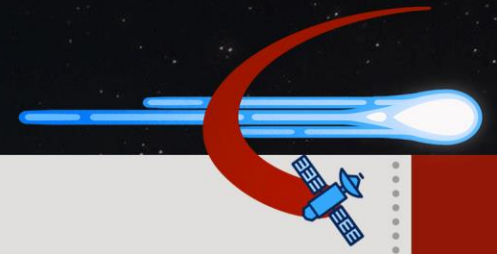
The *maximum* time between two discoveries is equal to *2.1 years*.

If in this period **no targets** are discovered a transfer towards a *back-up target* shall be planned.

μ [days]	210
σ [days]	185
Maximum time between two discoveries [days]	765
Maximum time between two discoveries [years]	2.1



DEPARTURE PHASE



DEPARTURE FROM THE MOON

- The Moon shall be *in between* the Sun and the Earth.
- High NRHO perilune radius → high altitudes comparable to the Moon *Sphere of Influence (SOI)*.

Suitable target detected in 2.1 years → satellite leaves the NRHO.

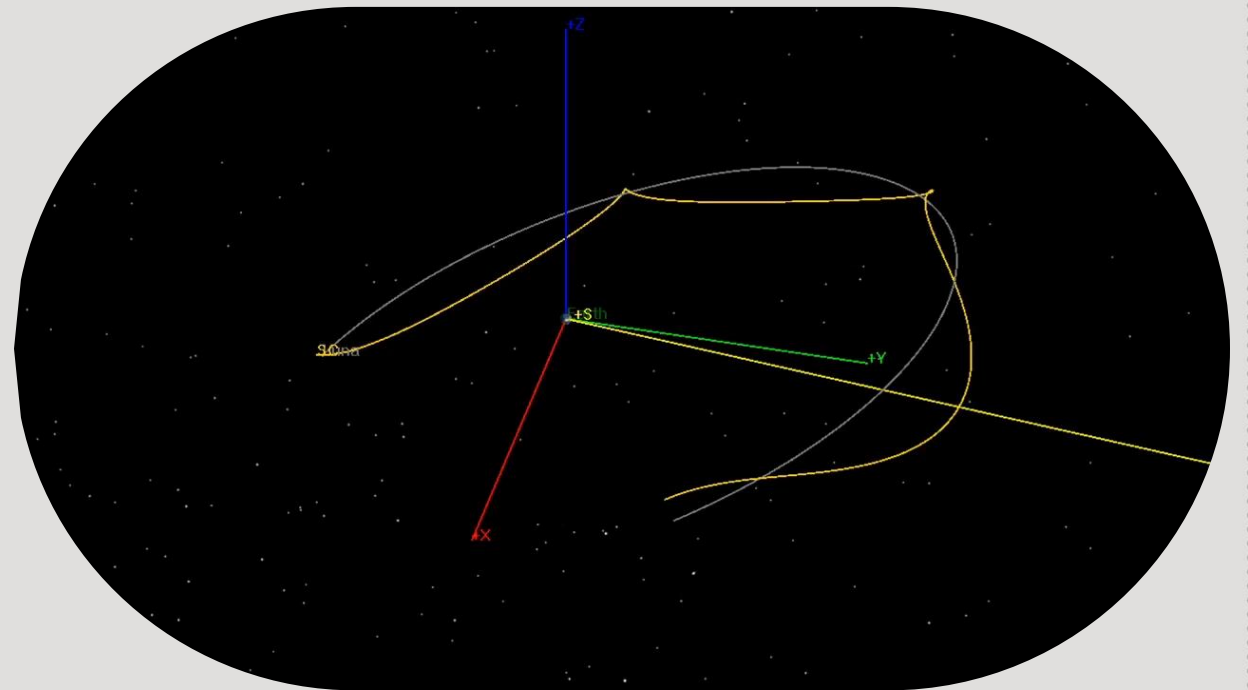
Departure maneuver opportunity once every 29.5 days

Departure performed with **chemical propulsion**
(monomethylhydrazine + mixed oxide of nitrogen MON-3)

- Departure delta-V to leave the Moon SOI is low (30-40 m/s).
- Departure delta-V to leave the Earth SOI is of about 200 m/s.

DEPARTURE FROM THE EARTH

- Impulsive maneuver to enter on a *transfer Ellipse*.
- Boundary of the *Earth's SOI* reached in **30.5 days**.
- Impulse of **small magnitude** concludes the departure phase → the *interplanetary transfer* towards begins.



TRANSFER PHASE

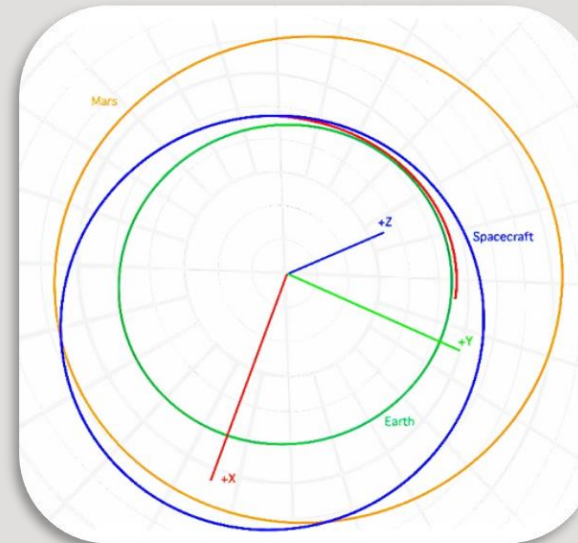
The interplanetary transfer is performed with **electrical propulsion** (*Xenon propellant*).

Two strategies are proposed:

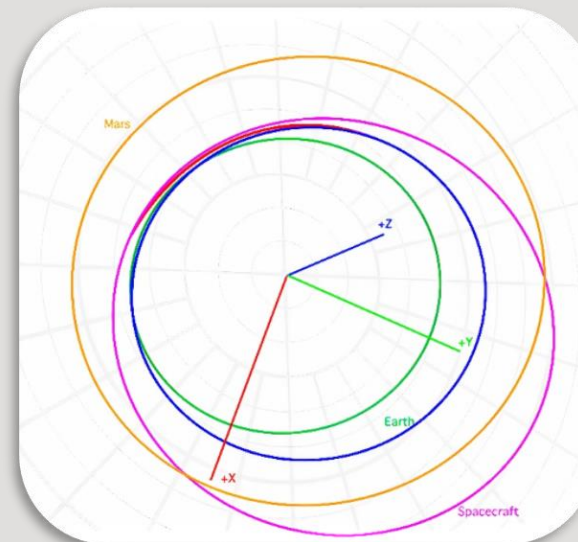
1. **Single finite burn** – Reaching a maximum distance of 1.39 AU from the Sun in about 1.5 years.
2. **Double finite burn** – Reaching a distance up to 1.58 AU from the Sun. The *maximum transfer time is of 2.6 years*. The *delta-V* necessary for this strategy is of about 3250 m/s.

The two strategies differ in terms of **transfer time**, with the first requiring *less time* than the second but covering a *smaller region*.

Depending on the **particular mission scenario**, one strategy may be preferable to the other



Single finite burn strategy



Double finite burn strategy

FLYBY PHASE

Flyby start: Payload instruments detect the first scientific data
Flyby end: Payload instruments not able to detect data anymore.
Maximum data obtainable: ~350 GB

Main role: relative velocity of the satellite with respect to the comet.

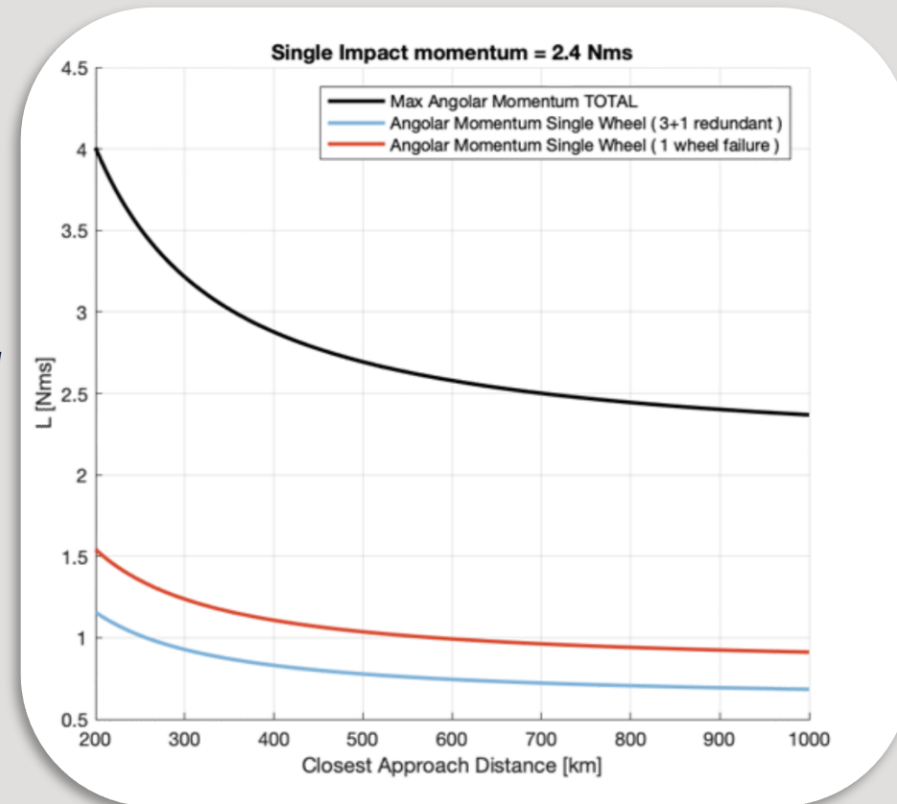
Closest approach (CA) distance: ~600 km (*nominally*)

Stabilization ensured for:

- Small particles impacts (≤ 10 mg) striking uniformly the satellite
- 100 mg particles impacts up to a CA distance of 600 km for *1 wheel failure*

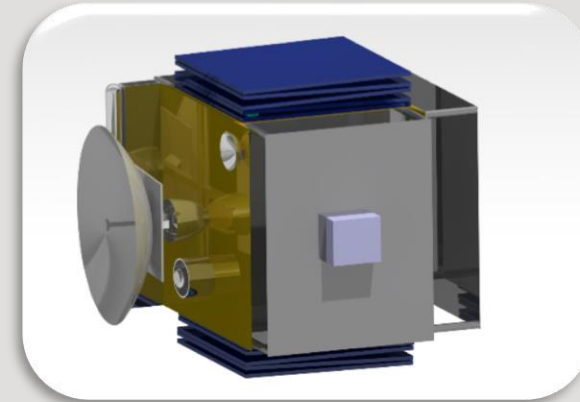
Chosen reaction wheels provide 1 N m s angular momentum

Maximum relative velocity	~70 km/s
Average relative velocity	~42 km/s
Average satellite heliocentric velocity	~25 km/s
Maximum satellite heliocentric velocity	~30 km/s



COMETARY ENVIRONMENT

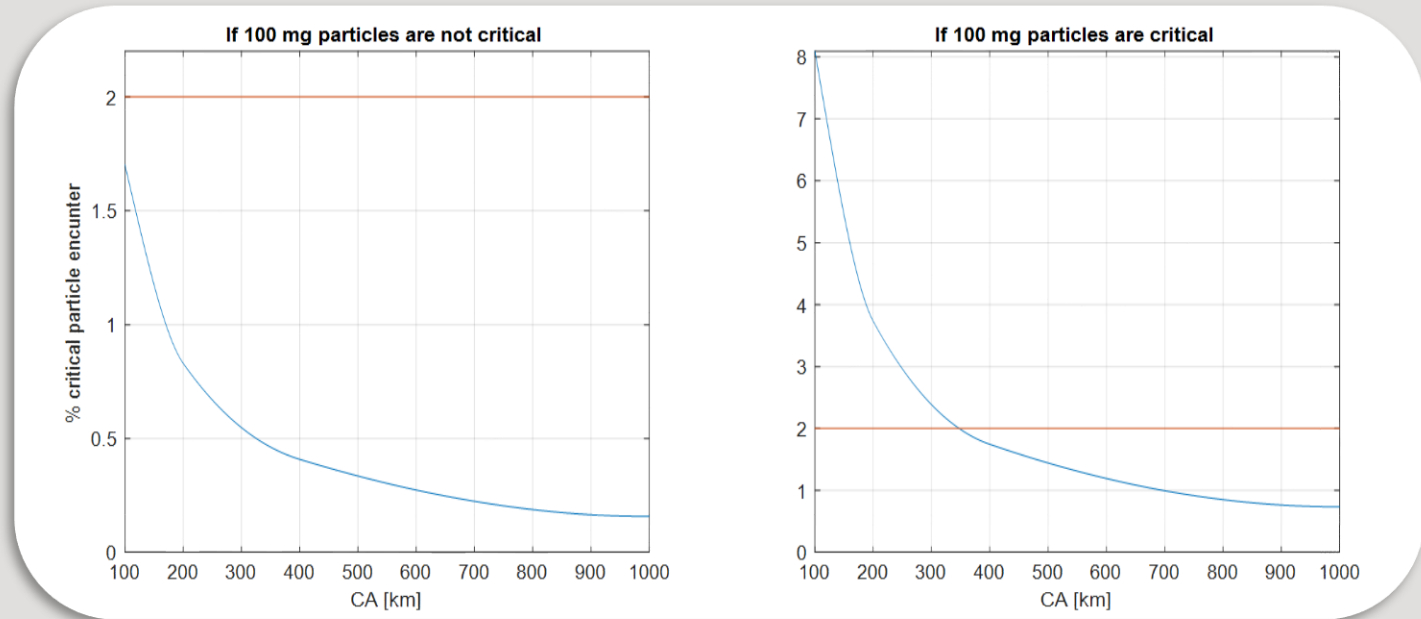
- Cometary environment simplified model made starting from Giotto and Rosetta observations.
- Probability of impact for particles of a given mass has been obtained for different CA distances.
- 100 mg particles are not critical up to flyby velocities of 64 km/s.
- For higher velocities, the CA distance shall be increased.



Double Whipple Shield with:

- 2.5 mm *Kevlar* wall
- 0.3 mm *aluminium* bumper
- 13 cm *spacing*

Designed through *ESABASE2*



DIPARTIMENTO DI
INGEGNERIA
INDUSTRIALE

POST-FLYBY PHASE

Data transmission time = 8 hours a day

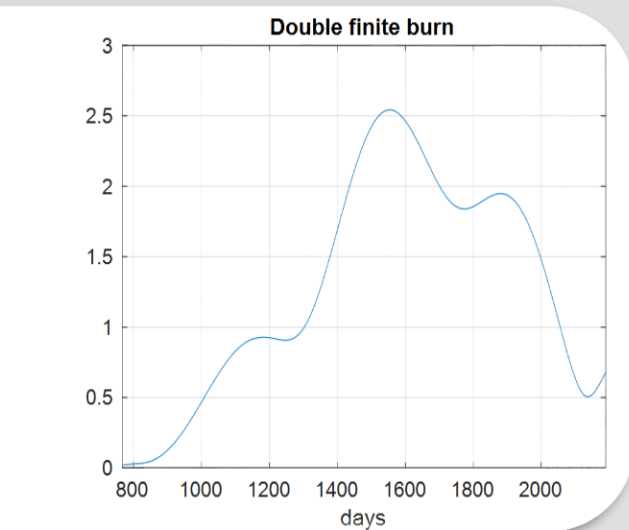
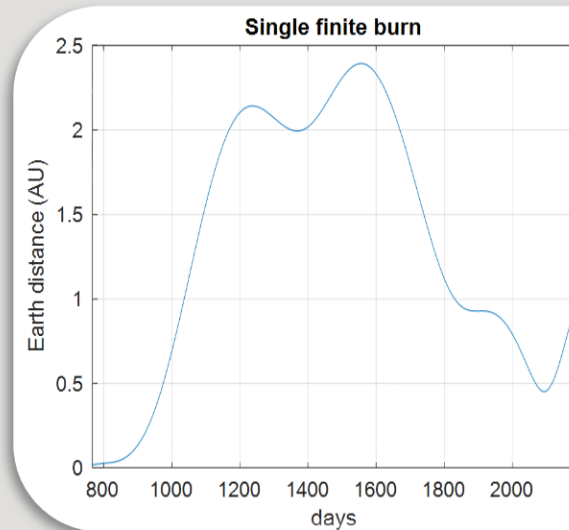
Duration of the phase depends on the duration of other phases

Simulation performed (worst case):

- Parking phase lasts 2.1 years
- Transfer phase lasts 2.6 years

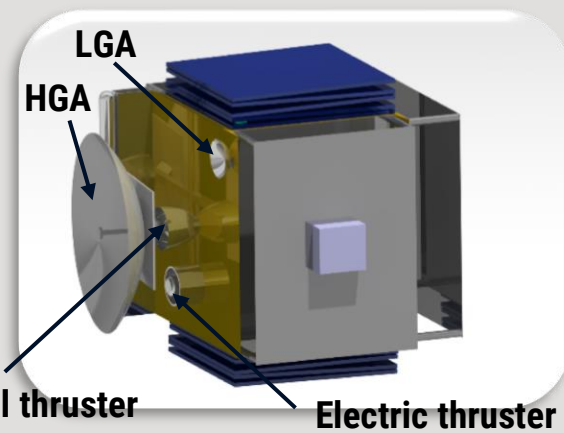
→ 1.3 years post-flyby duration

↓
30 GB data transferrable
(worst case)



Telecommunication characteristics:

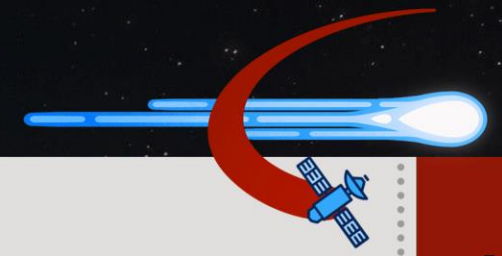
- 8.4 GHz X-band High Gain Antenna for data transmission
- 8.4 GHz X-band Low Gain Antenna for telemetry & command during the parking phase
- IRIS V2.1 X-band transponders for signal modulation
- Travelling Wave Tube Amplifier to amplify the power transmitted
- HGA mechanism to protect the antenna from cometary dusts and/or thruster plumes



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INGEGNERIA
INDUSTRIALE



SPACE SEGMENT DESCRIPTION

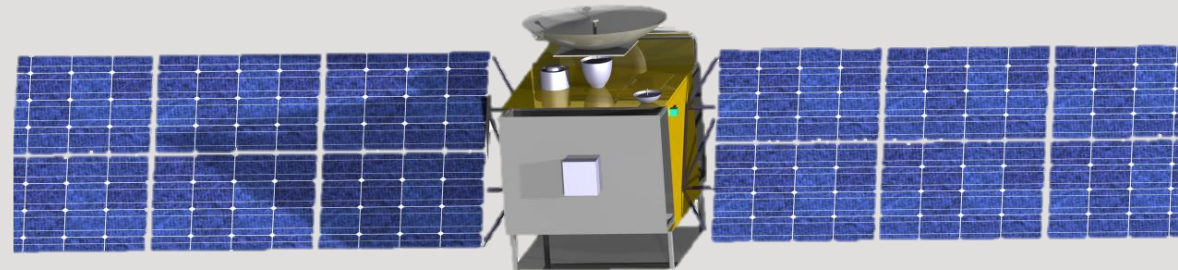


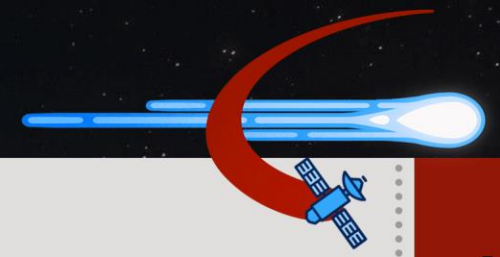
SPACE SEGMENT DESCRIPTION

- Different mission concepts selected through a **System Engineering** approach
- *Analytic Hierarchy Process* to perform *trade-offs* and select the final concept
- Mass & power margins taken into account according to *AIAA Regulations*
- Rough cost estimated of 35 M\$ (includes all mission operations costs)

Mass Budget [kg]		Power Budget [W]				
Structure	13.2		Parking	Transfer	Flyby	Comm.
Thermal Subsystem	0.810	No Margin	83	324	108	168
Attitude Control	6.86	Margin	45%			
Power Subsystem	21.5	Total	120	472	157	243
Propulsion	4.13	Link Budget				
Telecommunications	6.70	Distance from Earth [AU]		0.5	1	1.5
Payload	7.5	Data Rate [kbps]		66	16.7	7.4
Total Dry Mass (20% margin)	72.8	C/N ₀		53.7	47.7	44.2
Propellant (chemical electrical)	11.8 15.4	Volume (without shield) [m ³]		0.102		
Total	100.0	Dimensions (stowed) [cm]		86.0 x 78.0 x 77.2		

POWER SUBSYSTEM:
 Solar arrays (2.32 m²)
 Lithium-Ion secondary batteries
 STARBUCK Mini (PCDU)



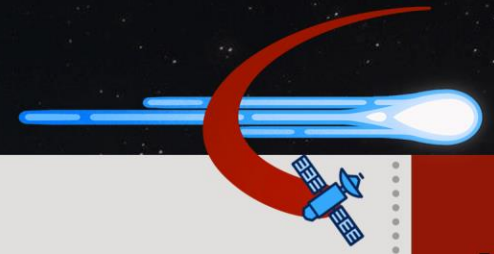


RISKS ANALYSIS

Impact					
5					
4		SAT-2 SAT-4			
3		SAT-3	SAT-1		
2			MA SAT-5 PAY COST		
1					
	1	2	3	4	5
					Likelihood

ID	Root Cause Description
PAY	Failure to gather data during the flyby
SAT-1	Failure due to environmental factors
SAT-2	Failure of satellite due to EPS malfunction
SAT-3	Failure in the telecommunication subsystem
SAT-4	Failure in the propulsion subsystem
SAT-5	Failure in the AOCS
MA	Wrong evaluation of orbits or manoeuvres
COST	Costs are higher than expected



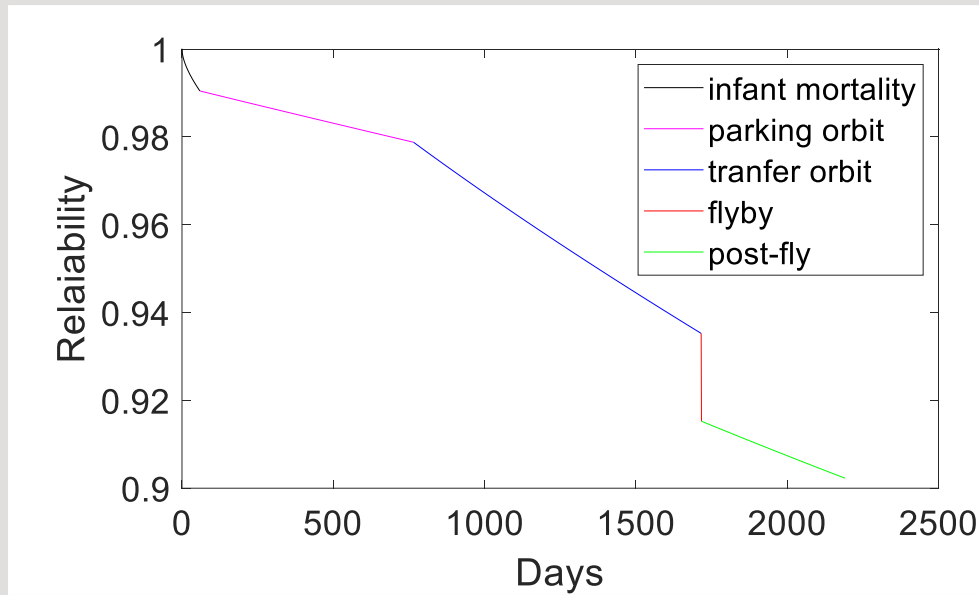


RELIABILITY ANALYSIS

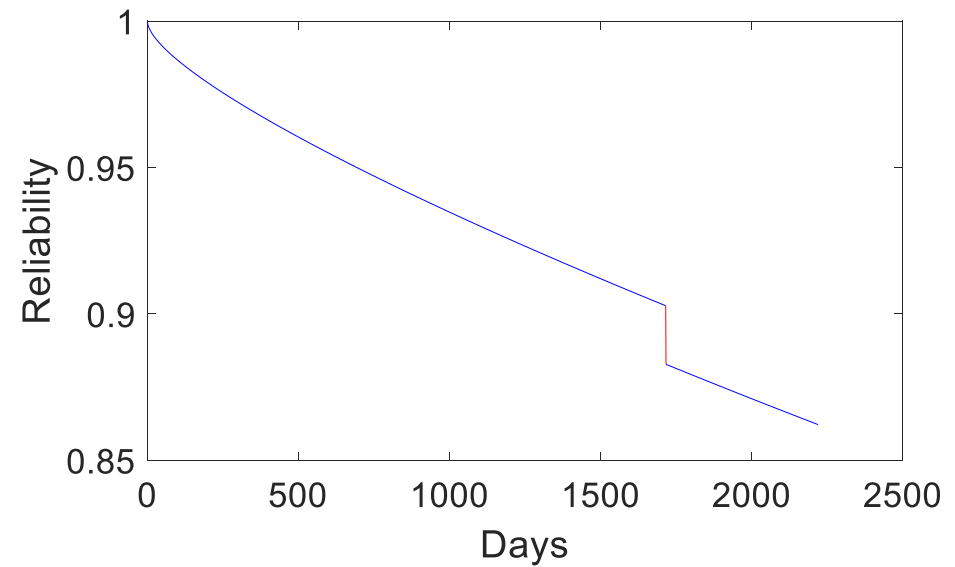
Twofold reliability analysis performed through *Weibull's distributions*:

- First analysis (a): specific mission scenario → only *most stressed* subsystem contributes to reliability in each phase.
- Second analysis (b): worst case scenario → *all the subsystems* are always considered on

$$0.86 < R < 0.90$$



(a)



(b)

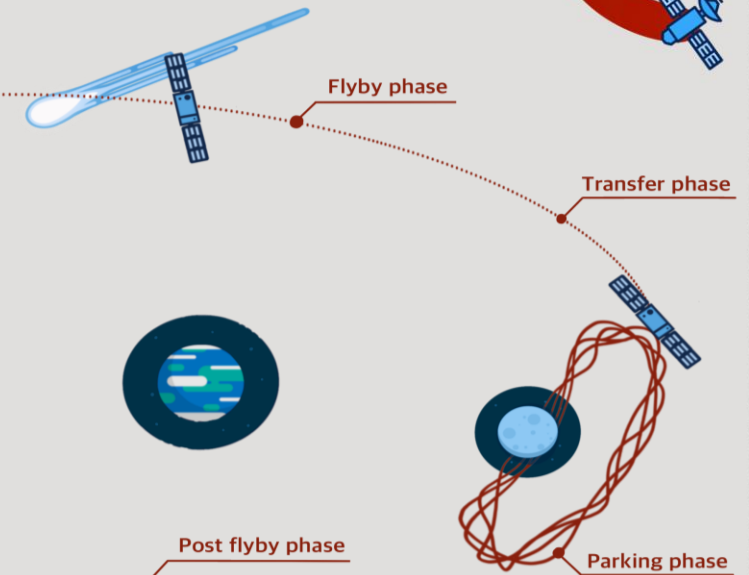


CONCLUSIONS

- Mission *feasible* in terms of risk assessment
- Mission *feasible* in terms of *mass and size constraints*
- Study *demonstrates* the feasibility of future scientific interplanetary missions with micro-nano satellites
- Flexibility of mission phases due of the probabilistic nature of the problem

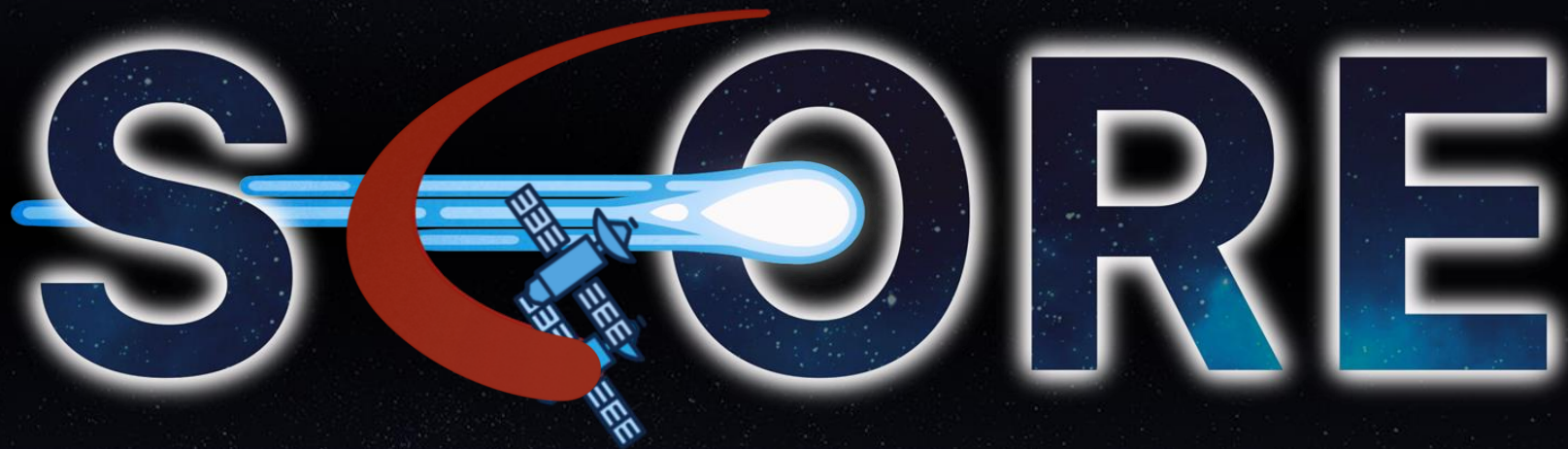
Critical Aspects of the mission:

- Cometary environment: need of a more accurate model
- Development of the electric thruster (TRL 5): development carefully monitored
- The **target will be known** once the satellite is on orbit → the specific mission scenario will be determined during the mission initial phases.



HT400 Electric Thruster

SCORE

The word "SCORE" is rendered in a bold, blue, sans-serif font with a white glow. The letter 'C' is replaced by a red swoosh. The letter 'O' is replaced by a blue rocket ship with a white flame, and a satellite is positioned behind it. The background is a dark blue space with white stars.

THANK YOU FOR YOUR ATTENTION!