



Gakugei 東京学芸大学  
Tokyo Gakugei University



# Microsatellites observing atmospheric and space electricity for the science of serious natural disasters: Challenge to their mitigations

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# Introduction

## Earthquake



**Mission 1:**  
Verification of earthquake ionospheric precursor for practical earthquake prediction



This photo is taken by Leggi il Firenzepost

## Tsunami



**Mission 2:**  
Investigation of tsunami ionospheric hole for early warning system



## Lightning



**Mission 3:**  
Study of lightning-related phenomena for lightning prediction.



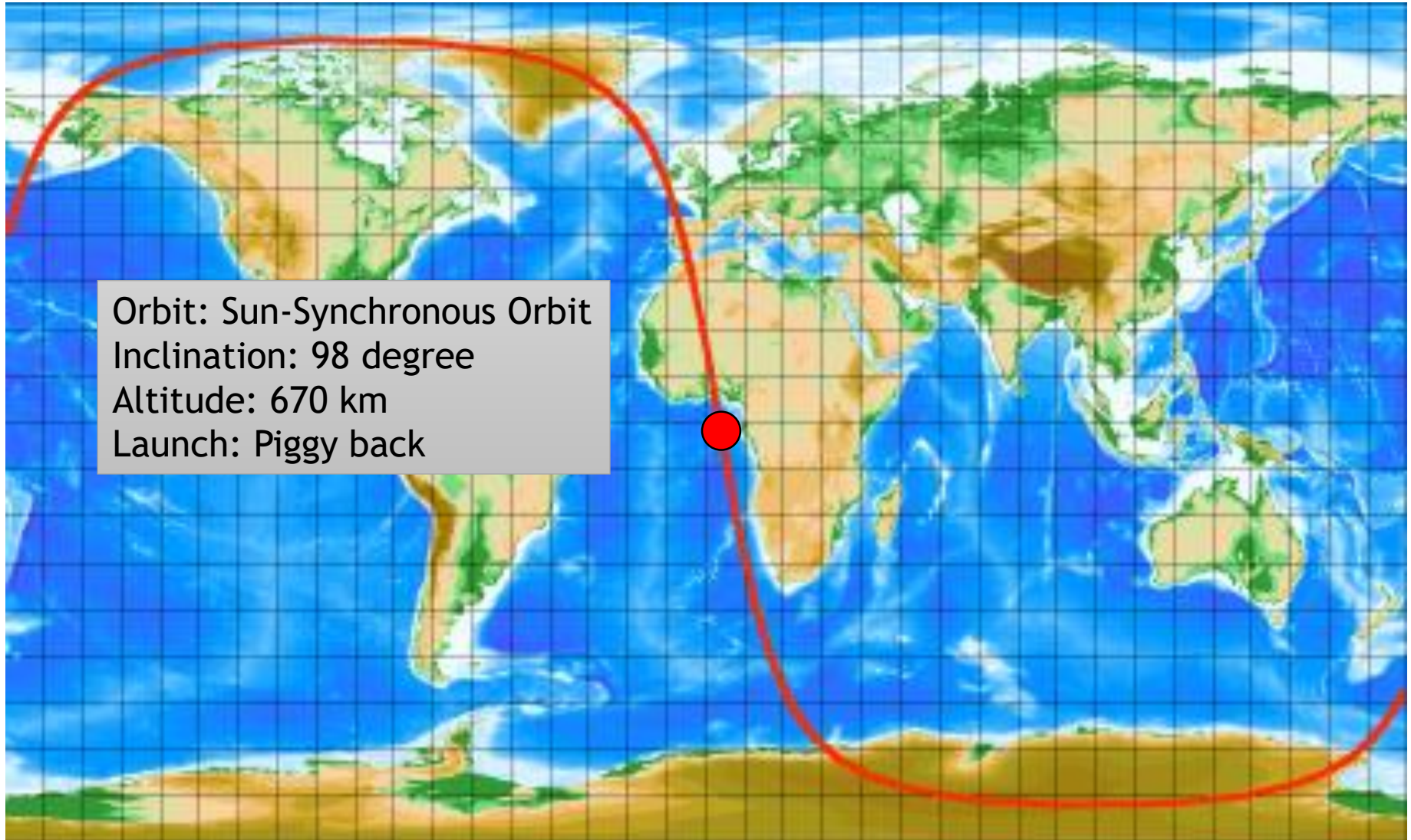
## Global warming



**Mission 4:**  
Study of global lightning for global warming understanding



# Plan A: Single satellite operation (Low cost plan)



# Plan B: 3 satellites constellation (Integrated plan)

## 3 satellites

Sun-Synchronous Orbit

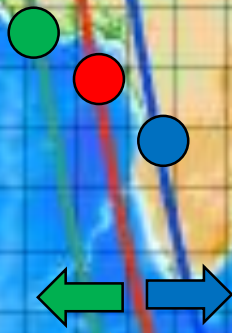
Inclination:

98, 97.8, 98.2 deg.

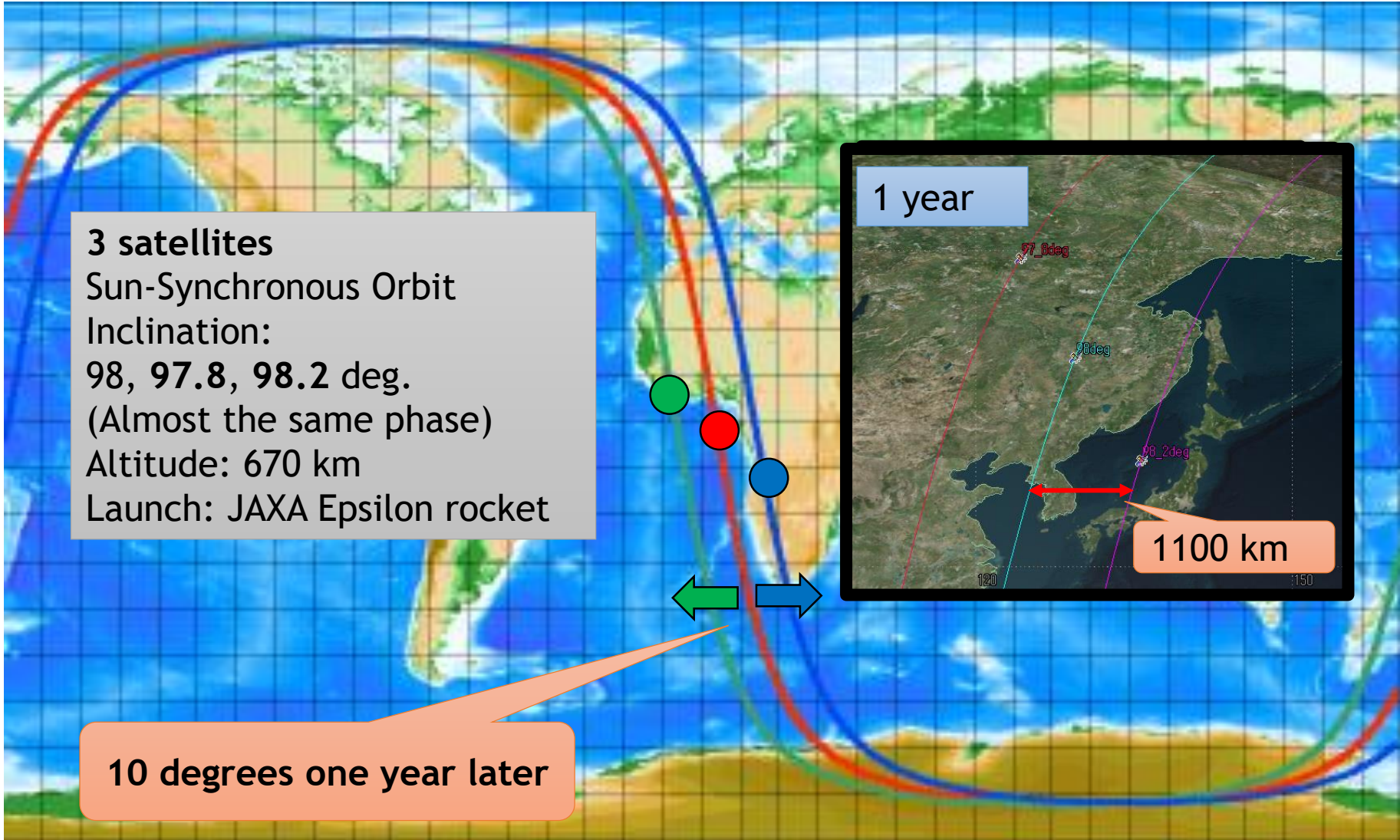
(Almost the same phase)

Altitude: 670 km

Launch: JAXA Epsilon rocket

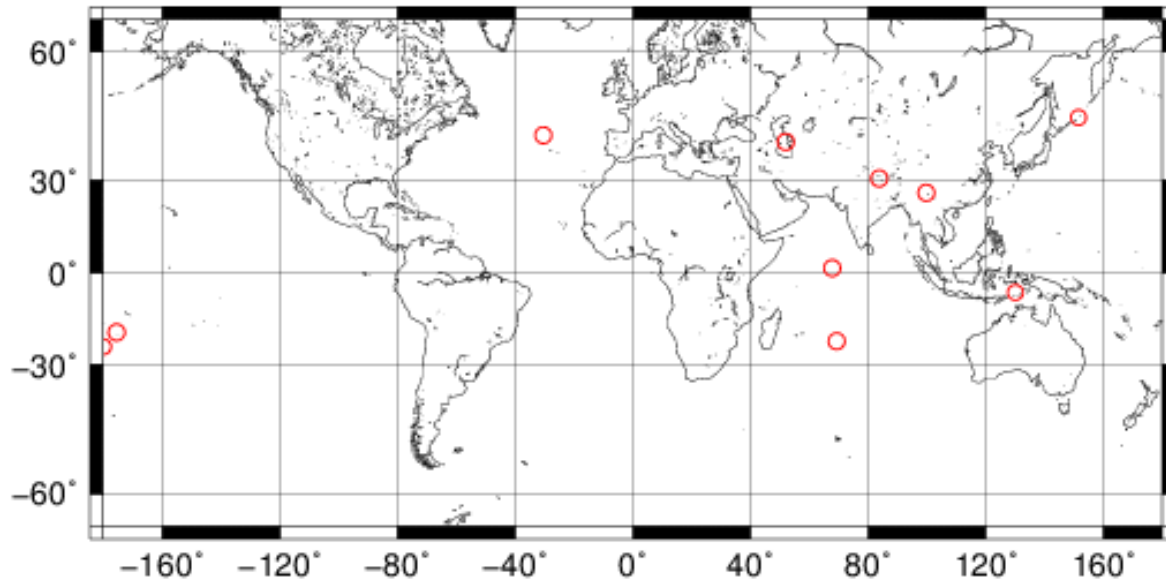


10 degrees one year later



# Mission 1 :

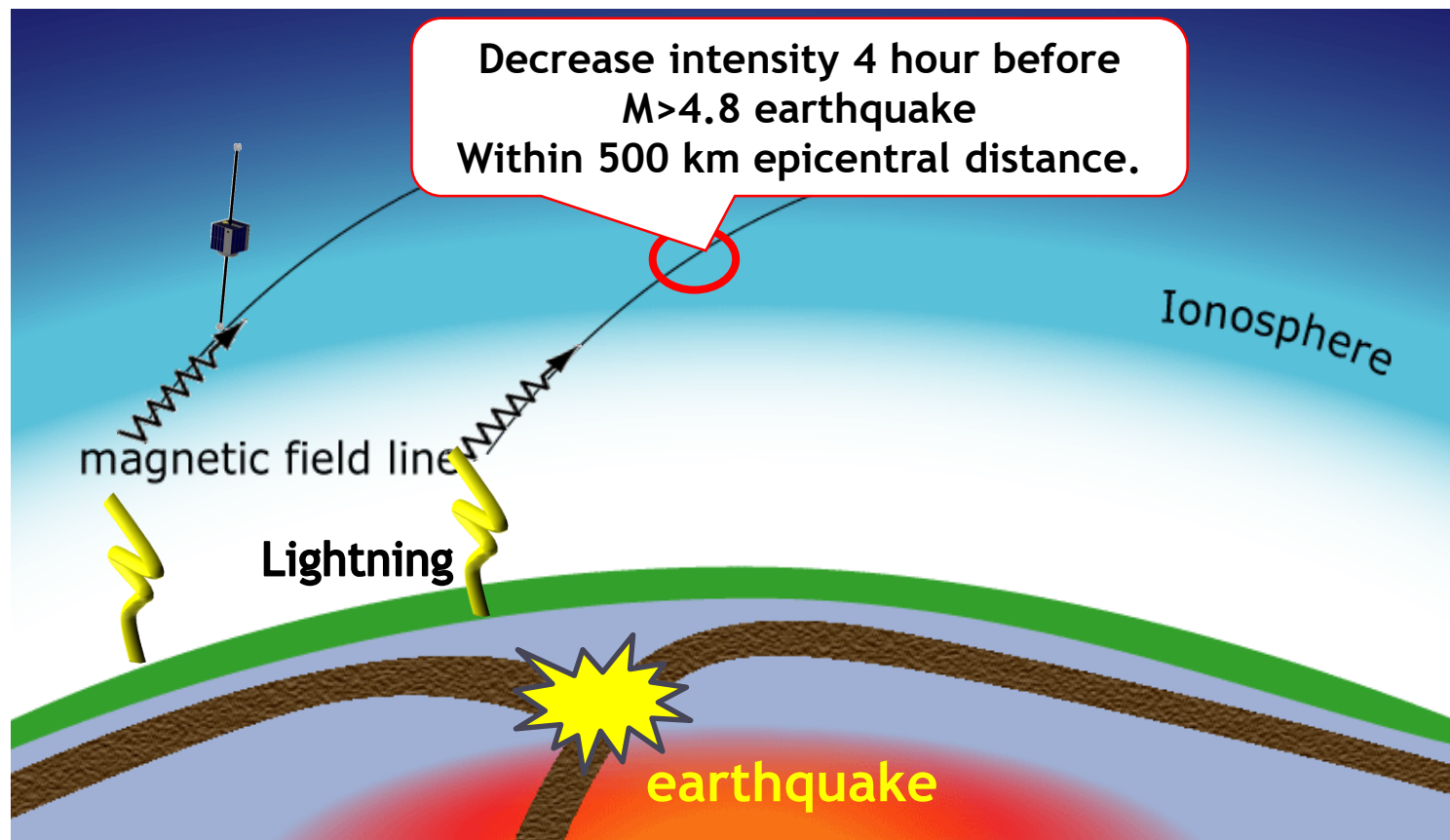
## Verification of earthquake ionospheric precursor for practical earthquake prediction (EQ mission)



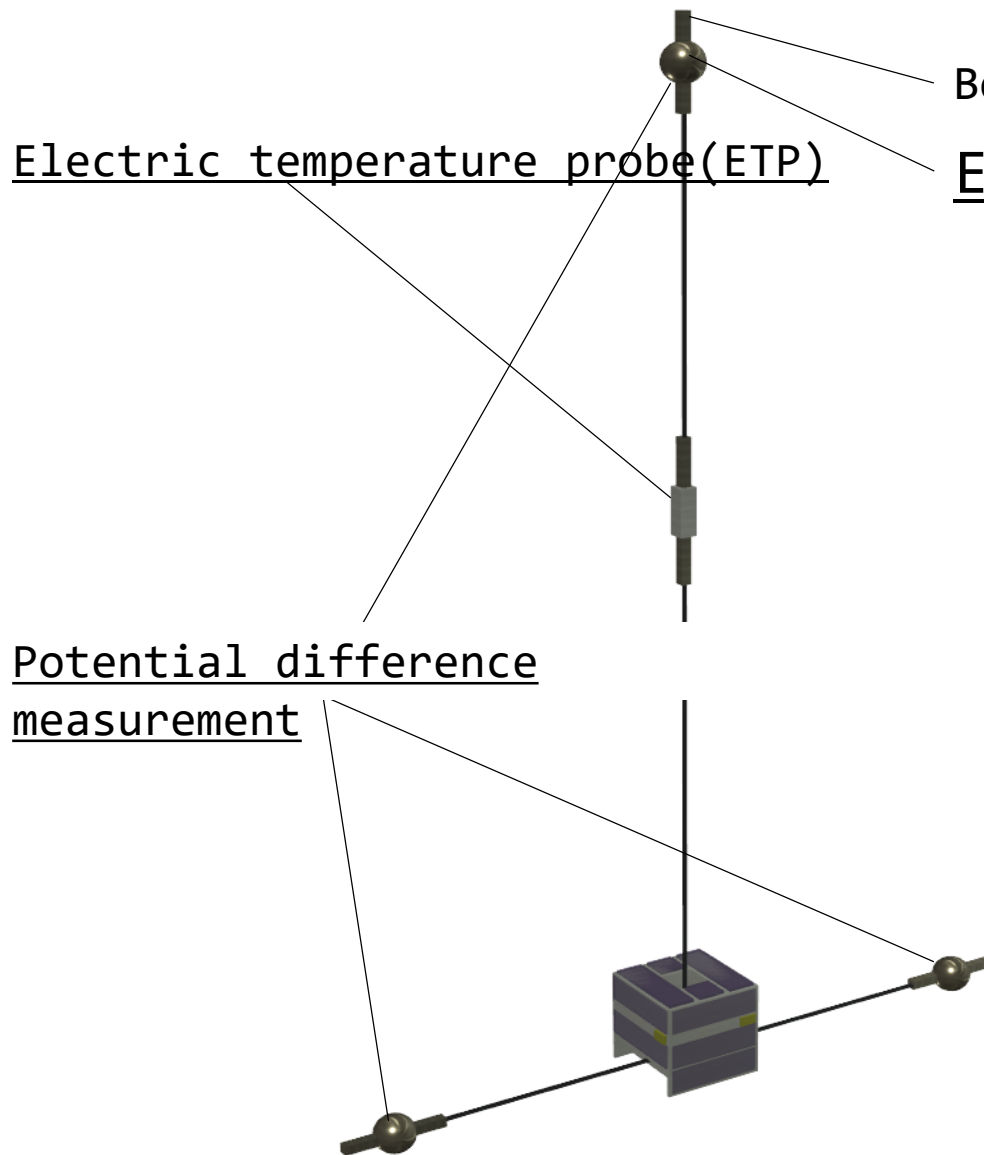
**No reliable and frequently repeatable precursor for a long time.**

# Pre-seismic decrease of VLF wave intensity at nighttime

Nemec et al., Geophys. Res. Lett. (2008)  
Kamogawa et al. (in preparation)



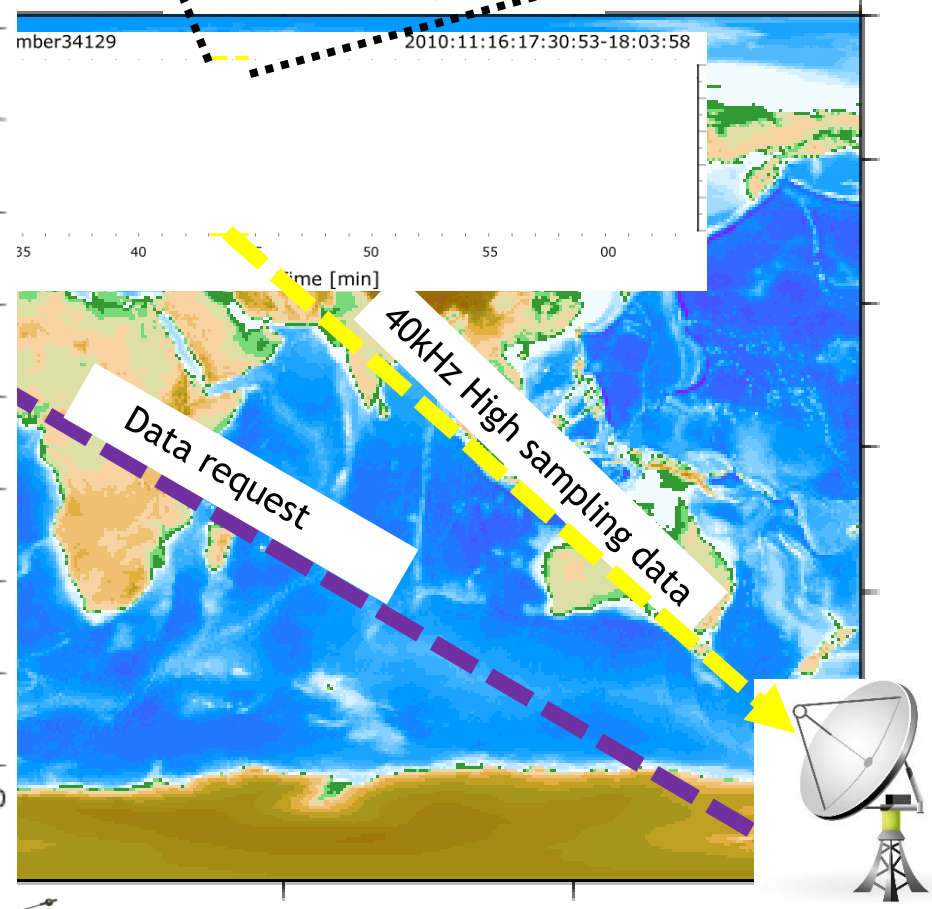
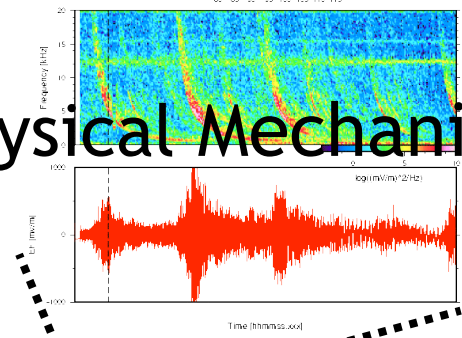
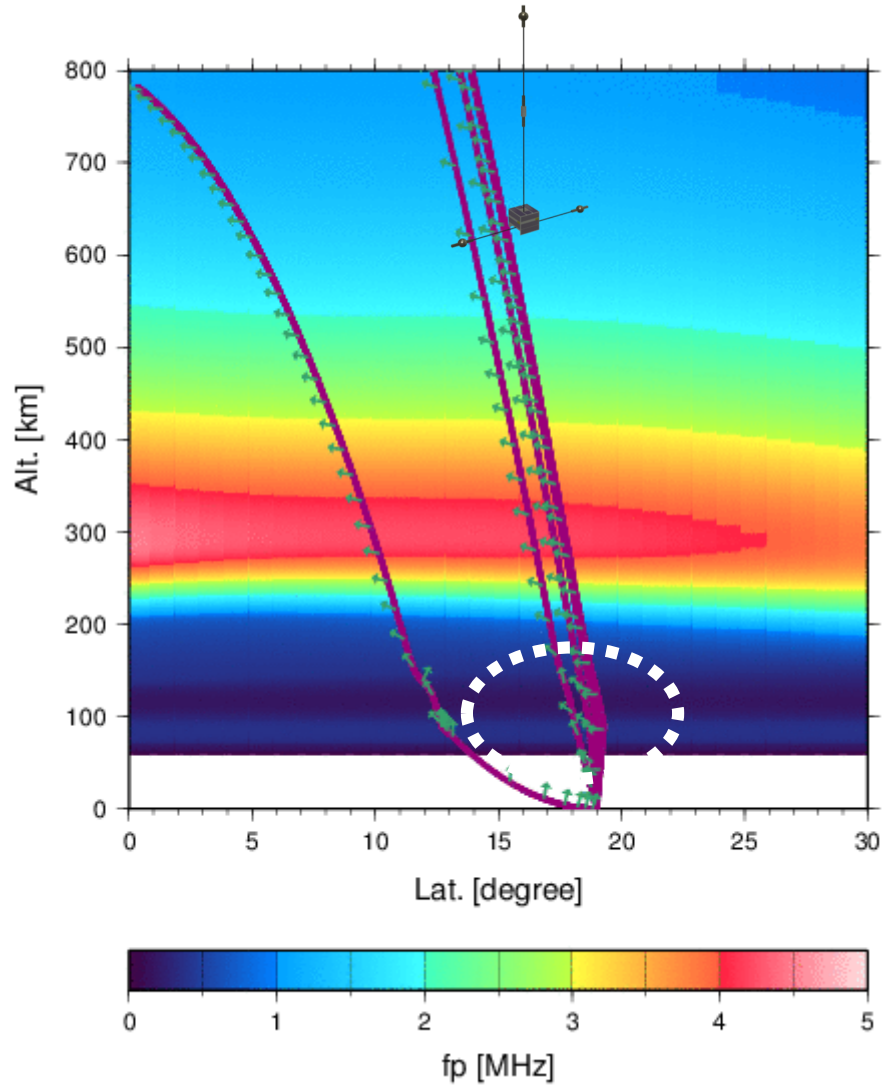
# Mission system



	Attitude control accuracy	Attitude determination accuracy
DC	1 deg.	7 deg.
AC	5 deg.	37 deg.

# Plan A: Understanding of Physical Mechanism

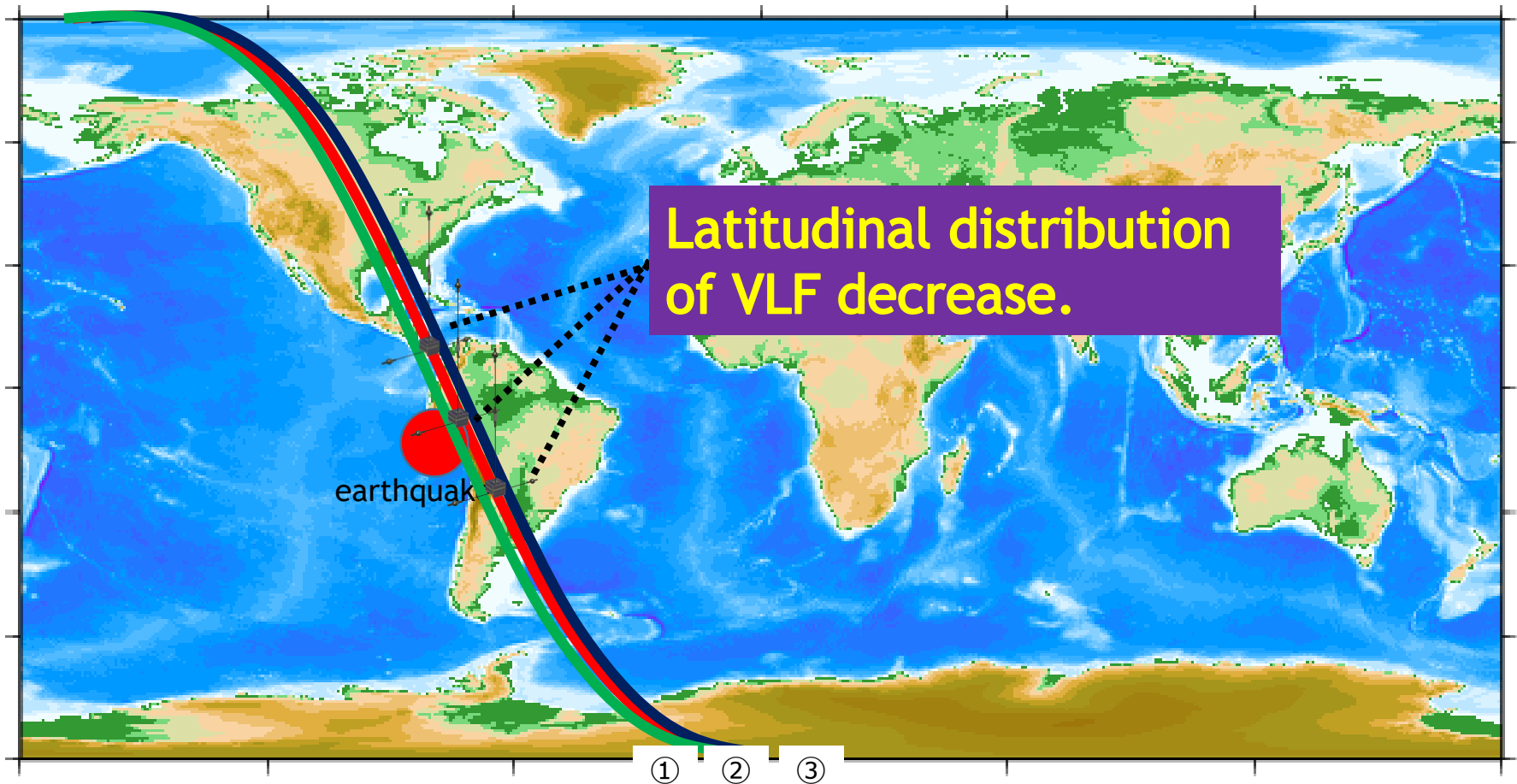
Ray path for the dipole field model 20060810UT4.5





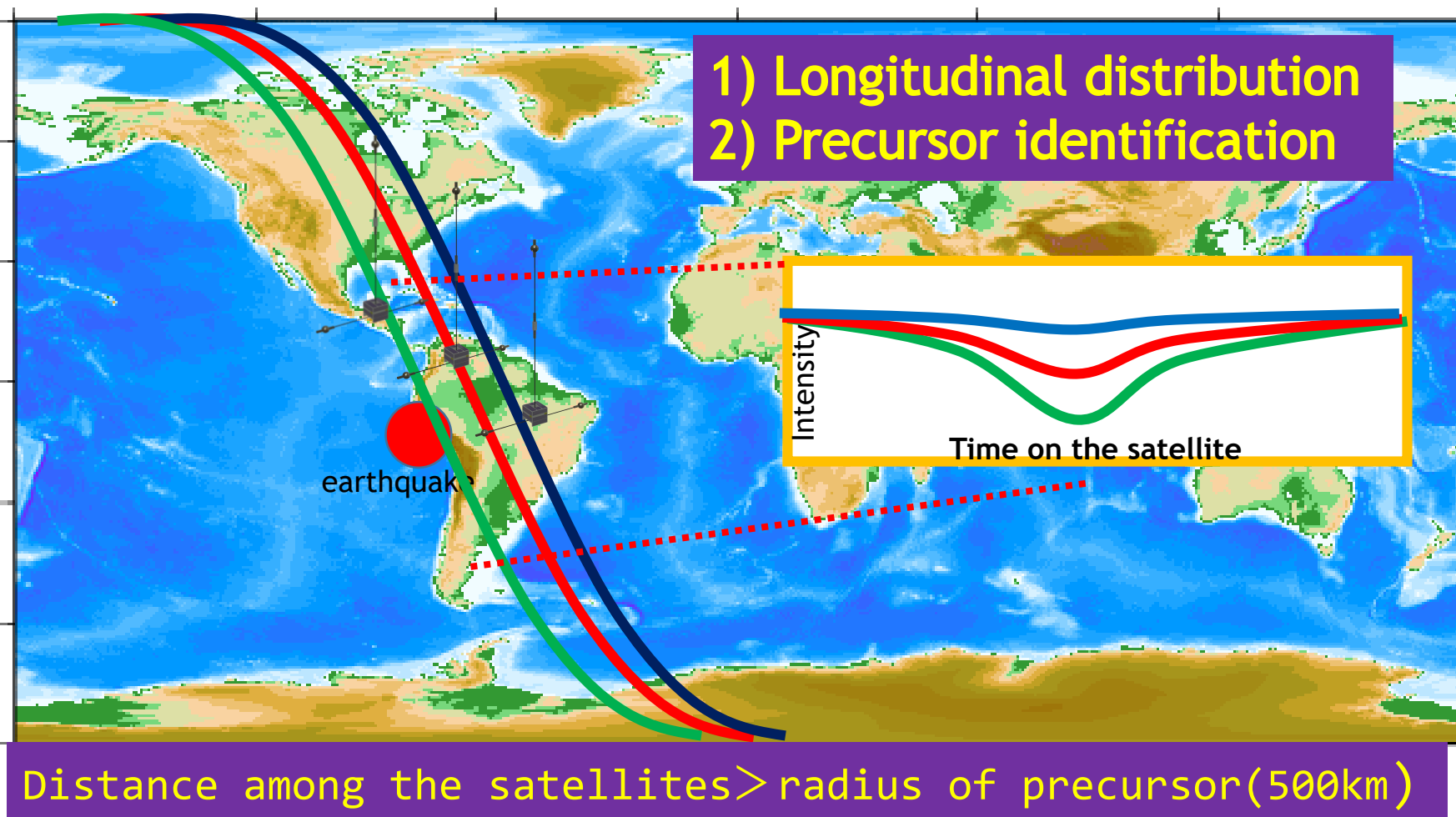
# Plan B

## Early phase: Investigation of Latitudinal Distribution



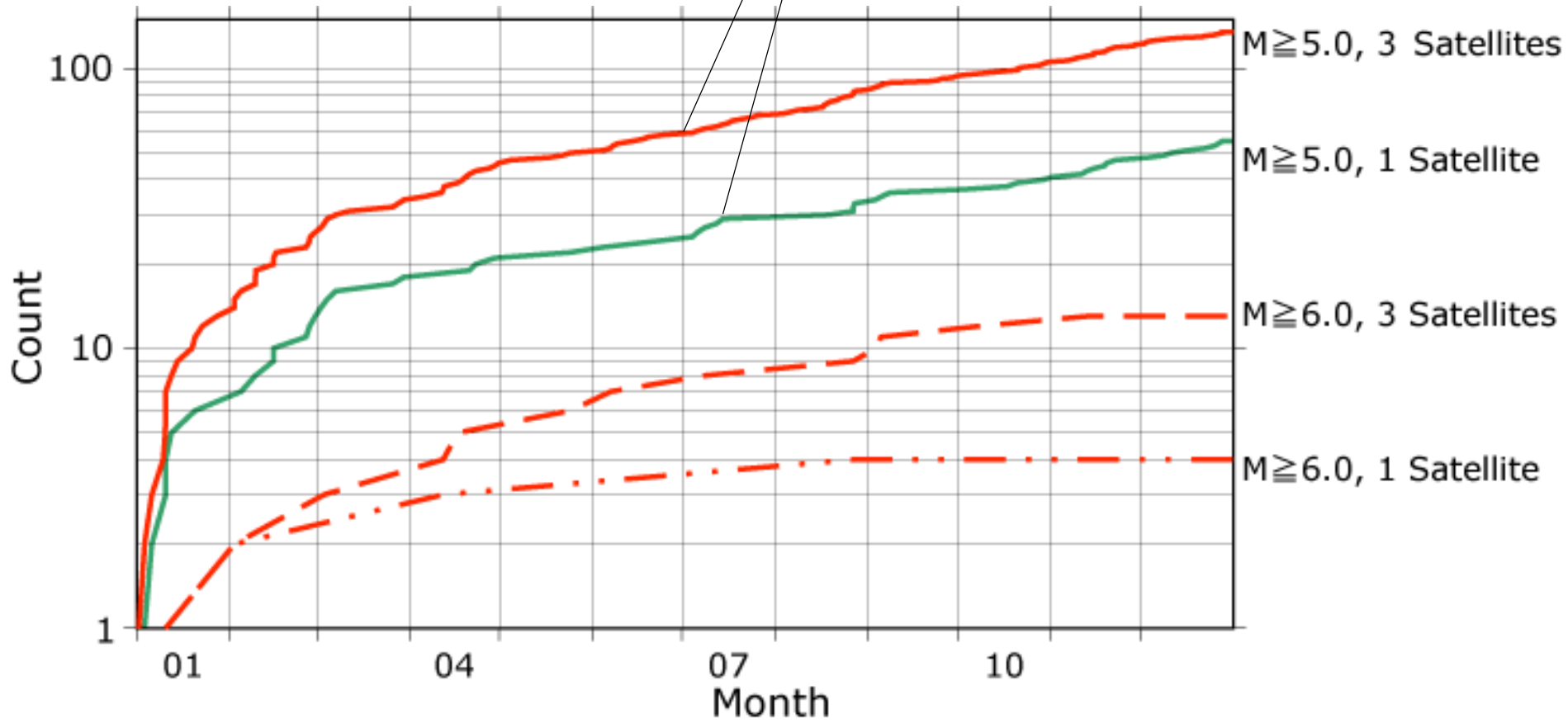
## Plan B

Late phase : Feasibility of Real-time Identification  
: Robust Verification of Phenomenon



Doubled events

## Expected Number of Earthquakes

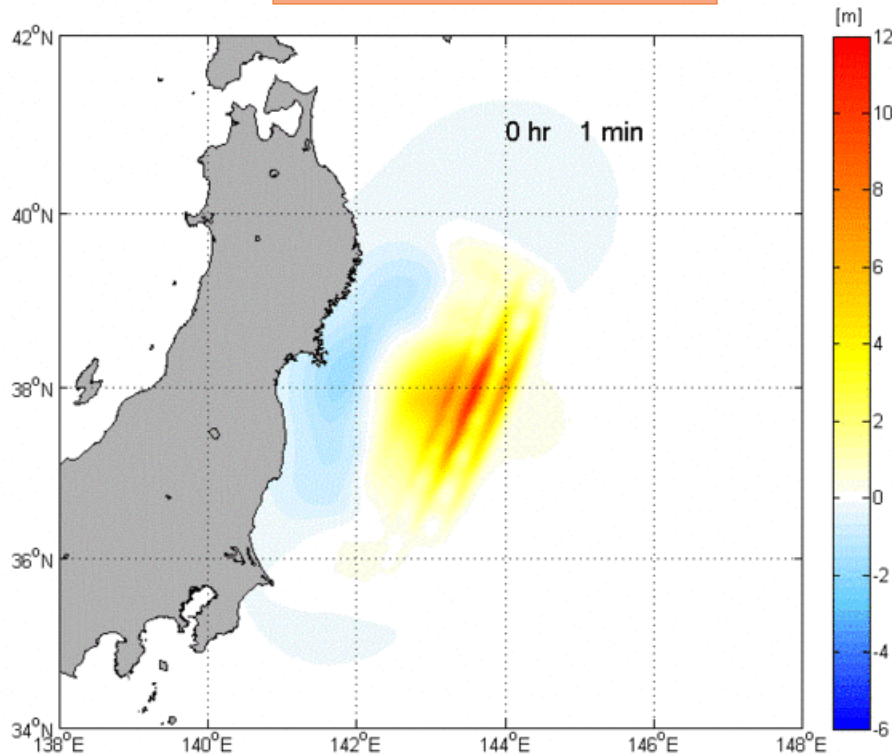


## Mission 2:

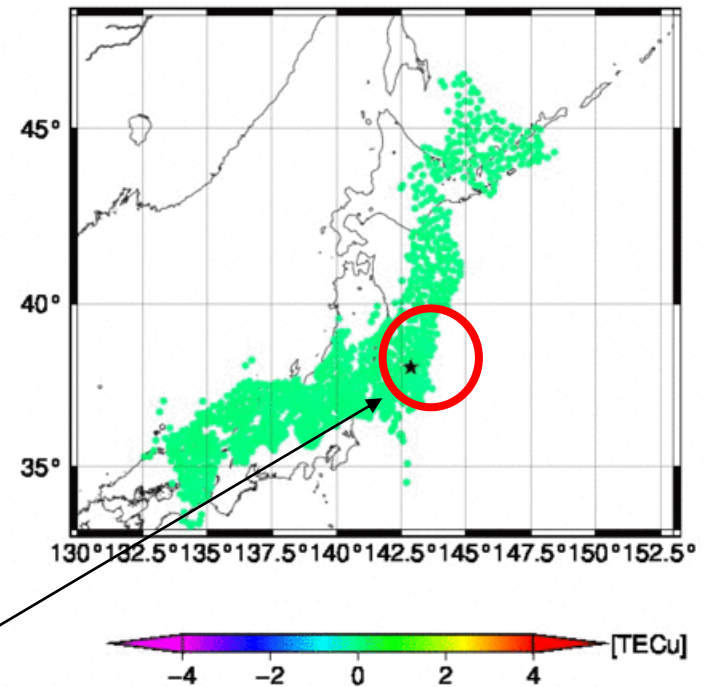
## Investigation of tsunami ionospheric hole for early warning system (Tsunami mission)

2011.3.11 Tohoku Earthquake (Japan)

## Height of Tsunami

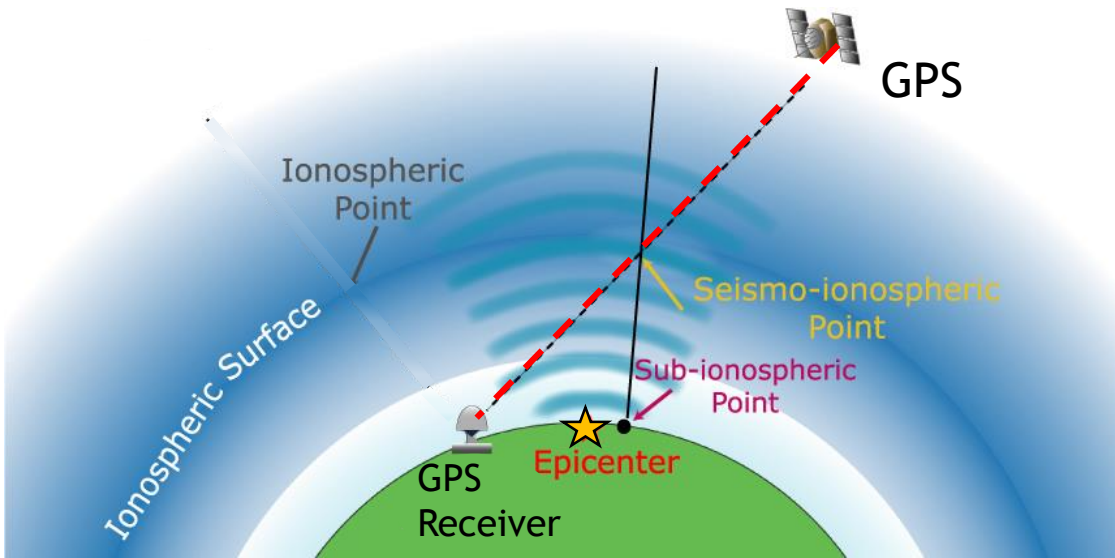


## Total Electron Content (TEC) in Ionosphere



We discovered **Tsunami Ionospheric Hole (TIH)**

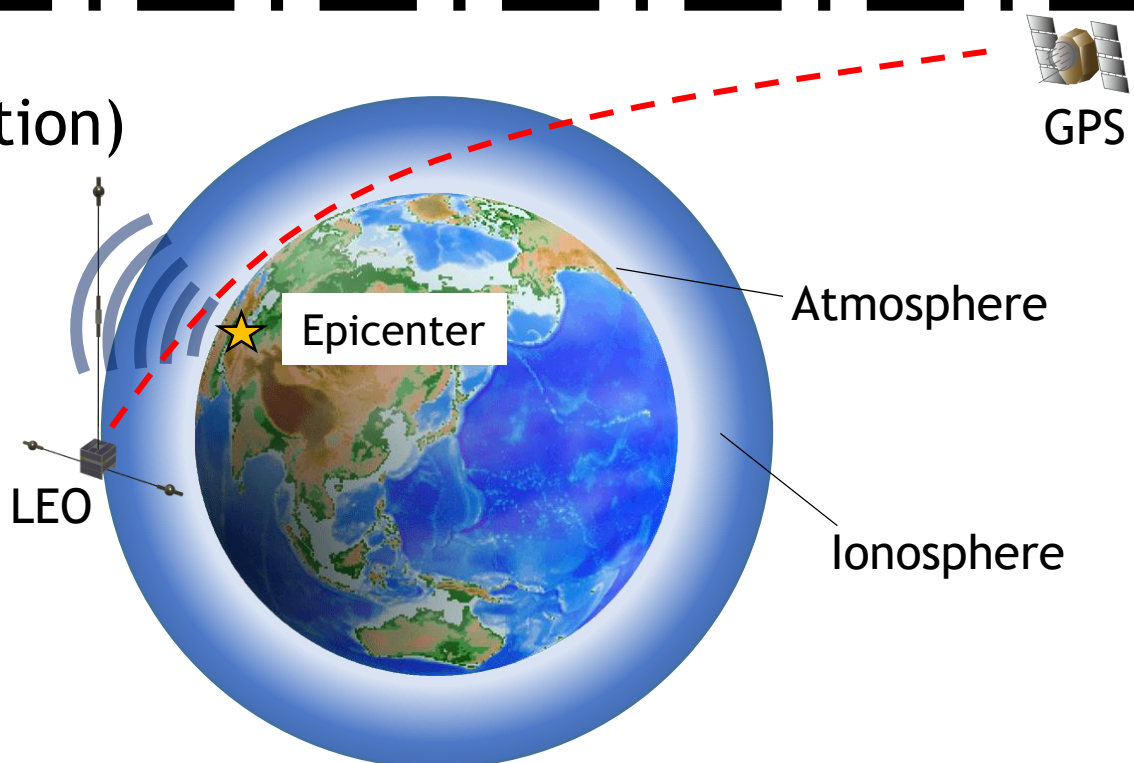
# Ground observation



GPS stations are not uniformly distributed

GPS stations are not far from the coast

# Satellite observation (GPS occultation observation)



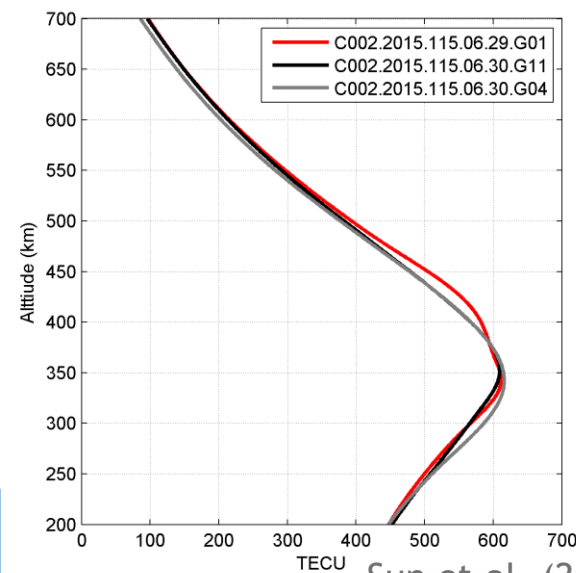
Satellite observation can monitor the whole area in the earth

# Observation

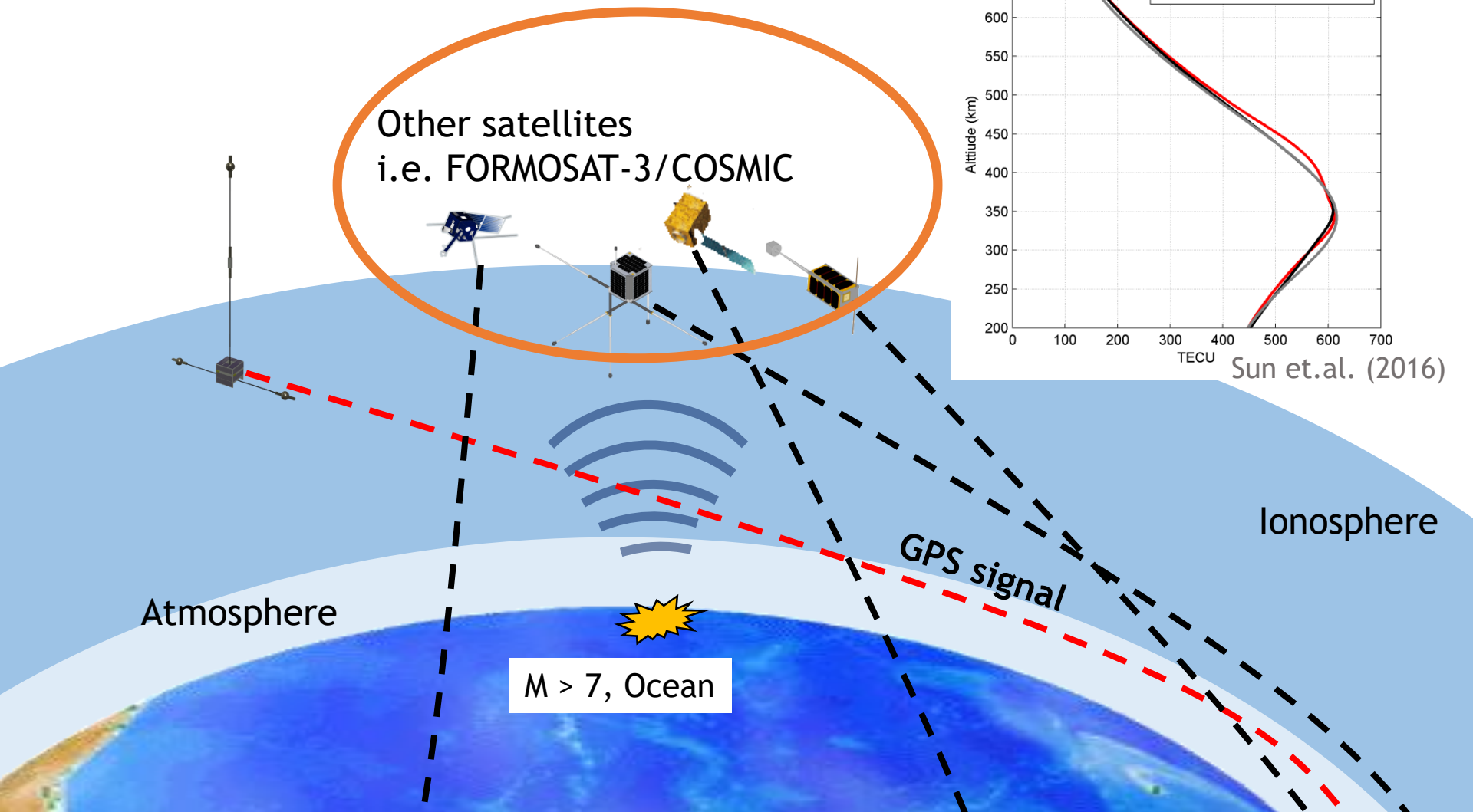
GPS receiver      Low power consumption  
                         Low cost  
                         Small

Other satellites  
i.e. FORMOSAT-3/COSMIC

### Electron Density Profile

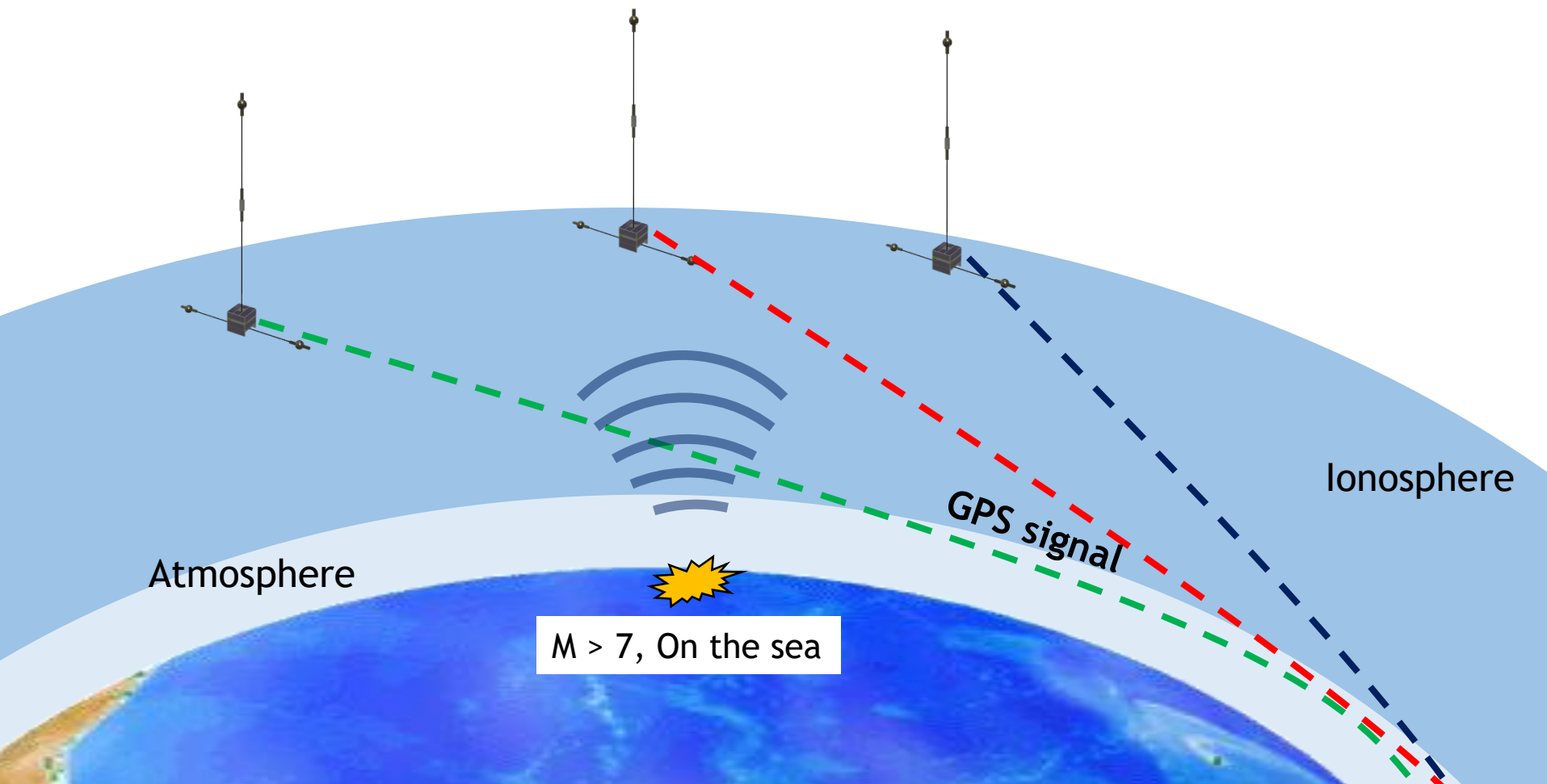


Sun et.al. (2016)



# Plan B: Constellation Observation

Verification feasibility for satellite-based tsunami early warning system from real time monitoring



# Mission 3: Study of lightning-related phenomena for lightning prediction. (Lightning mission)

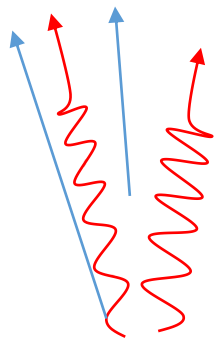


Photo given by Uchinada town Local government

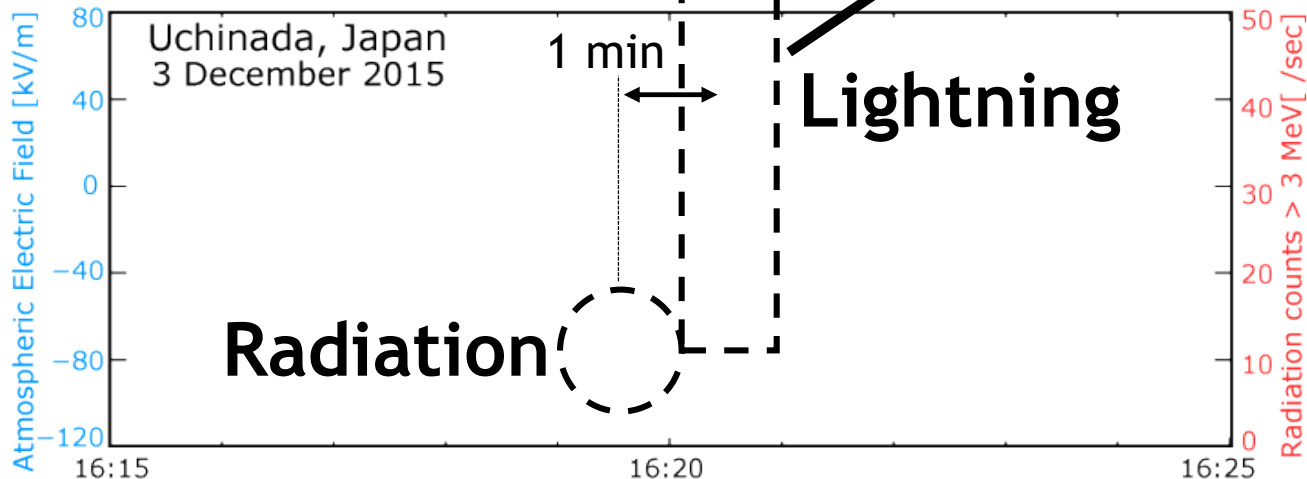
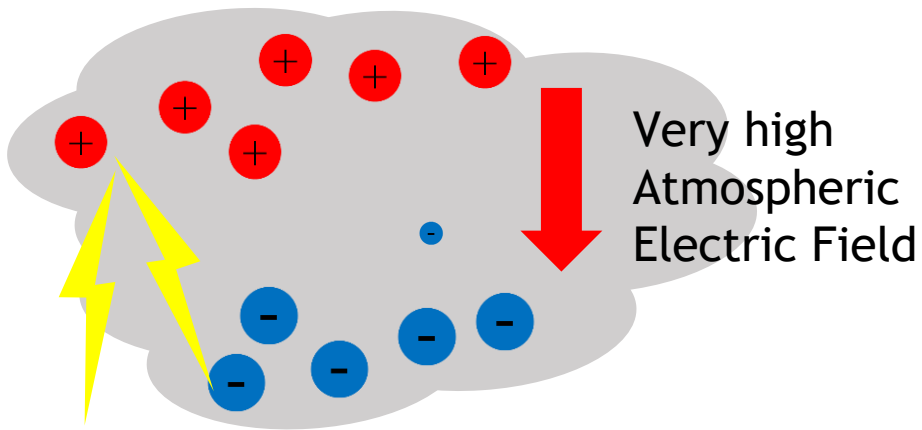
15 09/01 17:17:01.764 T. NO.0



# Gamma Radiation and Thundercloud

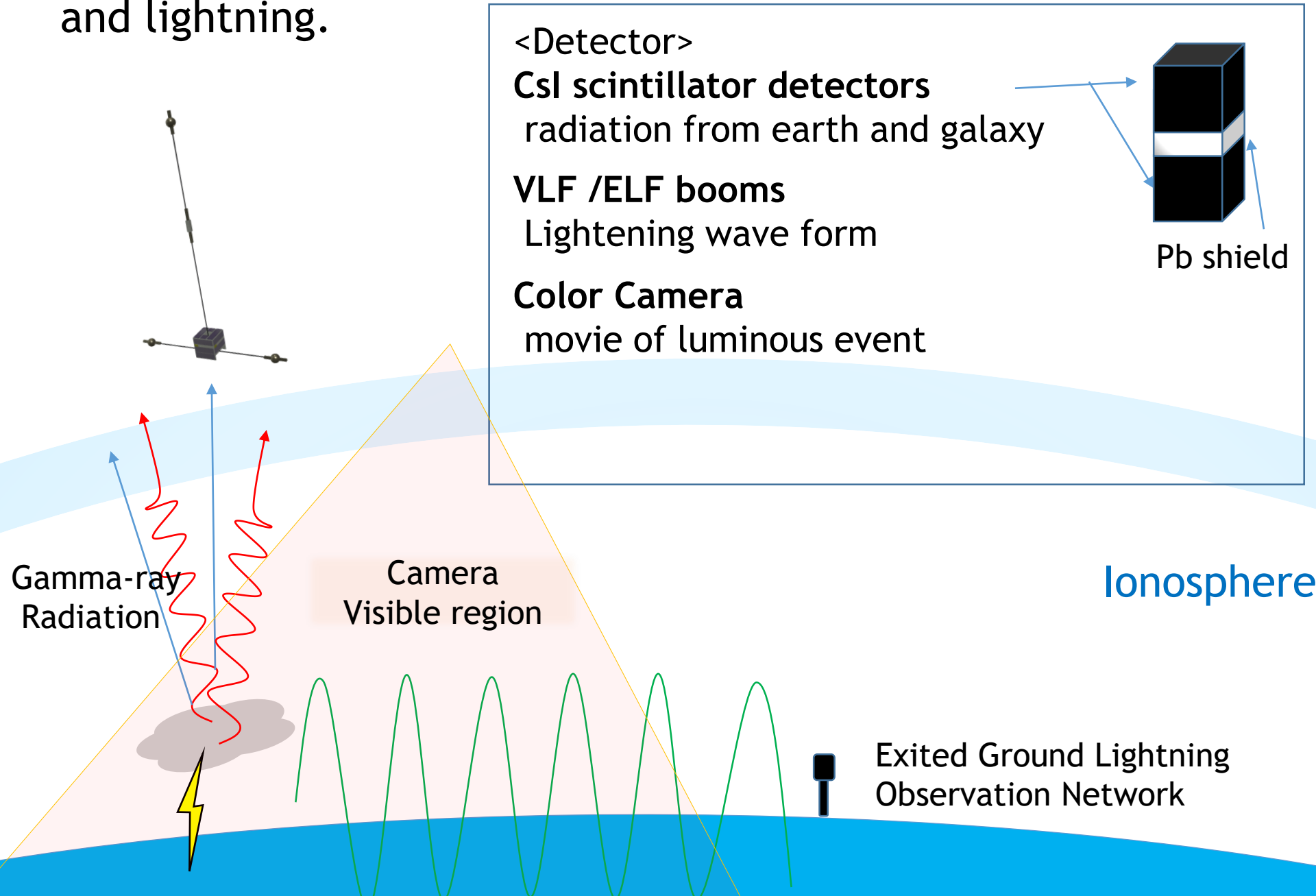


Gamma Radiation



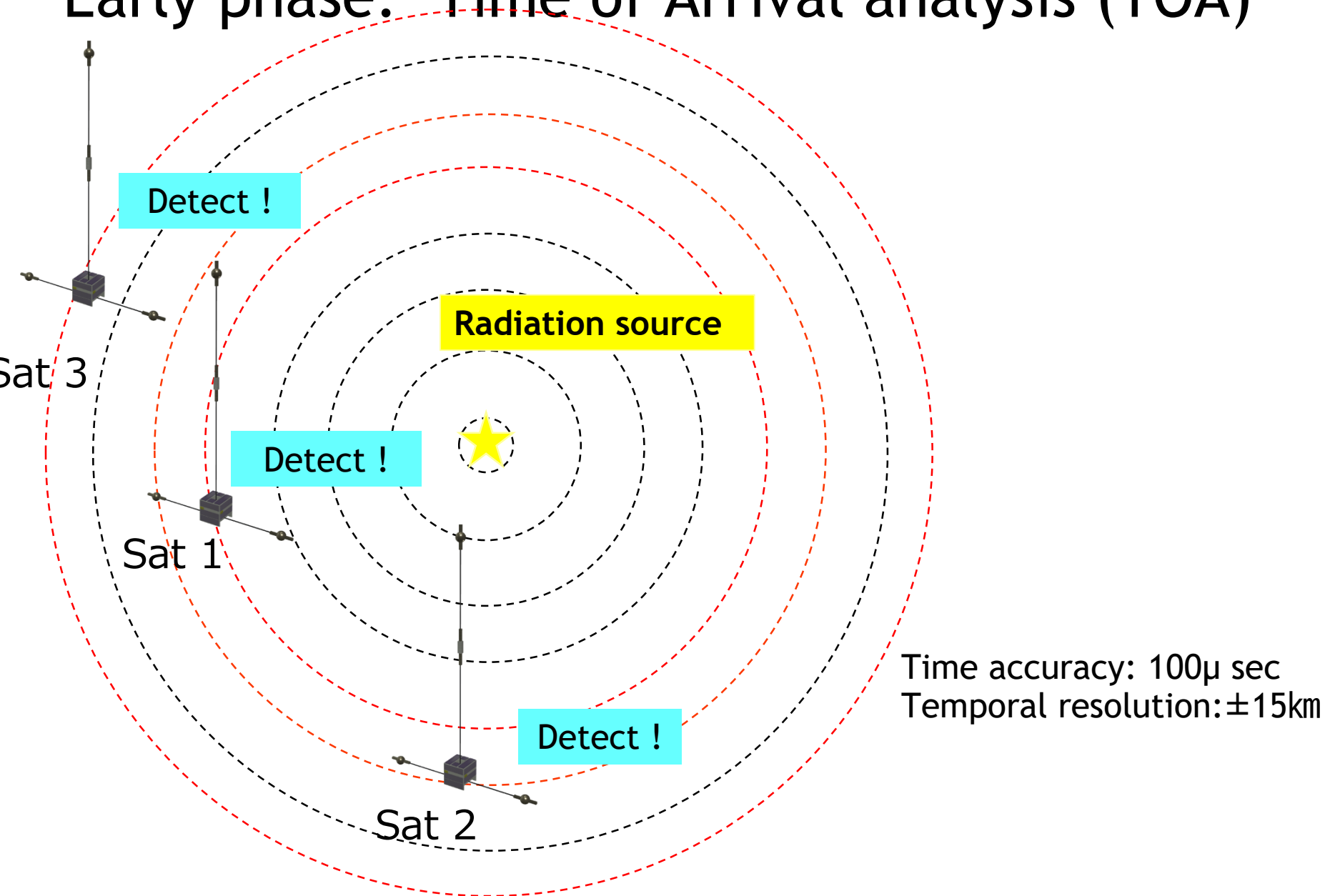
We will find new lightning prediction technique

# Plan A: Verifying statistical correlation between gamma ray and lightning.



# Plan B

## Early phase: Time of Arrival analysis (TOA)



# Mission 4

## Study of global lightning for global warming understanding (Global warming mission)



After JCCCA

CO<sub>2</sub> gas

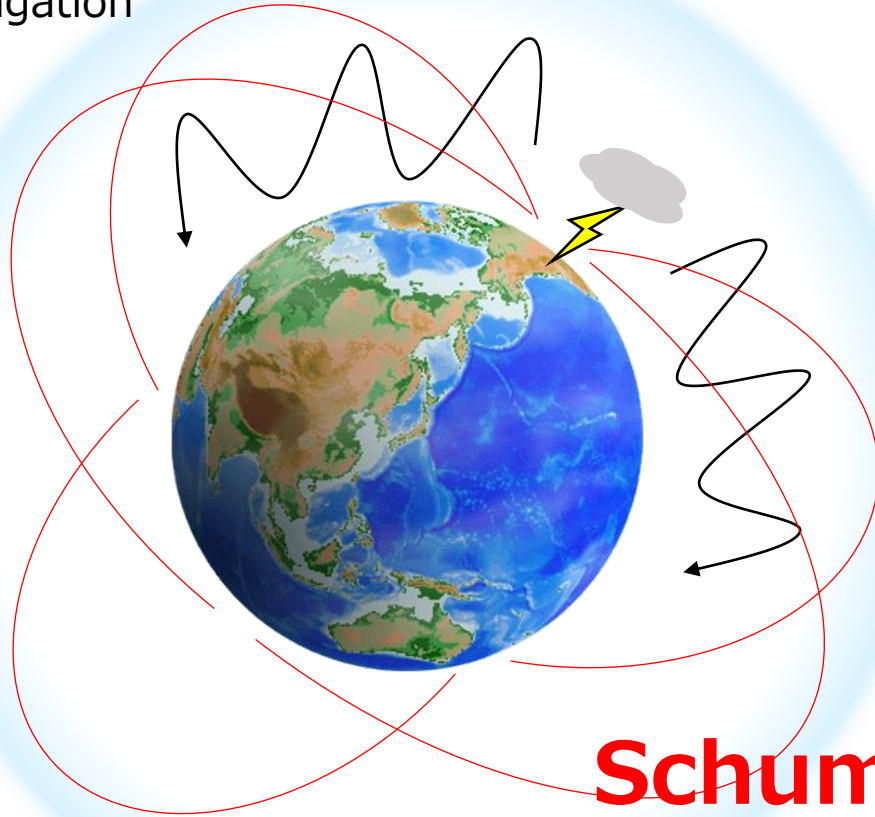
Atmospheric  
Electric Field

Cosmic-ray

Multiple factors?

# Global Electric Circuit (GEC)

ELF/VLF waveguide propagation



**Temperature**



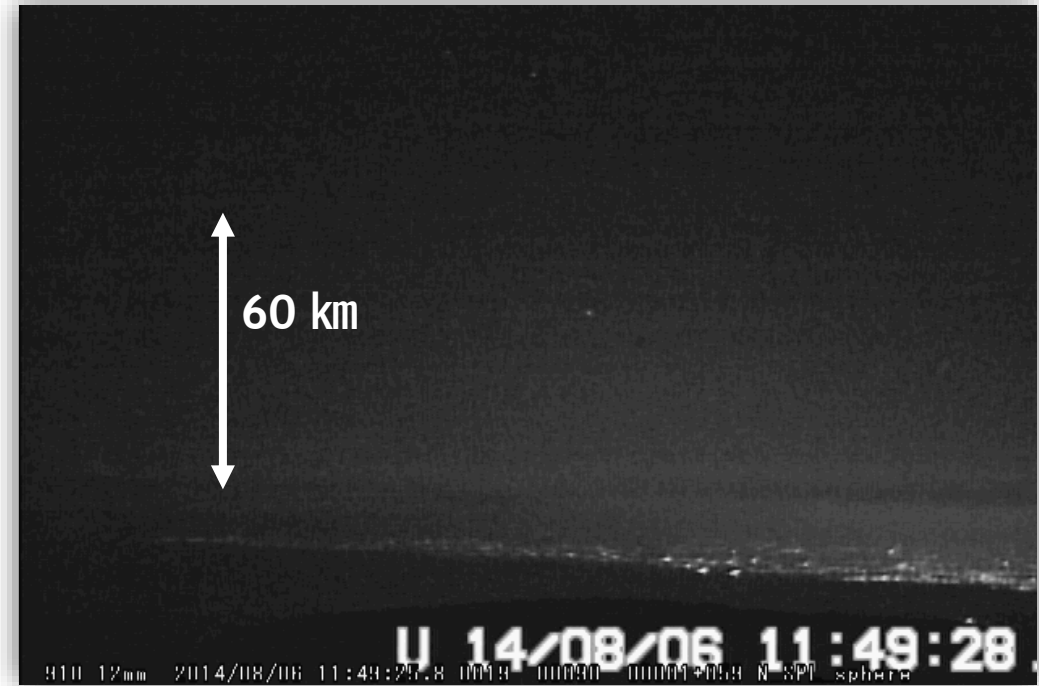
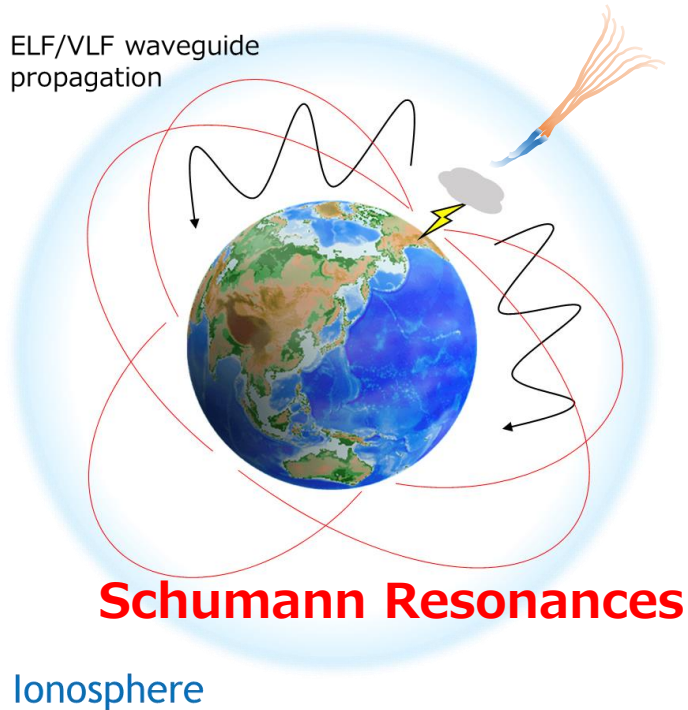
High correlation  
(Williams, Science, 1992)

**Schumann Resonances**

Ionosphere

# Transient luminous events (TLEs)

## Gigantic-Jet at summit of Mt. Fuji, Japan



TLE emit TLEs emit 10 times stronger ELF wave than Schumann resonances

➔ We investigate how much TLE event influence on the Schumann resonance intensity for global warming study.

# Plan A : Schumann resonance and TLE observation

<Detector and object >

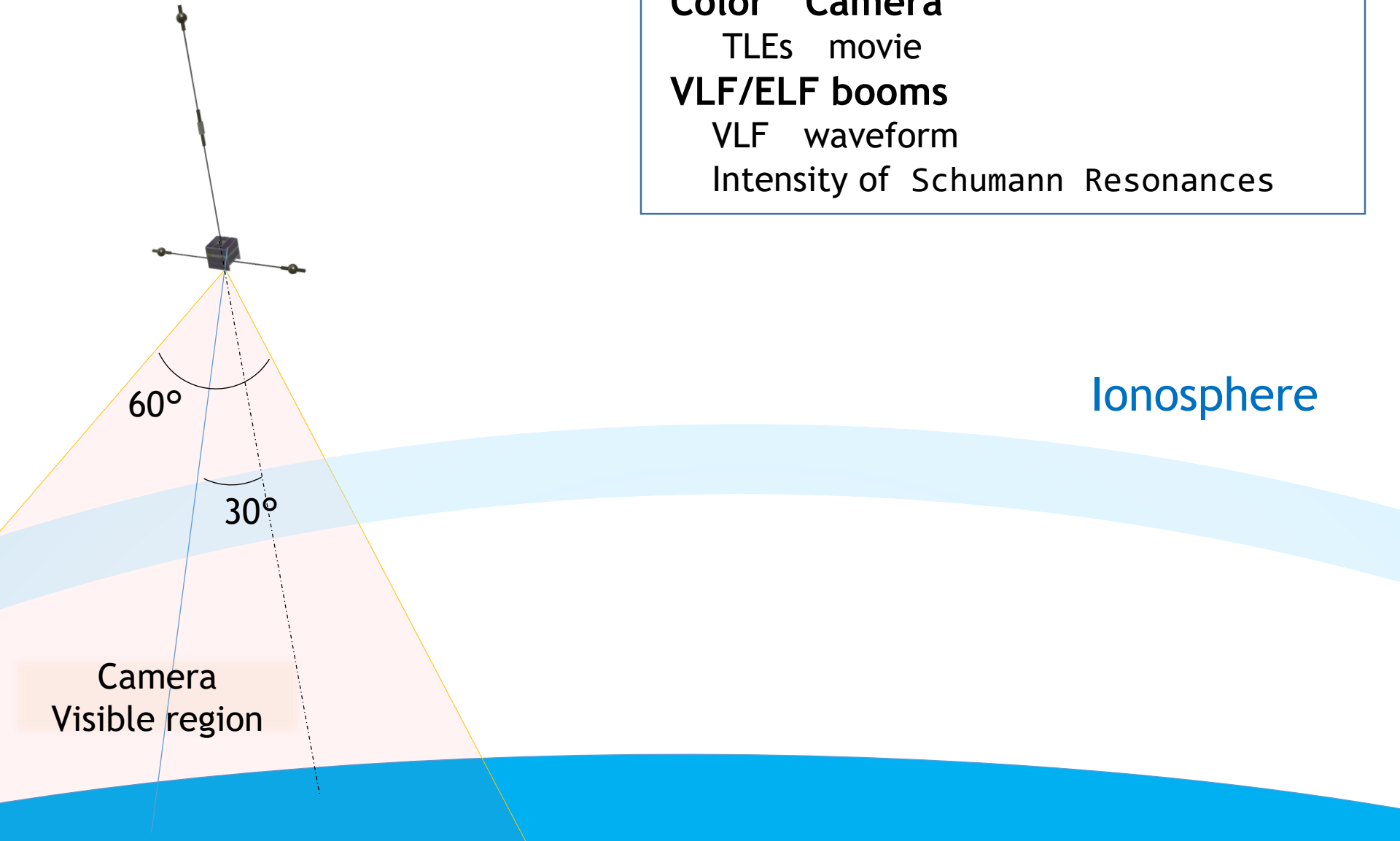
**Color Camera**

TLEs movie

**VLF/ELF booms**

VLF waveform

Intensity of Schumann Resonances



Ionosphere

Camera  
Visible region

# Plan A : Schumann resonance and TLE observation

<Detector and object >

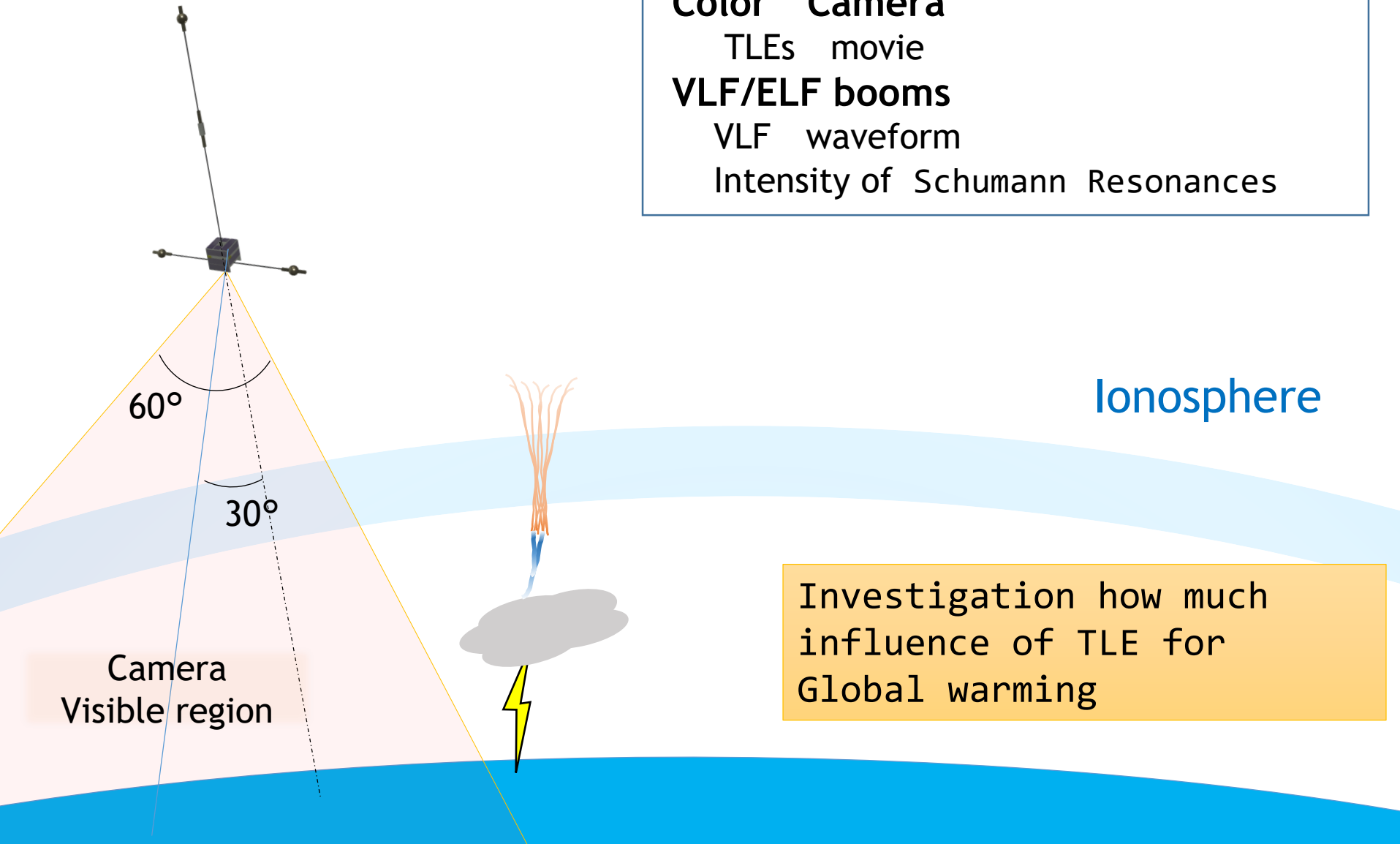
**Color Camera**

TLEs movie

**VLF/ELF booms**

VLF waveform

Intensity of Schumann Resonances



Ionosphere

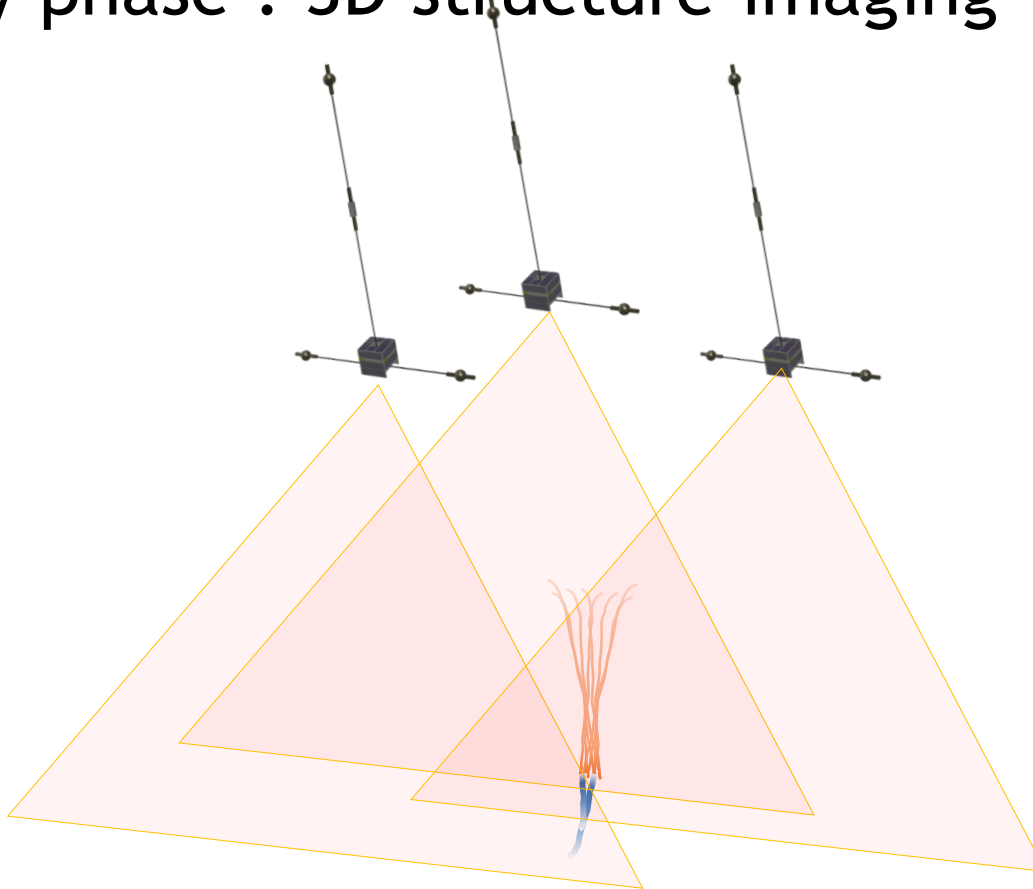
Investigation how much influence of TLE for Global warming

Camera Visible region



# Plan B

## Early phase : 3D structure imaging



- Constructing 3D structure
- Investigating discharge current for GEC study

# Operation Sequence

Earth quake mission

Latitude distribution

Longitude distribution

Feasibility test

100 events

Tsunami mission

Vertical density Profile

Real time monitoring

Lightning mission

Verification of Lightning prediction

TOA

Global warming mission

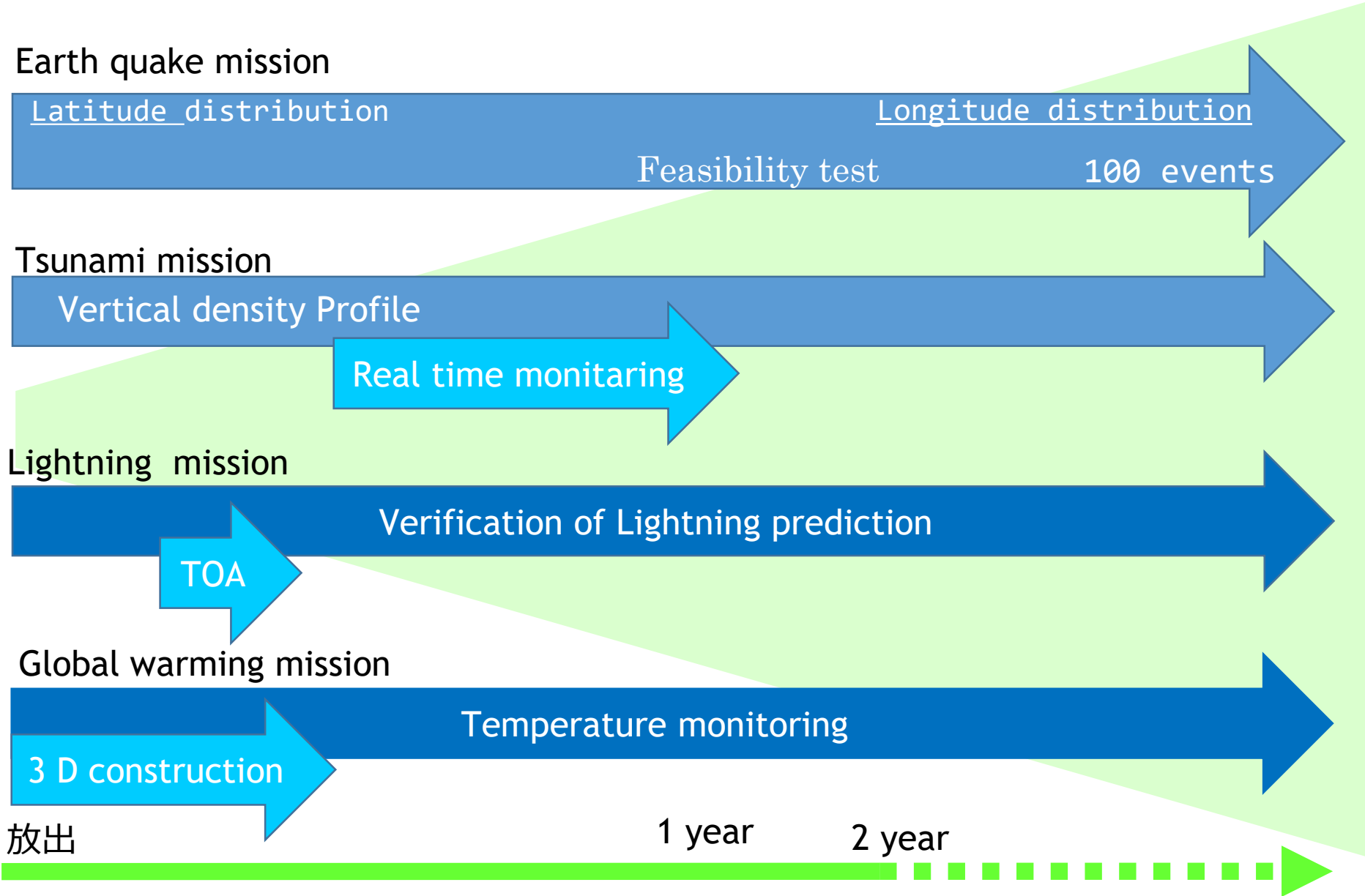
Temperature monitoring

3 D construction

放出

1 year

2 year

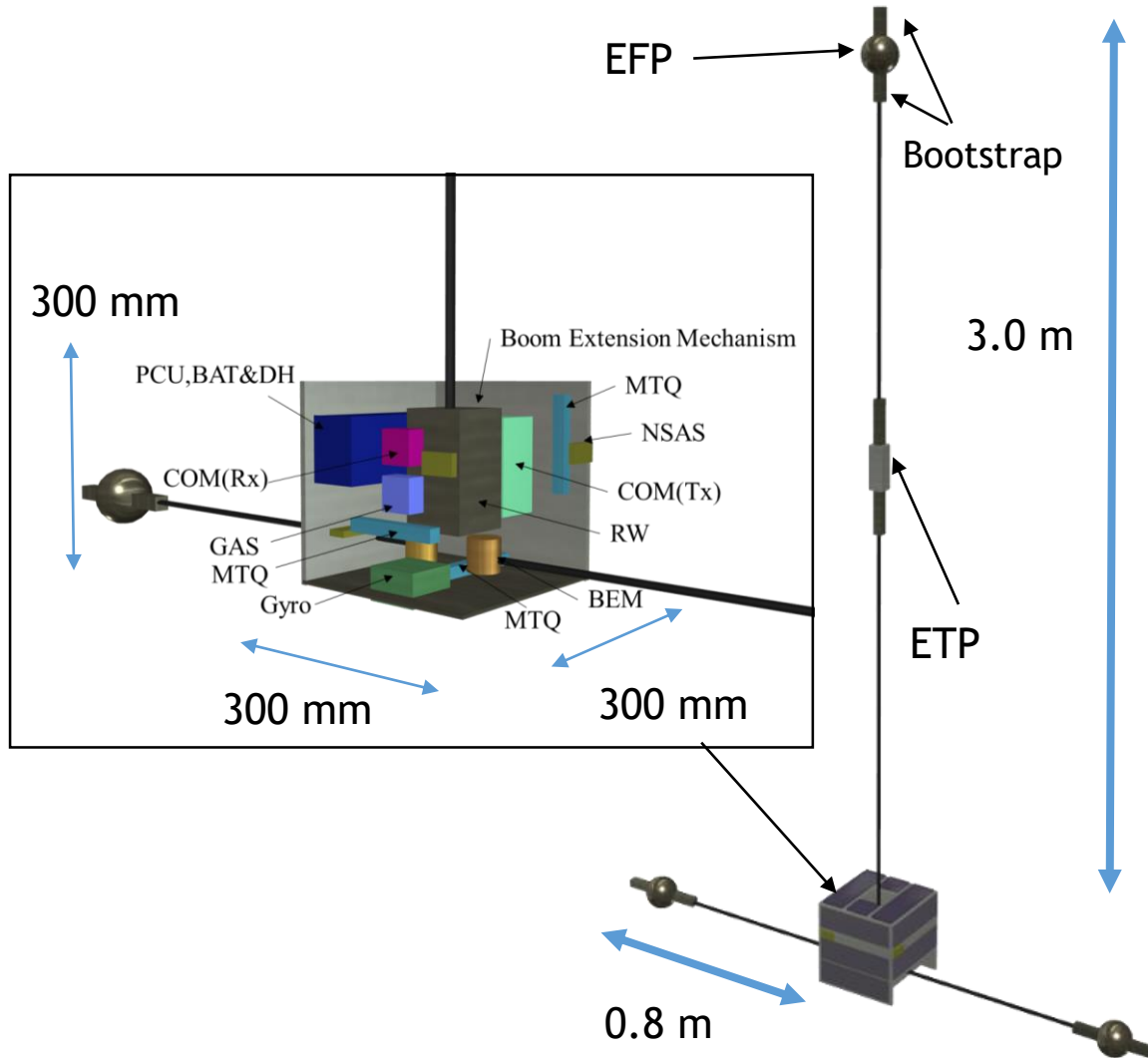


# Mass and Power Distribution

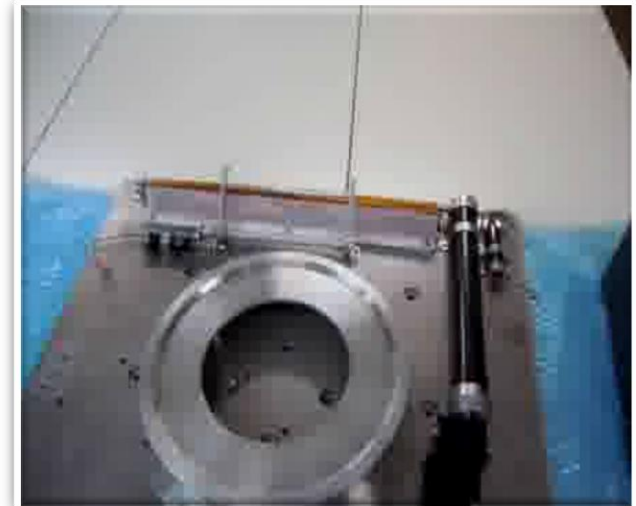
	Components	Qty.	Mass[kg]	Power[W]
EPS	Solar Panel	1	0.6	
	Power control Unit	1	1.5	1.5
	BAT	1	1.5	
COM	Antenna	2	0.1	N/A
	Transmitter and Receiver	1	0.74	TX ON: 4.6
				TX OFF: 1
DH	On-board Computer	1	0.8	5
ADCS	Geomagnetic Acquisition Sensor	1	0.14	0.1
	NSAS	6	0.3	0.9
	MEMS Gyro	1	0.2	0.3
	IR Earth Sensor	1	0.03	0.13
	Magnetic Torquer	3	1.2	1.5
	Reaction Wheel	3	3.3	2.1
<b>Mission</b>	EFP	3	0.3	0.3
	ETP	1	0.1	0.2
	CSI	2	9.4	0.8
	OPC	1	0.02	1
	GTO	1	0.2	1.5
STR	HEATER	1	0.1	1
	STRUCTURE + Harness	1	16.97	N/A
	BOOM (Long)	1	1.1	N/A
	BOOM (Short)	2	1.4	N/A
<b>TOTAL</b>			<b>40</b>	<b>20.9</b>

# Spacecraft System Overview

## Components layout



Boom



First natural frequency of longer boom is estimated to be around 3 Hz.

# Attitude Determination and Control

Mission requirement for electric field measurement

	determination	Control
DC	1	7
AC	5	37

(degree)



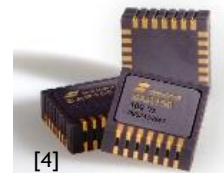
[1]



[2]

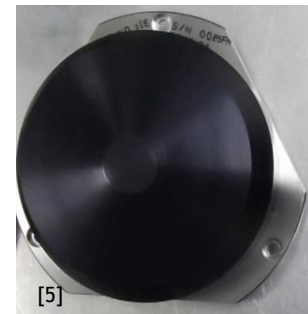


[3]



[4]

- Attitude determination with
  - 6-sun sensors
  - Earth sensor (or star sensor)
  - 3-axes geomagnetic aspect sensor
  - 3-axes mems gyro
- Zero-momentum control with 3 axis RWs
- Unloading with 3-axes magnetic torquers
- **Avoidance of the boom natural frequency**



[5]

[1] <http://www.axelspace.com>  
 [2] Kaplan. C., LEO Satellites: Attitude Determination and Control Components ; Some Linear Attitude Control Techniques  
 [3] <http://www.tierra.co.jp/>  
 [4] <http://www.sensor.com/>  
 [5] <https://makesat.com/products/reaction-wheel>

# Communication

## S-band

- Up: 4 kbps, PCM-PSK-PM
- Down: 64-300 kbps, BPSK

## Operation

- High frequency -> Japan & Norway GS
- Low frequency -> Japan GS

## Expected communication data amount

### Only Japan GS

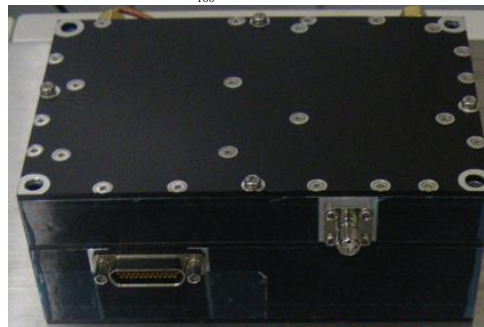
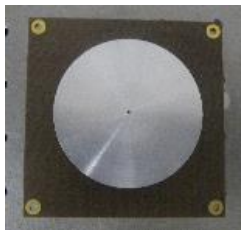
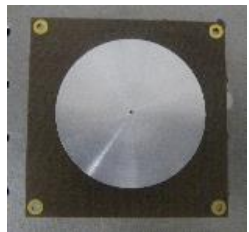
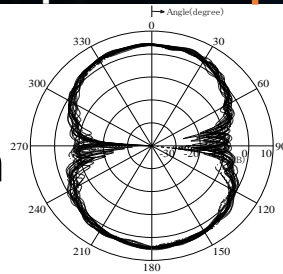
- 1.7MB/Day @ 4kbps
- 100 MB/Day @ 256 kbps ⇔ Requirement 70 MB /Day



Access Summary	Norway	Japan
Access Frequency [num./Day]	14.7	5.3
Averaged Access Duration [hour/Day]	3.0	0.9

## 2 patch antenna

-> omni antenna pattern for emergency



## Downlink Link Margin @ 256kbps

- Larger than 0 dB for worst case

## Uplink Link Margin

- Larger than 10 dB for worst case

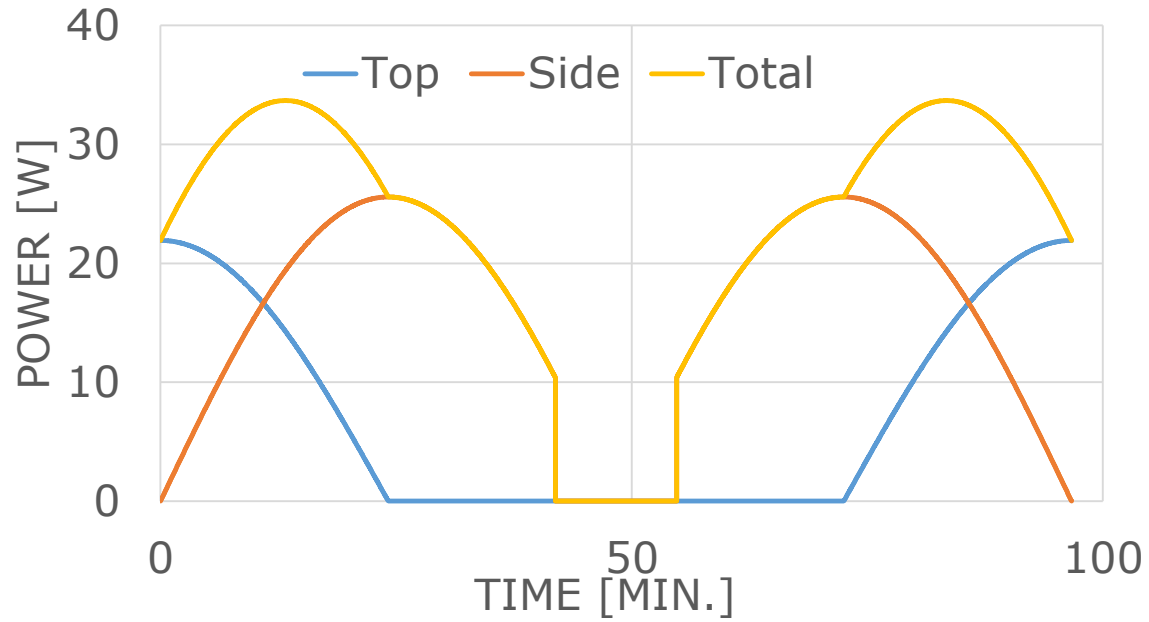
Enough Link Margin,  
Enough data communication amount

# Power

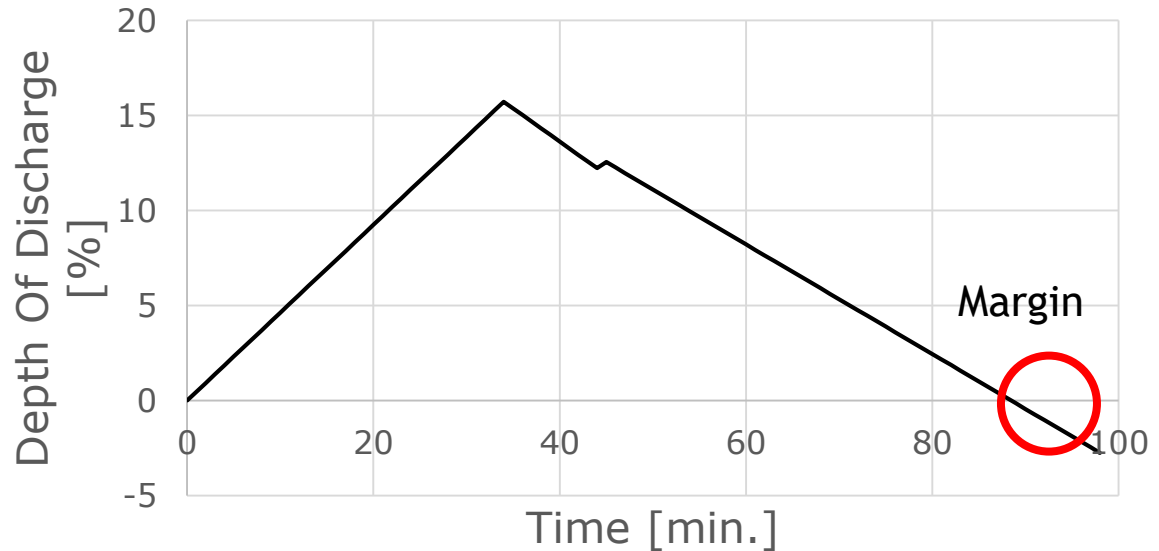
Maximum Power Supply  
-> 34 w

Body mount (7S2P / 7S3P)  
(Solar cells are covered with SnO<sub>2</sub>)

## Generating power



## Power Balance



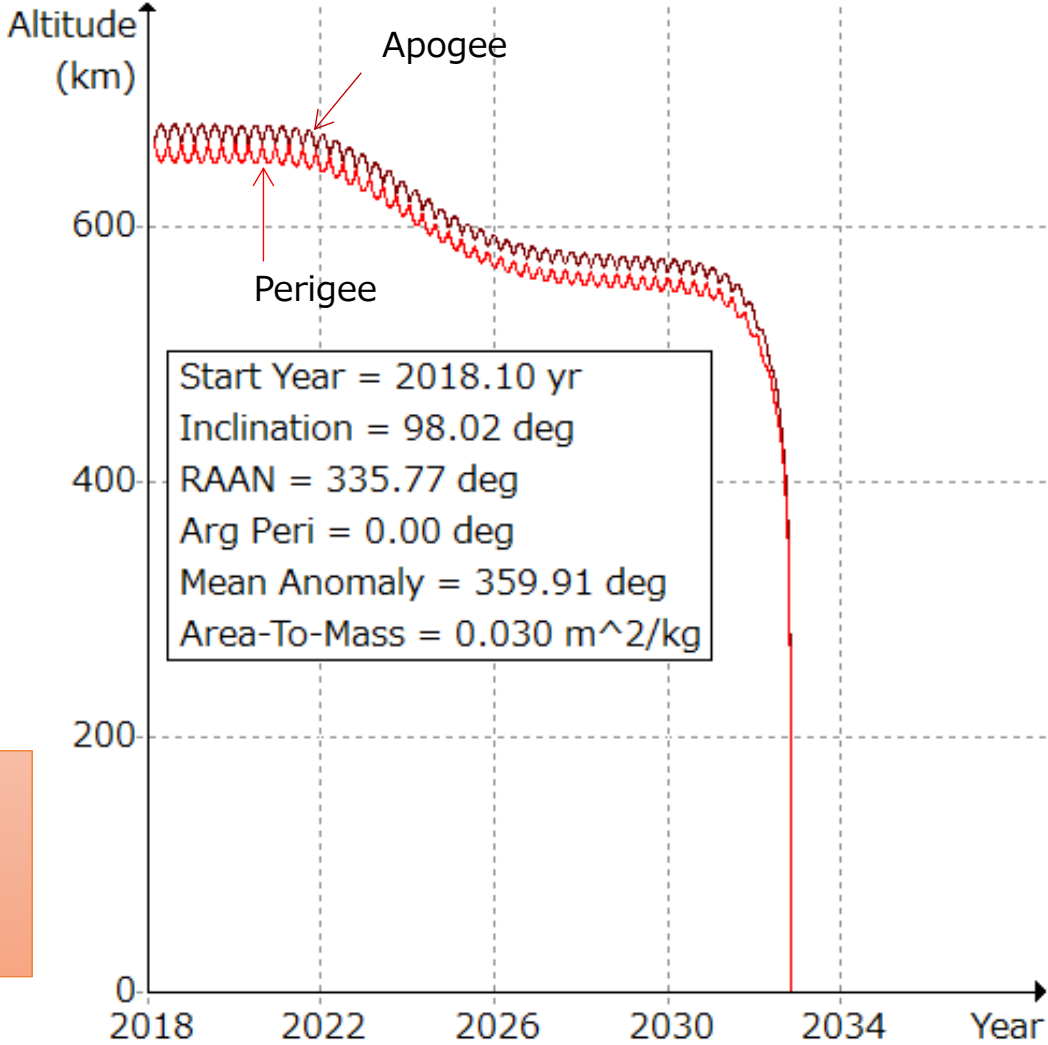
# De-orbiting Analysis

DOM-1500



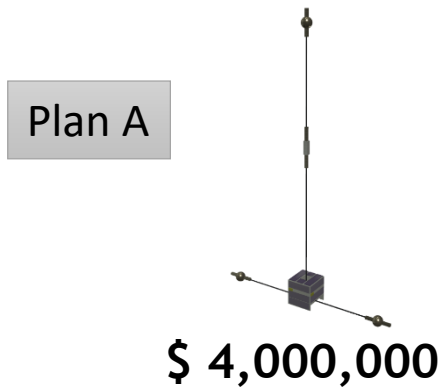
Effective Area :  
total surface area / 4

De-orbiting can be achieved  
within 20 years





# Cost Estimate and Funding



## Team and Funding Sources:

Several university groups

Competitive scientific research funding

e.g. Grant-in-Aid for Scientific  
Research (KAKENHI)



International university group

Several national institute in Japan

Space Agency: JAXA

National project

10 years after our mission ...



14 Oct 2016 02:10.000

Time Step: 10.00 sec

# Concluding remarks

- Natural disasters often produce signals of atmospheric and space electricity before and after the event.
- Electricity monitoring is applicable to other natural disasters such as land sliding, volcano eruption, and forest fire.
- Space-based monitoring of such signals contributes to mitigate the disaster.
- On-board electromagnetic measurement is small, simple and matured technology and low cost.
- This project will show “innovative science” using microsatellites.

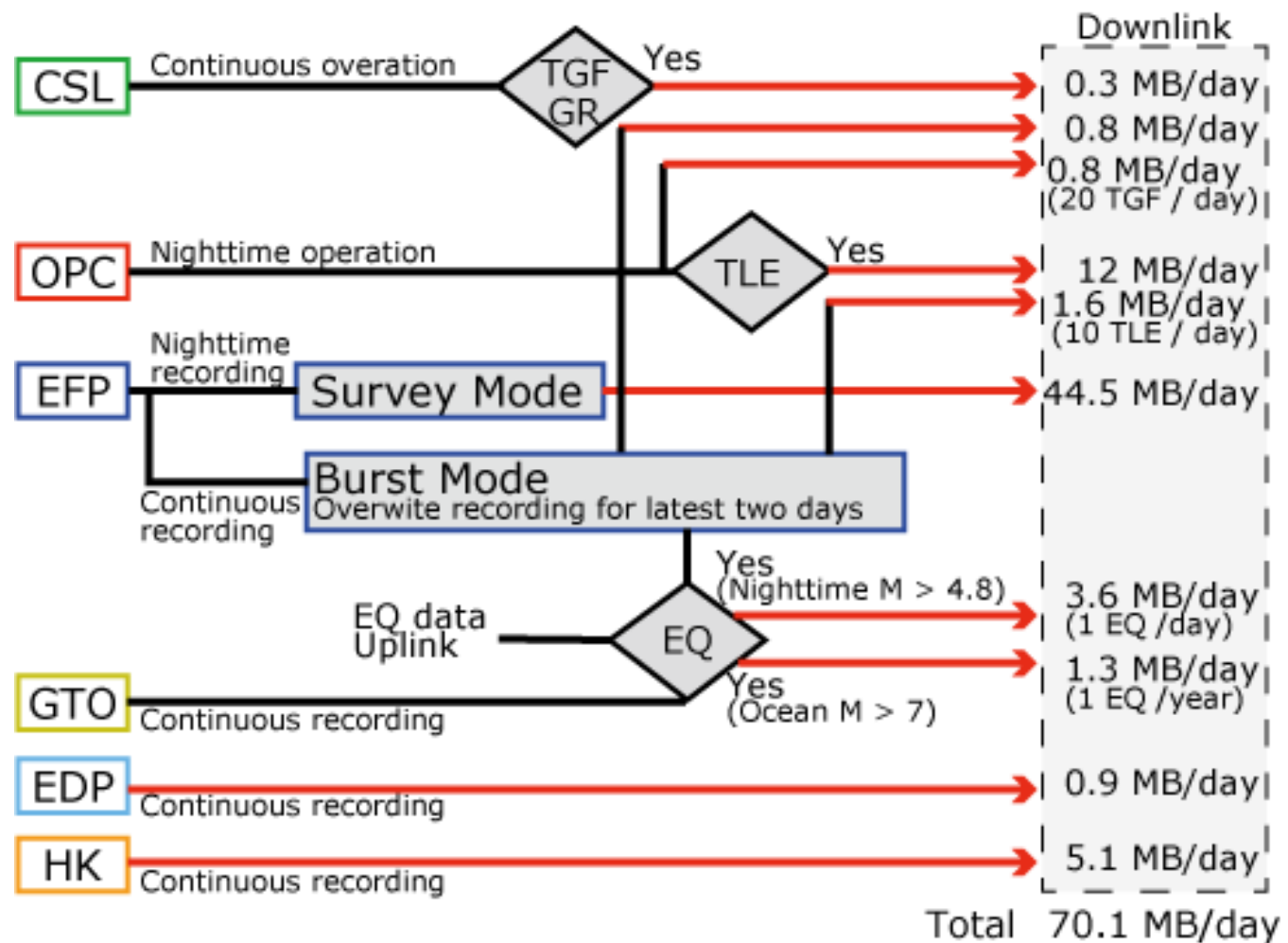
**Thank you for your attention!!**

Disturbance Estimation

-> 0.14Nms/rev

Gravity gradient	$0.16 \times 10^{-5}$	Nm
Solar Radiation Pressure	$0.0022 \times 10^{-5}$	Nm
<b>Residual magnetism</b>	<b><math>2.3 \times 10^{-5}</math></b>	Nm
Aerodynamic	$0.011 \times 10^{-5}$	Nm

# Mission data



# boom



[1] <http://www.i-qps.net/i-qps/service.html>

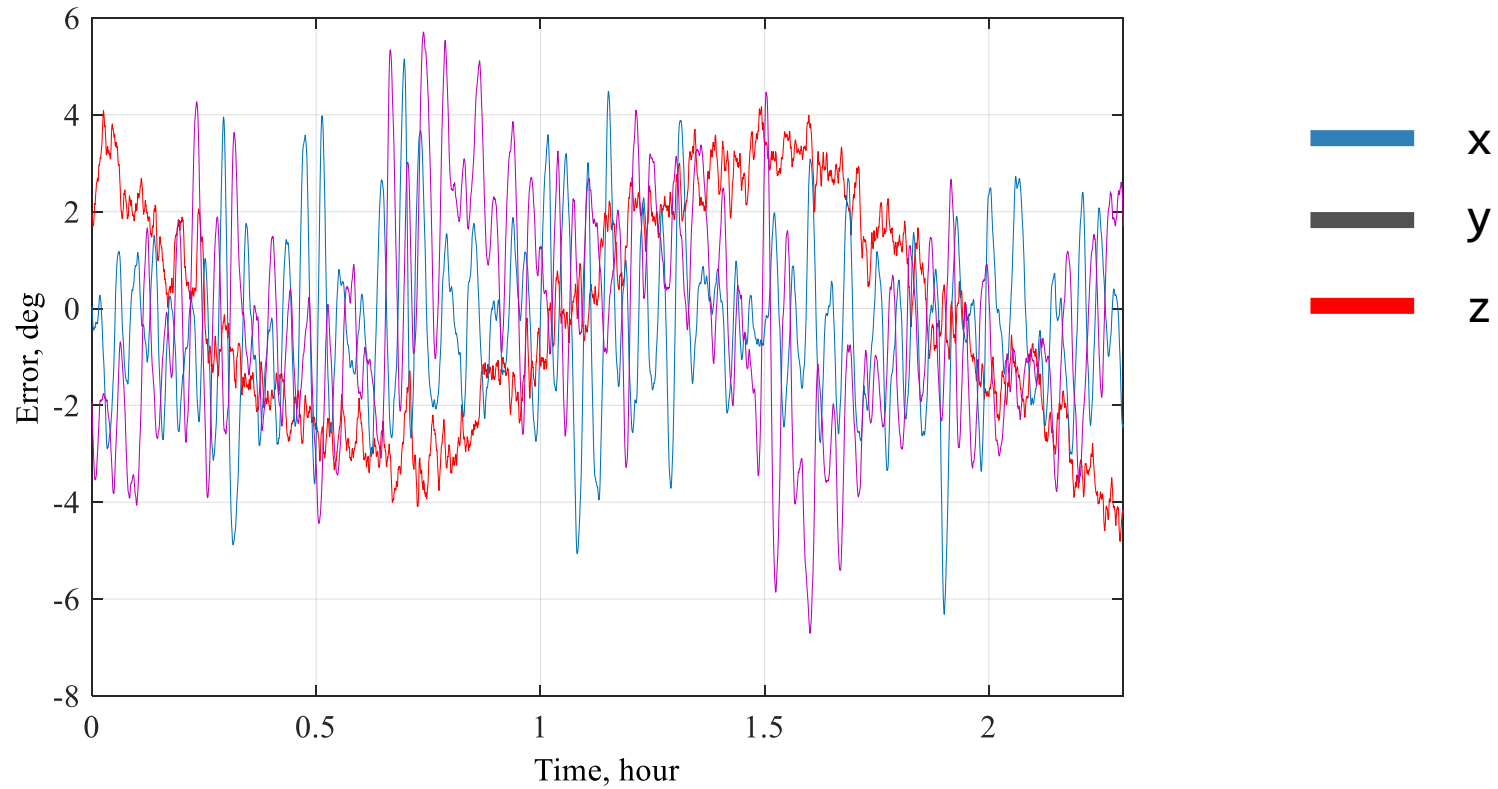
[2] Yokomatsu et al; Deployment Experiments on Stiffened Tri-axial Tubular CFRP Boom for Boom-Membrane Integrated Space Structures

# Mass and Power Distribution

	Components	Qty.	Mass[kg]	Power[W]
EPS	Solar Panel	1	0.6	
	Power control Unit	1	1.5	1.5
	BAT	1	1.5	
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	Transmitter and Receiver	1	0.74	TX ON: 4.6
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DH	On-board Computer	1	0.8	5
ADCS	Geomagnetic Acquisition Sensor	1	0.14	0.1
	NSAS	6	0.3	0.9
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	IR Earth Sensor	1	0.03	0.13
	Magnetic Torquer	3	1.2	1.5
	Reaction Wheel	3	3.3	2.1
Misson	EFP	3	0.3	0.3
	ETP	1	0.1	0.2
	CSI	2	9.4	0.8
	OPC	1	0.02	1
	GTO	1	0.2	1.5
STR	HEATER	1	0.1	1
	STRUCTURE + Harness	1	16.97	N/A
	BOOM (Long)	1	1.1	N/A
	BOOM (Short)	2	1.4	N/A
<b>TOTAL</b>			<b>40</b>	<b>20.9</b>

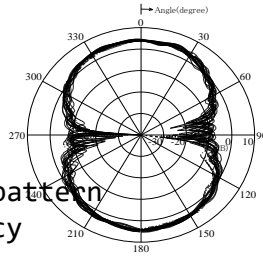
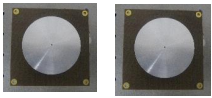


# Simulation Results of Attitude Control incl. determination error



# Communication

- S-band system
  - Up: 4 kbps, PCM-PSK-PM
  - Down: 64-300 kbps, BPSK
- Japan and Norway GS
- Expected communication data amount
  - 7MB/Day @ 4kbps
  - 530 MB/Day @ 300kbps



[1]

2 patch antenna  
-> omni antenna pattern  
for emergency

## Downlink Link Margin @ 256kbps

	Unit	min	max
Transmit EIRP	dBW	-7.13	-1.63
Transmitter Power	W	0.300	0.300
Internal Loss	dB	-3.9	-3.9
Antenna Gain	dBi	2	7.5
Free propagation range	km	2485.20	666.00
Atmospheric absorption loss	dB	-0.00305	-0.00305
Rain fade	dB	-0.08	-0.00557
Receiving G/T	dB/K	6.54	6.54
Antenna Gain	dBi	35.58	35.58
Internal Loss	dB	-2.50	-2.50
System noise temp.	K	450	450
Receiving C/N0	dBHz	60.54	77.55
Required Eb/N0(BER : 10 <sup>-5</sup> )	dB	6	6
Symbol Rate	kbps	256	256
Required C/N0	dBHz	-60.08	-60.08
Hardware loss	dB	-1.5	-1.5
<b>Link Margin</b>	<b>dB</b>	<b>0.20</b>	<b>17.21</b>

## Access Summary

	Norway	Japan
Access Frequency [num./Day]	14.7	5.3
Averaged Access Duration [hour/Day]	3.0	0.9



## Uplink Link Margin

	Unit	min	max	Command Margin	min	max	
Transmitter Power	W	10	10	Required Eb/No	9.303	9.303	dB
Internal Loss	dB	-0.9	-0.9	Modulation Loss	11.14	11.14	dB
Antenna Gain	dBi	36.2	36.2	Demodulation Loss	2.4	2.4	dB
Transmit EIRP	dBW	45.3	45.3	Coding Gain	2.498	2.498	dB
Free propagation range	km	2485.2	666	Required C/No	56.36	56.36	dBHz
Free propagation loss	dB	-166.8	-155.33	<b>Link Margin</b>	<b>11.3</b>	<b>22.78</b>	<b>dB</b>
Atmospheric absorption loss	dB	-0.003	-0.003				
Rain fade	dB	-0.079	-0.079	Carrier Margin	min	max	
Antenna Gain	dBi	-10	-10	Required S/N	10.00	10.00	dB
Internal Loss	dB	-0.90	-0.90	Modulation Loss	2.325	2.325	dB

Enough Link Margin, Enough data communication

Received C/N0	dBHz	67.70	79.13				
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## Access Summary

	Norway	Japan
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Expected communication data amount

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Antenna Gain	dBi	2	7.5
Free propagation range	km	2485.20	666.00
Atmospheric absorption loss	dB	-0.00305	-0.00305
Rain fade	dB	-0.08	-0.00557
Receiving G/T	dB/K	6.54	6.54
Antenna Gain	dBi	35.58	35.58
Internal Loss	dB	-2.50	-2.50
System noise temp.	K	450	450
Receiving C/N0	dBHz	60.54	77.55
Required Eb/N0(BER : 10 <sup>-5</sup> )	dB	6	6
Symbol Rate	kbps	256	256
Required C/N0	dBHz	-60.08	-60.08
Hardware loss	dB	-1.5	-1.5
<b>Link Margin</b>	<b>dB</b>	<b>0.20</b>	<b>17.21</b>

## Uplink Link Margin

	Unit	min	max	Command Margin	min	max	
Transmitter Power	W	10	10	Required Eb/No	9.303	9.303	dB
Internal Loss	dB	-0.9	-0.9	Modulation Loss	11.14	11.14	dB
Antenna Gain	dBi	36.2	36.2	Demodulation Loss	2.4	2.4	dB
Transmit EIRP	dBW	45.3	45.3	Coding Gain	2.498	2.498	dB
Free propagation range	km	2485.2	666	Required C/No	56.36	56.36	dB Hz
Free propagation loss	dB	-166.8	-155.33	<b>Link Margin</b>	<b>11.33</b>	<b>22.78</b>	<b>dB</b>
Atmospheric absorption loss	dB	-0.003	-0.003				
Rain fade	dB	-0.079	-0.079	Carrier Margin	min	max	
Antenna Gain	dBi	-10	-10	Required S/N	10.00	10.00	dB
Internal Loss	dB	-0.90	-0.90	Modulation Loss	2.325	2.325	dB
System noise temp.	K	700	700	Required C/No	42.32	42.32	dB
Received G/T	dB/K	-39.35	-39.35	<b>Link Margin</b>	<b>25.37</b>	<b>36.81</b>	<b>dB</b>
Received C/N0	dBHz	67.70	79.13				

Mass	40 kg	
Volume	300 x 300 x 300 mm	
Maximum Power Supply	34 W, Body mount (7S2P / 7S3P) (Solar cells are covered with SnO <sub>2</sub> )	
Power storage	5.8 Ah (Max DOD:15.7 %)	
Lifetime	2 year	
ADCS	Attitude Determination Accuracy	1° NSAS(6), GAS(1), MEMS Gyro(1), IR Earth Sensor(1)
	Attitude Control Accuracy	7° RW(3), MTQ(3)
COM	S-band Up: 4 kbps, PCM-PSK-PM Down: 64-300 kbps, BPSK	
Ground Station	Norway (Svalbard) , Japan	

	Plan A	Plan B
<b>Total</b>	<b>4,000,000</b>	<b>30,000,000</b>
Mission Payloads		
Satellite Bus		
IA&T		
Ground Systems		
Project Management & SE		

# Funding Sources

## Grant-in-Aid for Scientific Research (KAKENHI)

- Up to 300 M JPY expected
- Funded Tohoku University's  
SPRITE-SAT

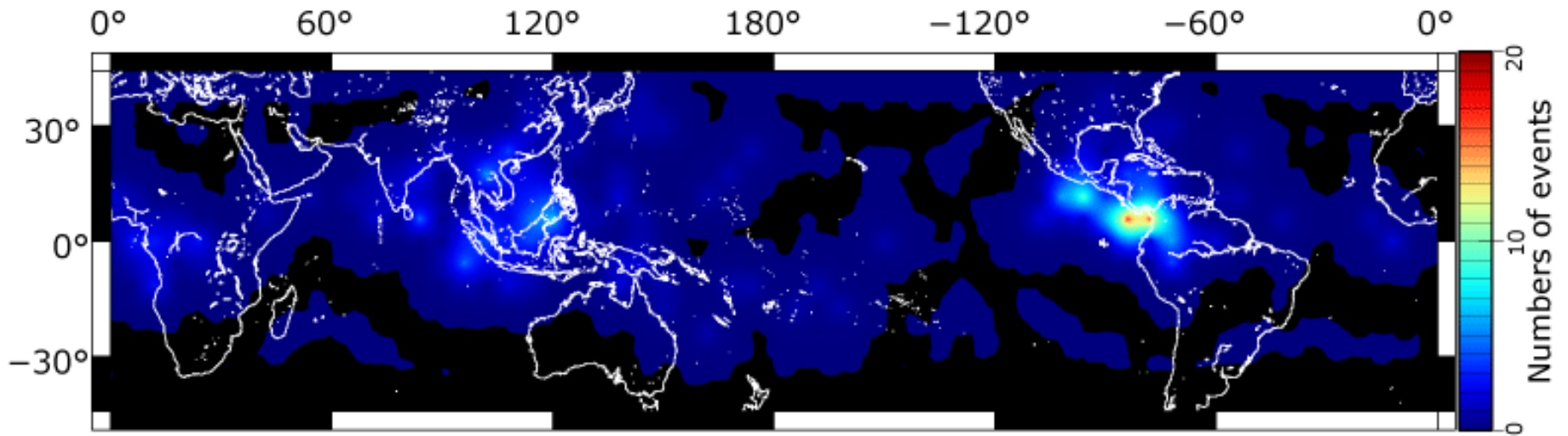


## Official Development Assistance (ODA)

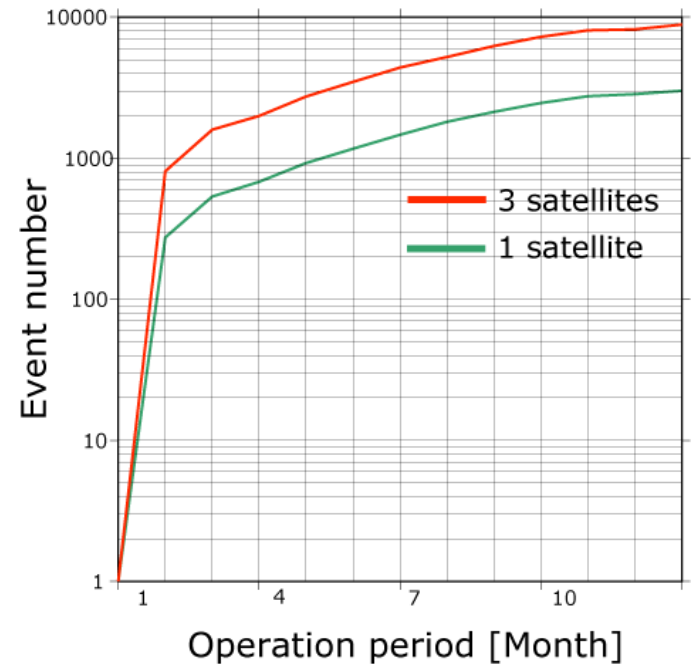
- First case: 40 B JPY to Vietnam
- Especially for countries  
participating in UNISEC's CanSat  
Leadership Training Program



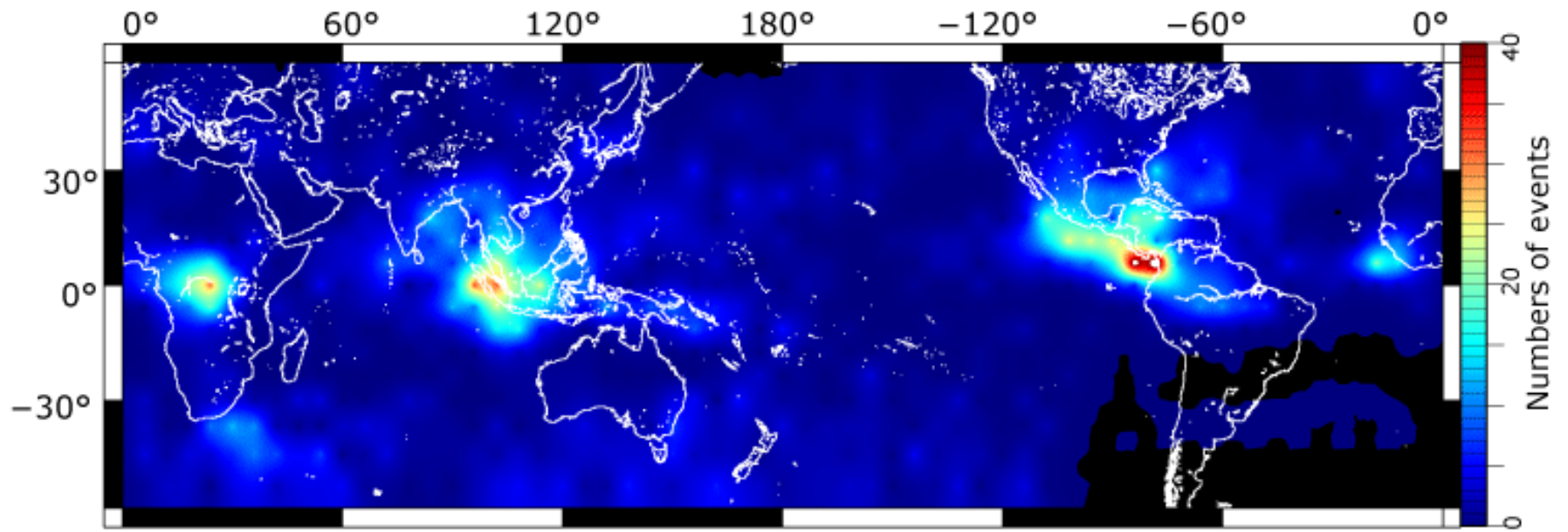
Proposal will be submitted to JSPS (KAKENHI) this month by Prof. Kamogawa for next fiscal year funding.



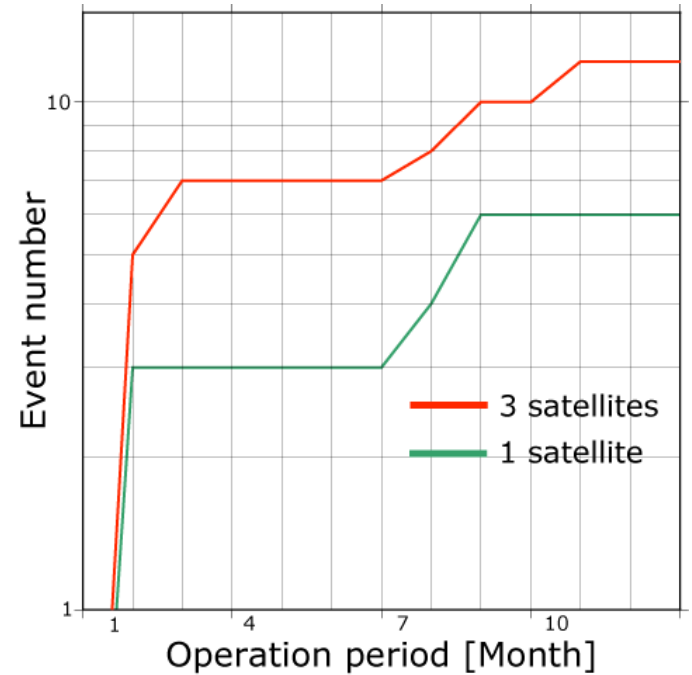
TLE map and Event number



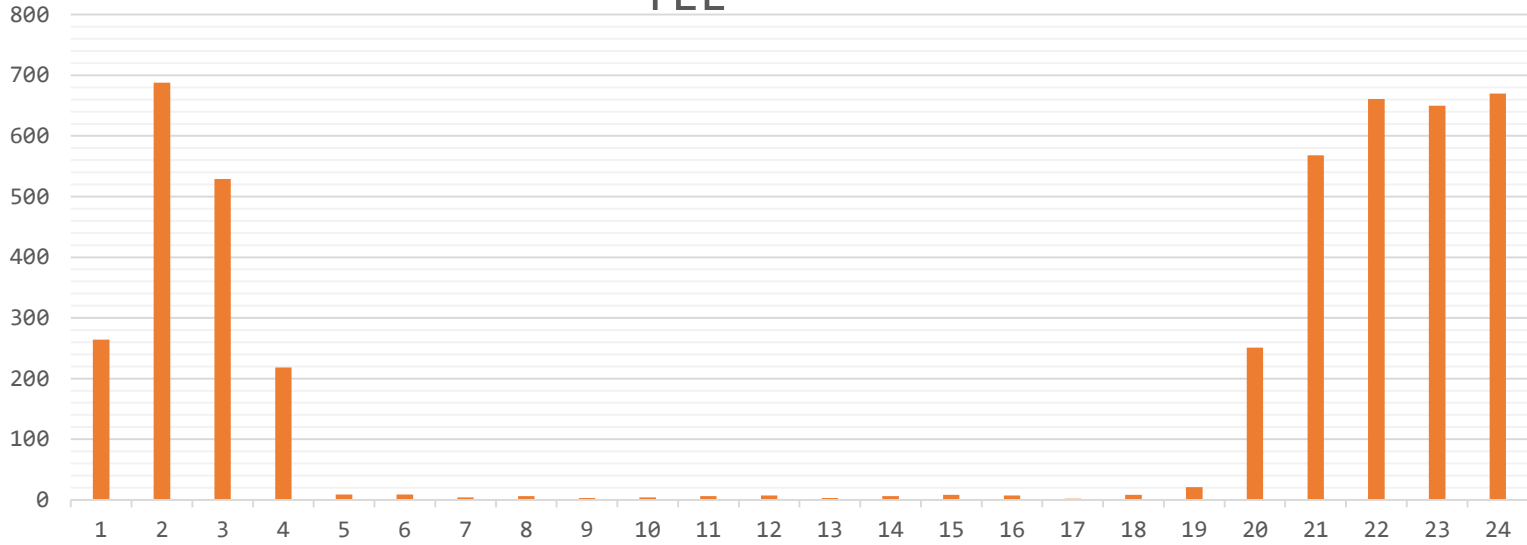




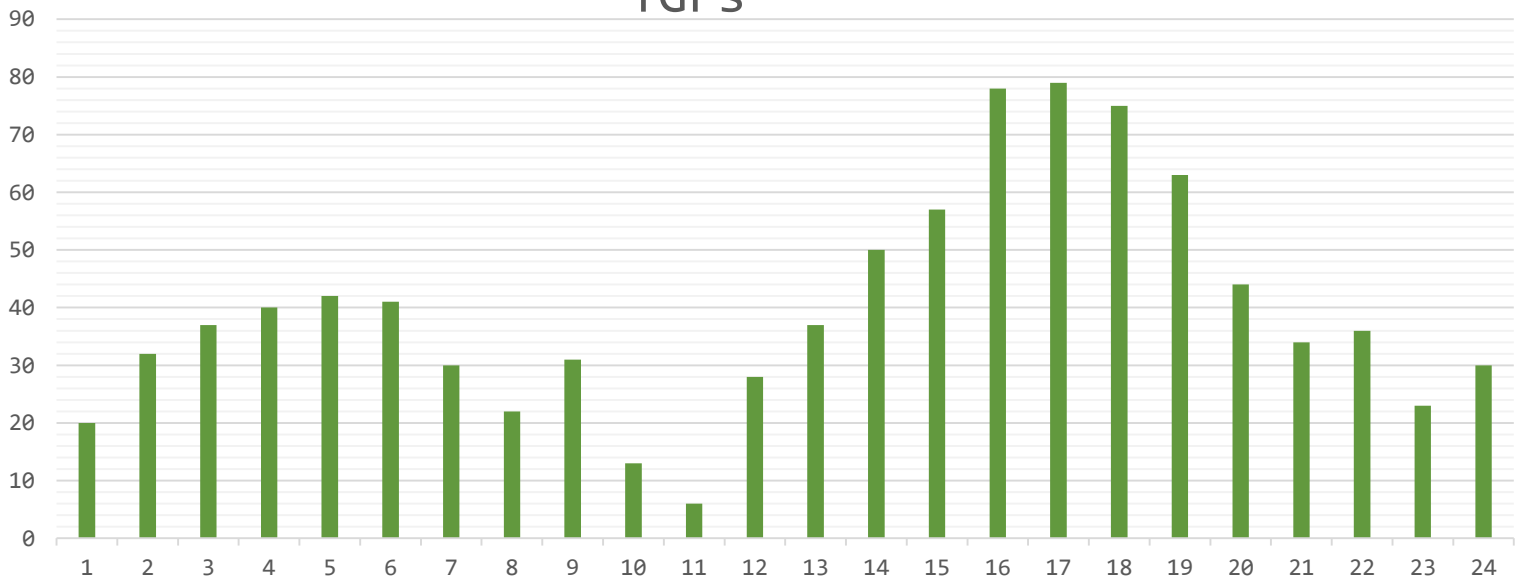
TGFs map and Event number

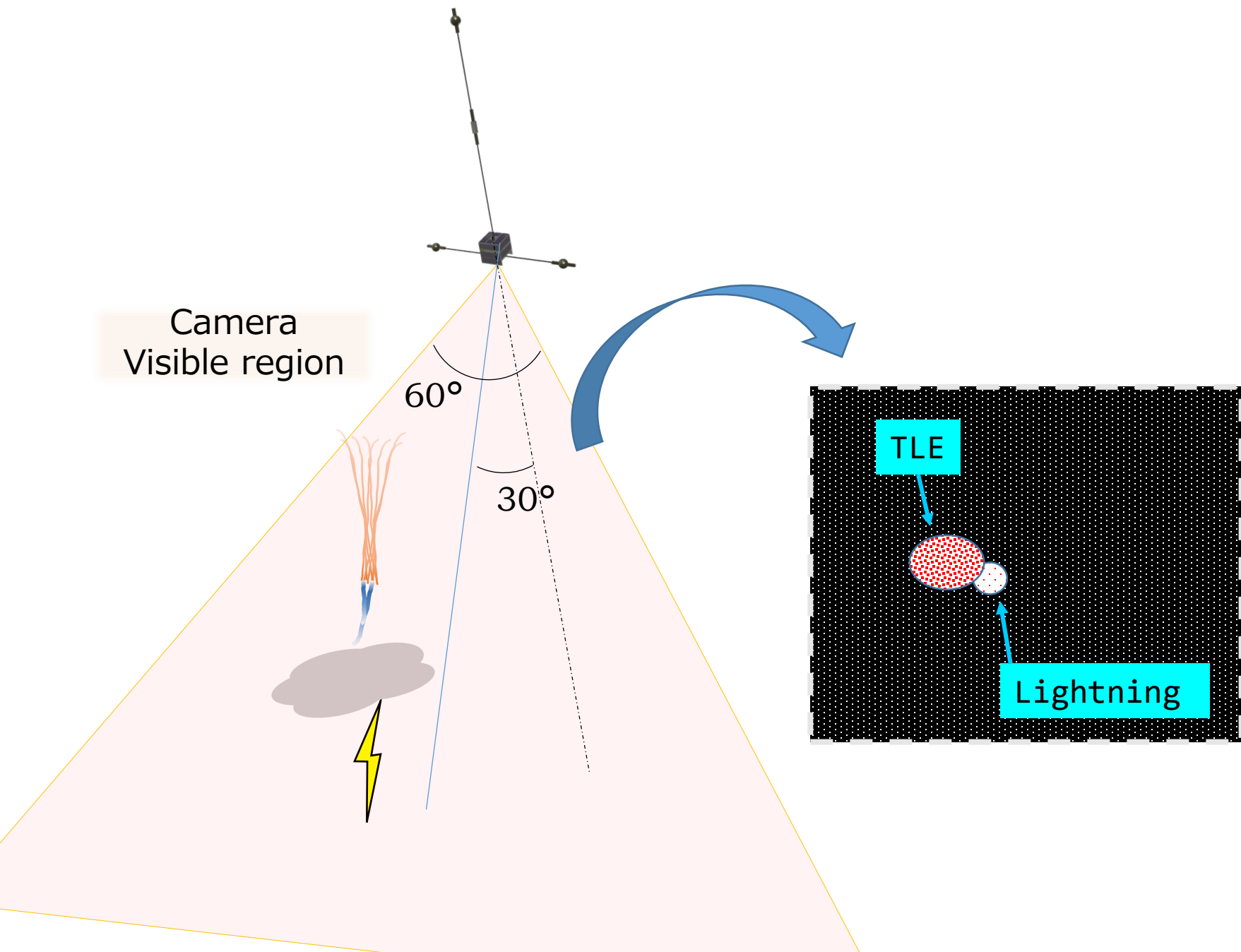


# TLE



# TGFs





Camera Visible region

$60^\circ$

$30^\circ$

TLE

Lightning

# How to estimate TGF and TLE events

# Purpose

estimation of detect number using orbit calculation

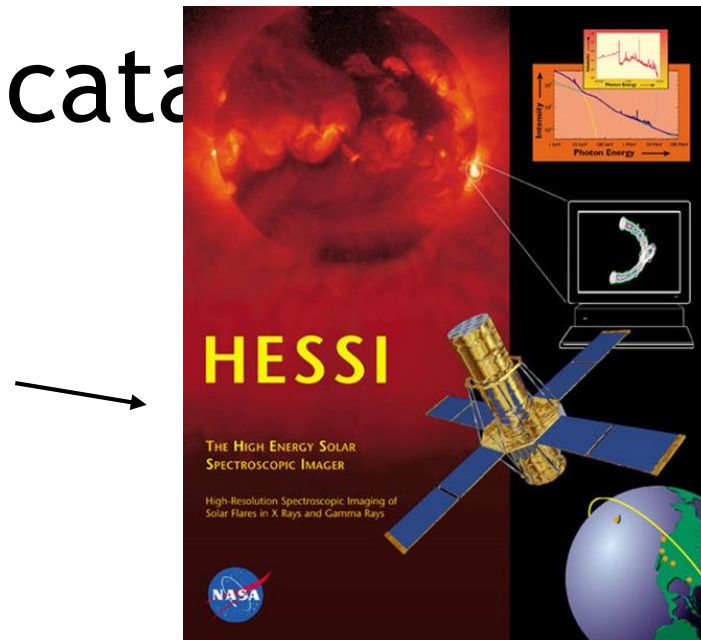
→ We need global cata

	Ground observation
Lightning	WWLLN
Earthquake	USGS
TCE	



JEM-GLIMS mission

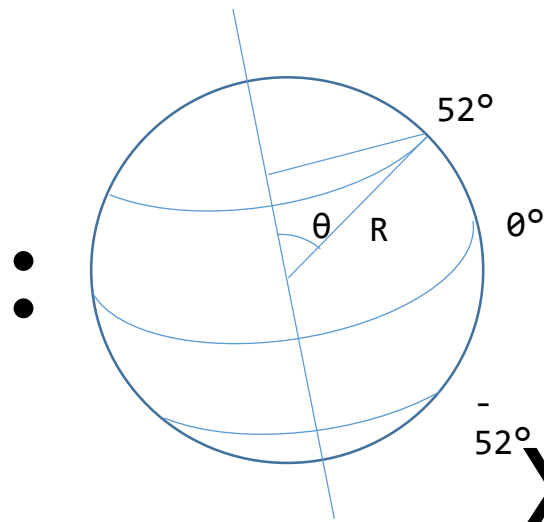
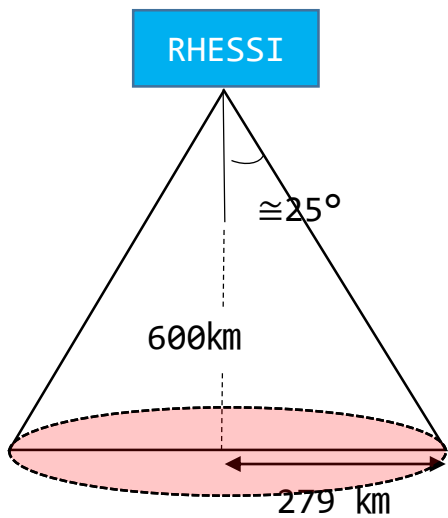
[http://www.nasa.gov/mission\\_pages/station/research/experiments/121.html](http://www.nasa.gov/mission_pages/station/research/experiments/121.html)



<http://hesperia.gsfc.nasa.gov/rhe>

- Satellite have only visible data then
- We estimated whole events by rate of satellite detectable area and orbital area.

# Estimation of whole data amount



Visible area : S1  
 $S1 = \pi r^2 = 3.14 * 279 * 279$   
 $= 2.44 \times 10^5$  [km<sup>2</sup>]

^2]

Ornbitral area : S2  
 $S2 = 2 * \int_0^{(90-38)^\circ} 2\pi R \sin\theta * R d\theta$   
 $= 4\pi R^2 [1 - \cos\theta]_0^{52^\circ}$   
 $= 4 * 3.14 * 6371 * 6371 * (1 - 0.61)$   
 $= 1.98 \times 10^8$  [km<sup>2</sup>]

= 975 :

~~X~~  
 $X = 6.92 \times 10^5$

# How to make TGF standard model

1) Download data set as follows.

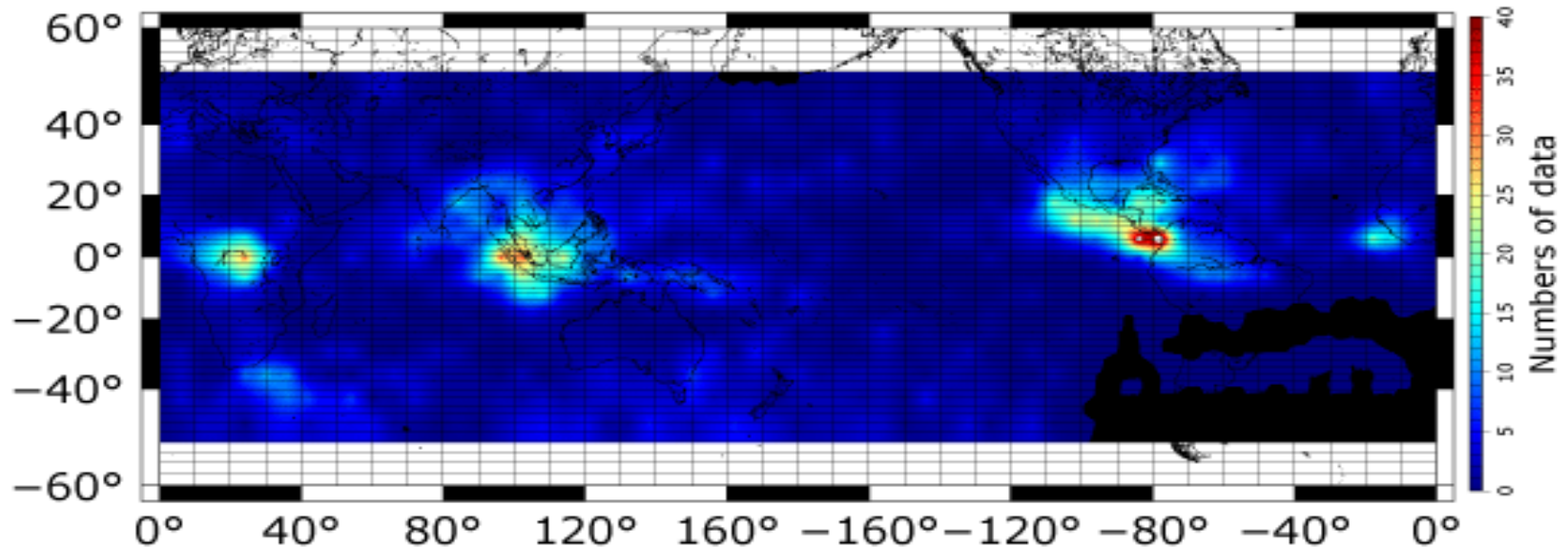
[http://scipp.ucsc.edu/~dsmith/tgflib\\_public/data/](http://scipp.ucsc.edu/~dsmith/tgflib_public/data/)

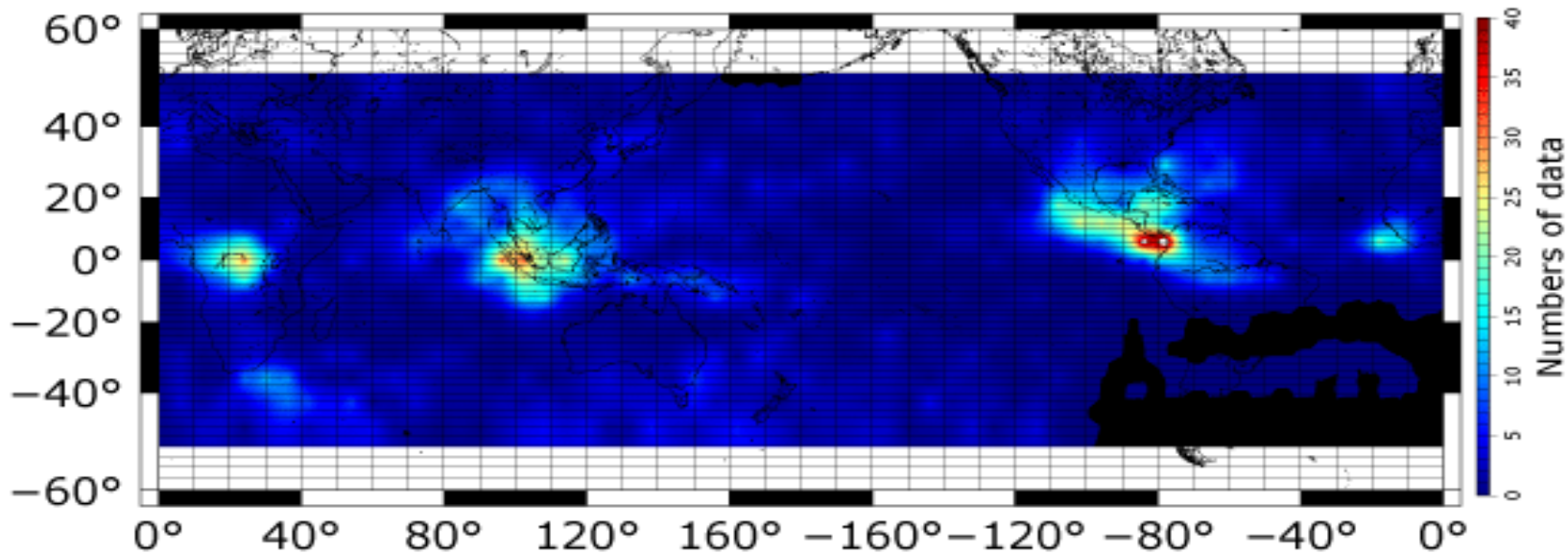
2) Separate events in bin

a) Space :  $6^\circ$  ( longitudinal and latitudinal)

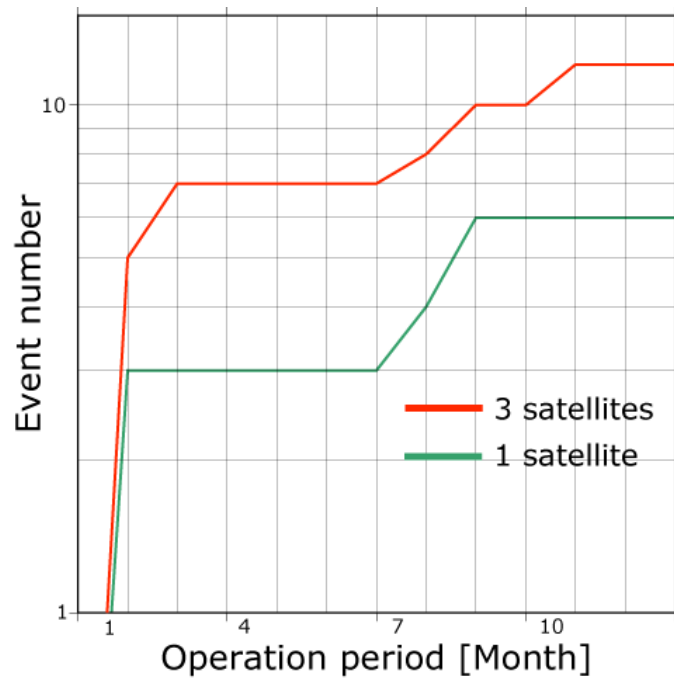
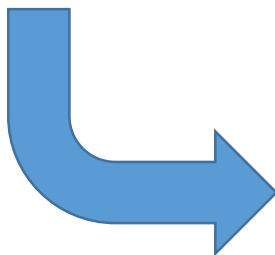
b) Season : Summer ( 5 ~ 10 ) and Winter ( 11 ~ 4 )

c) Time zone : Day ( LT 6 ~ 18 ) and Night ( LT 0 ~ 6, 18 ~ 24 )

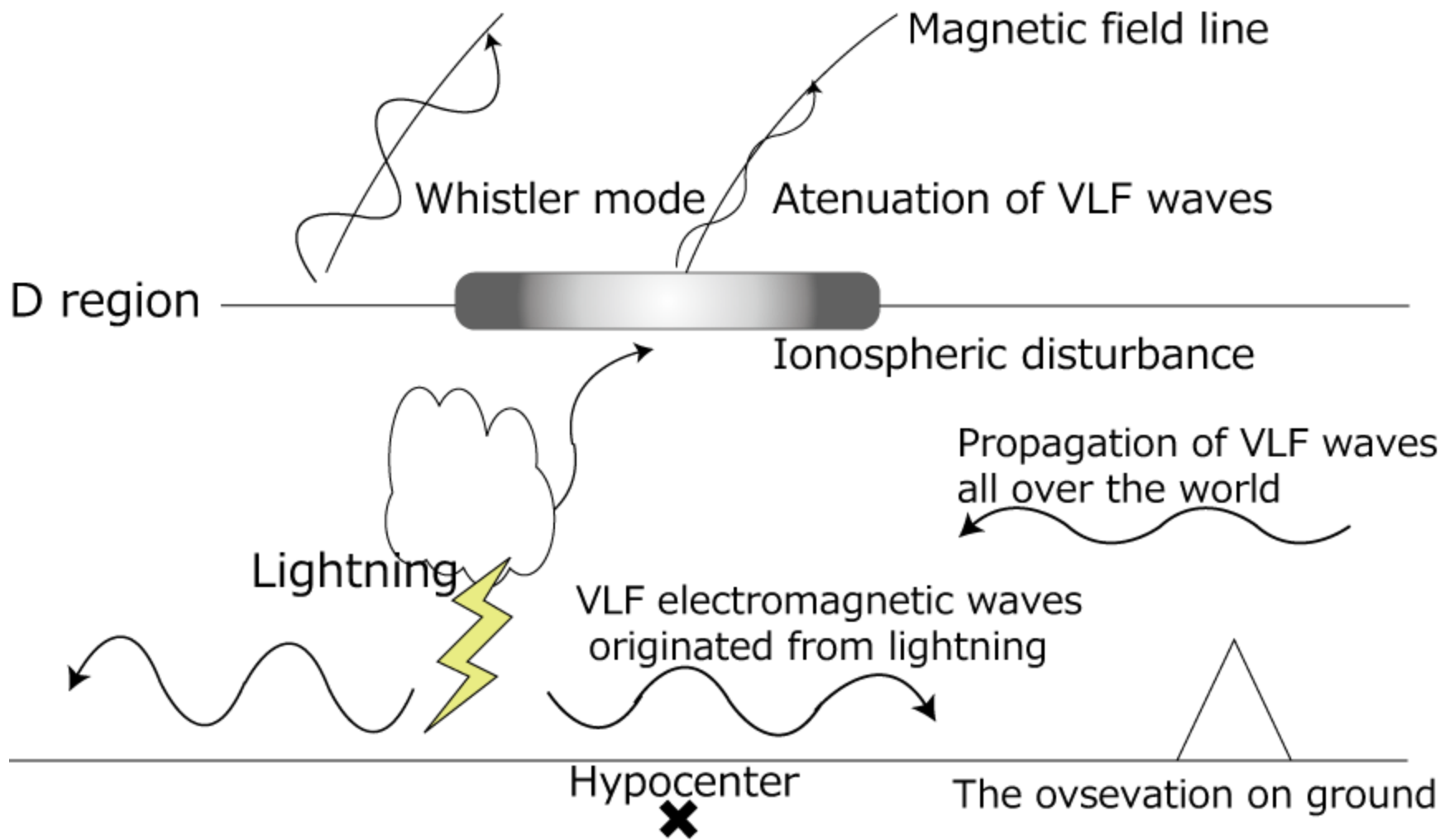


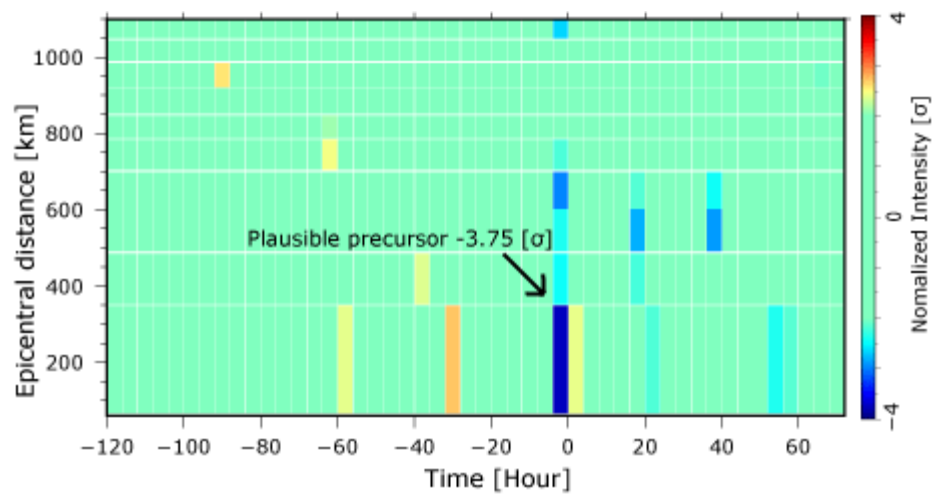


Orbit calculation  
Using GOSAT

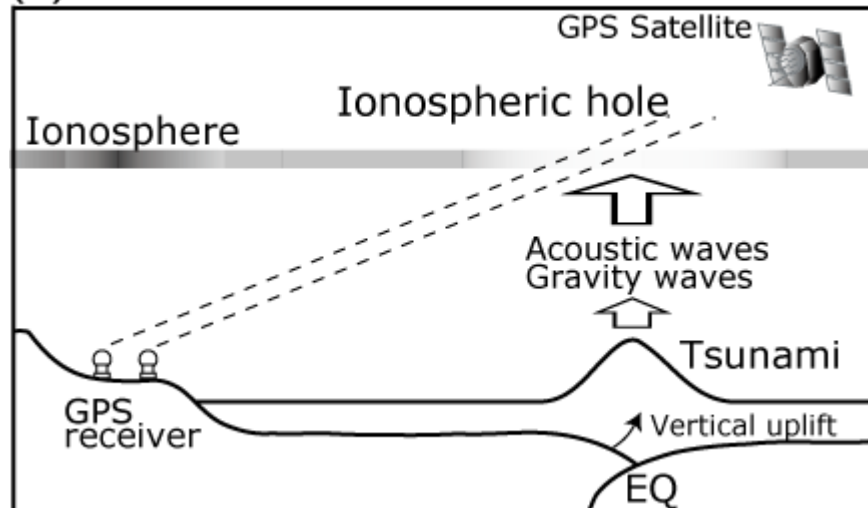




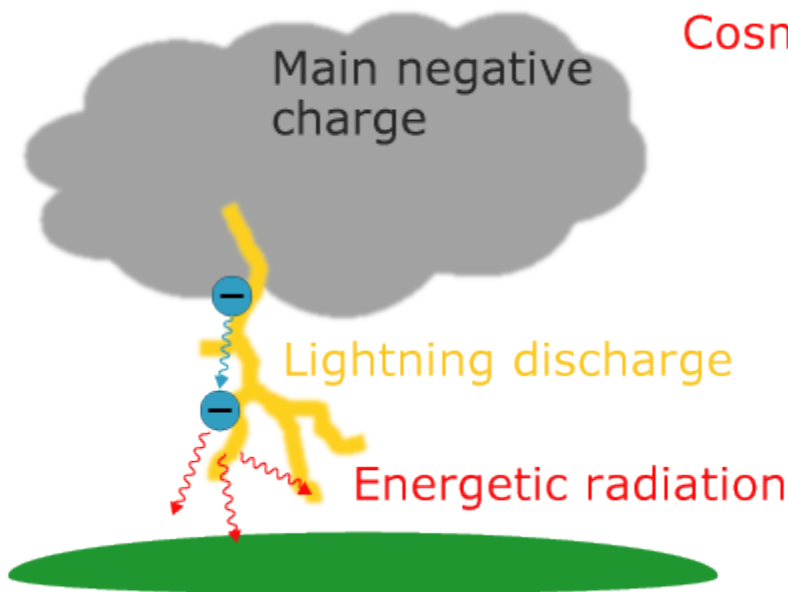




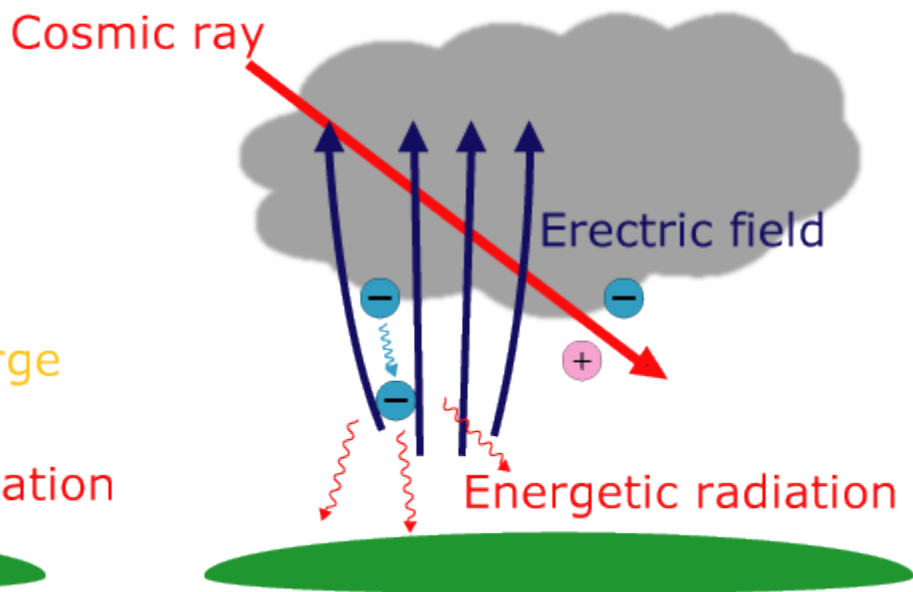
(a)

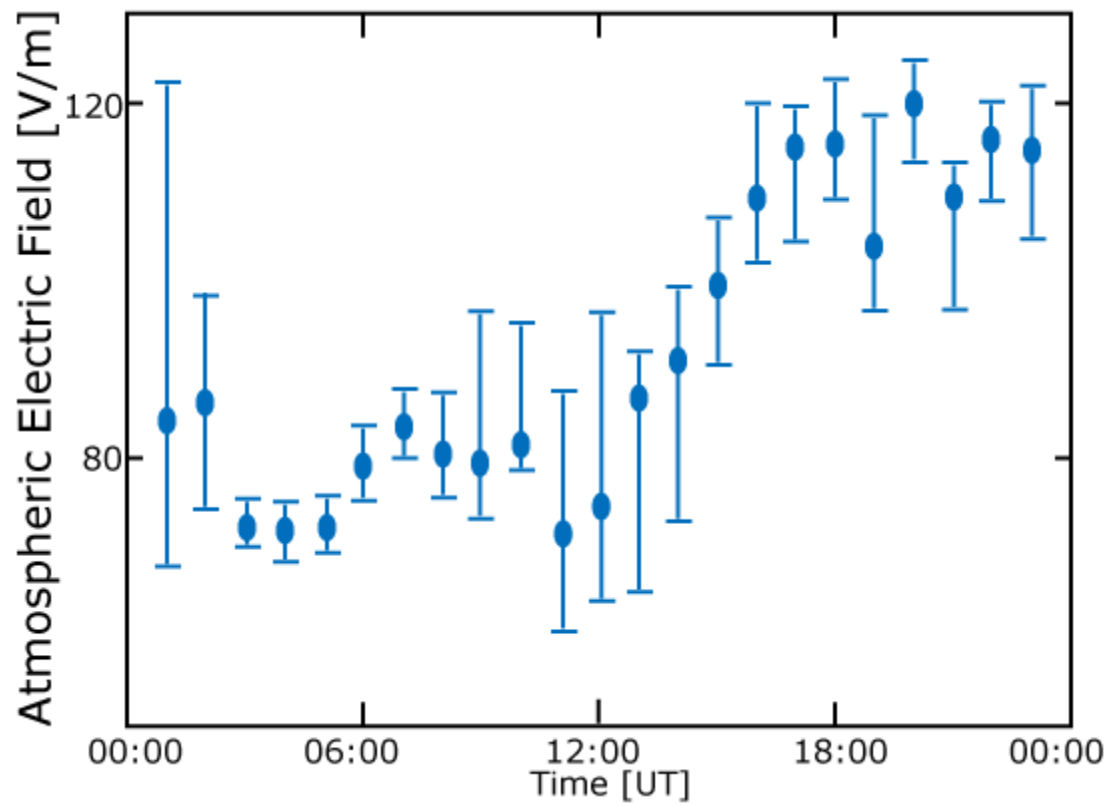


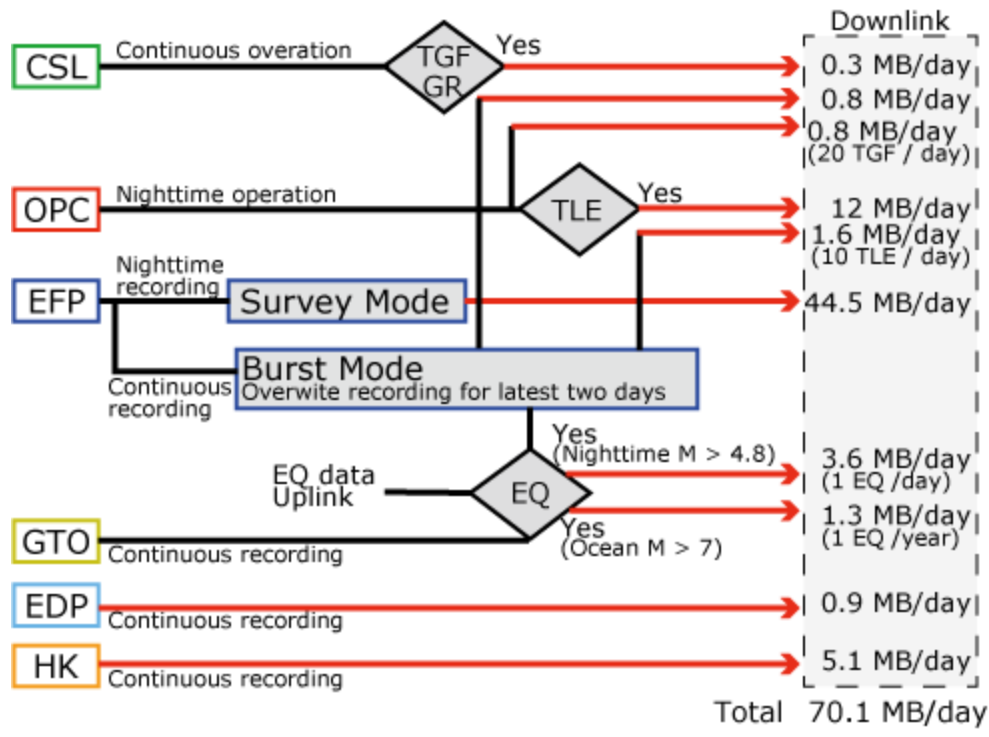
### SBR

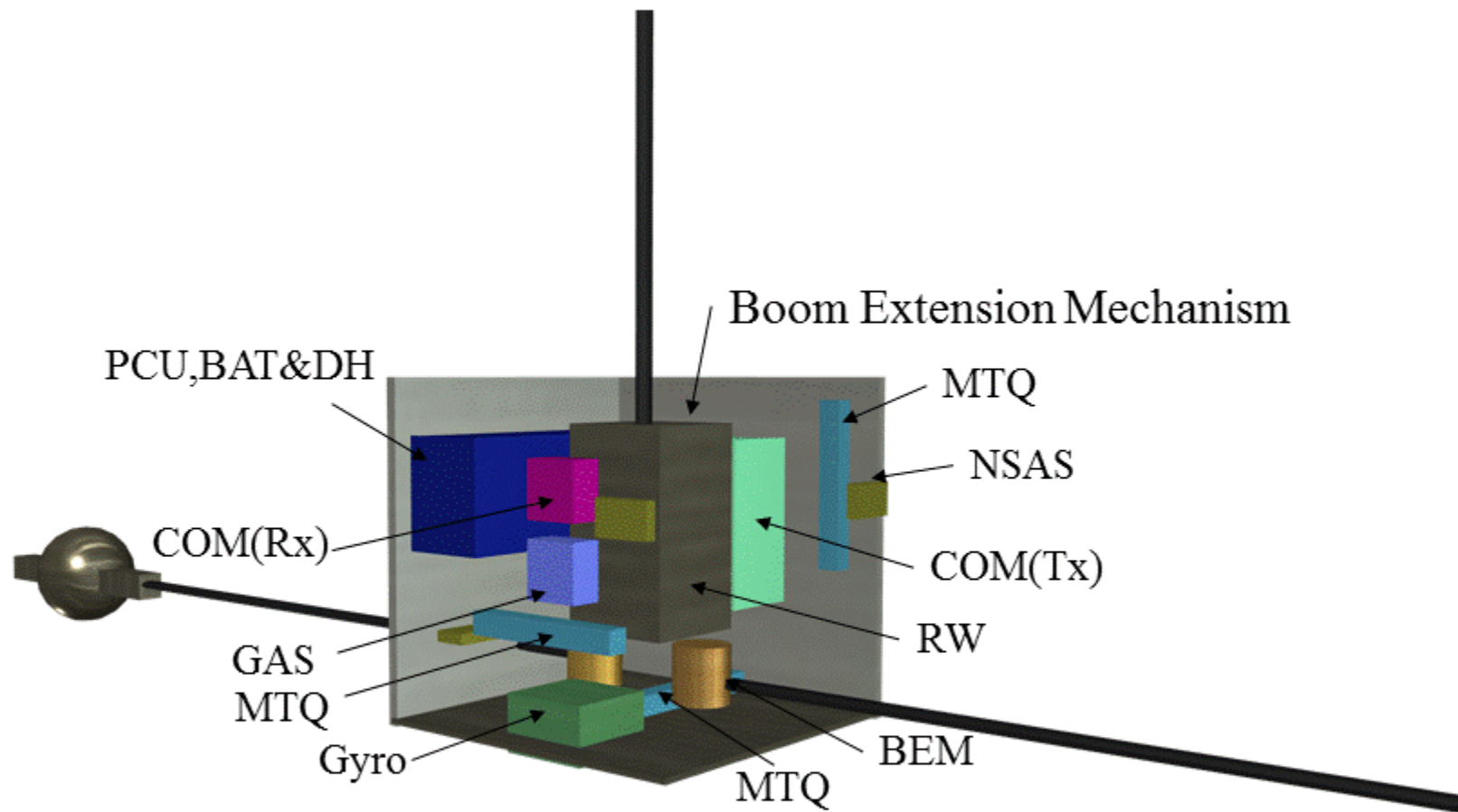


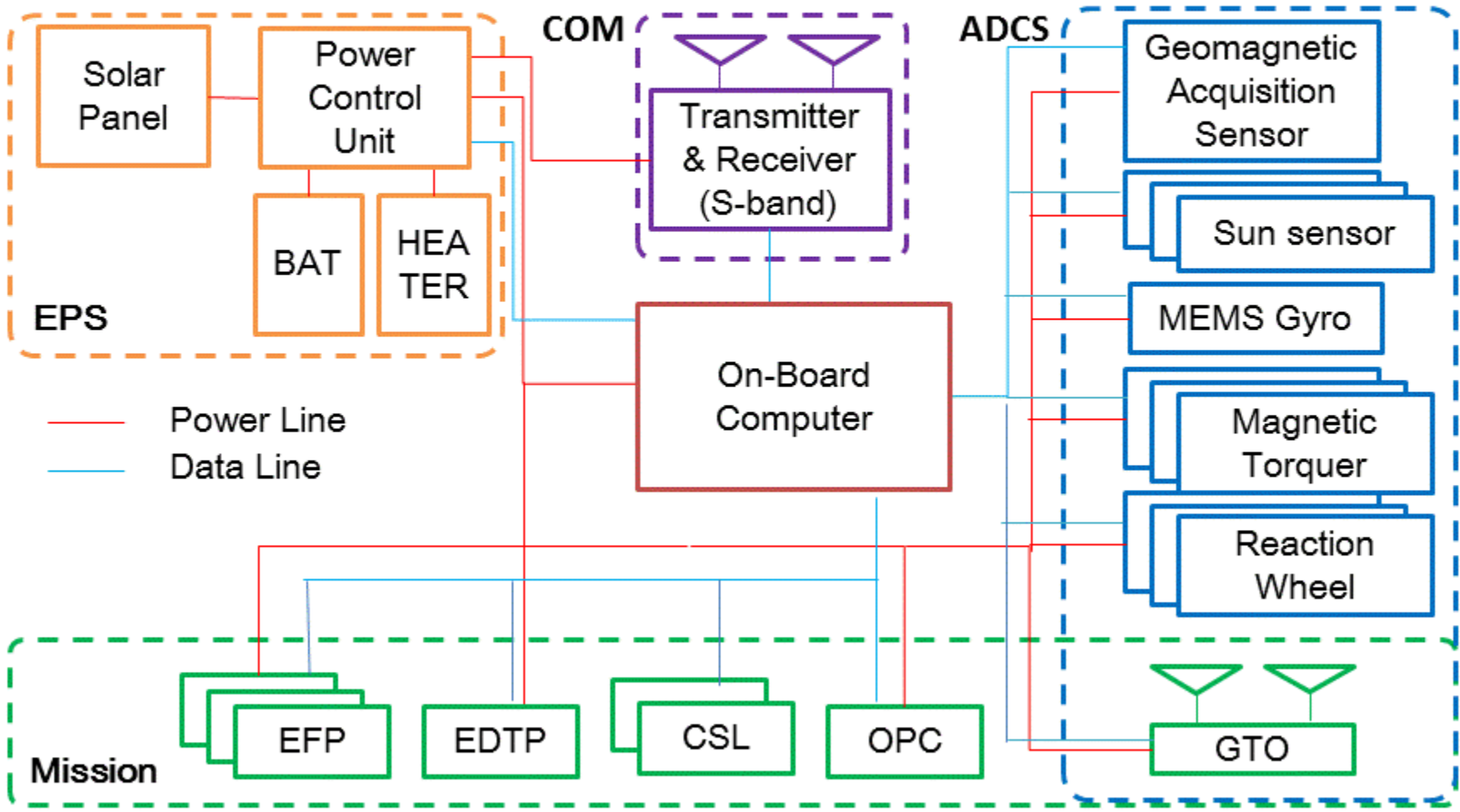
### LBR



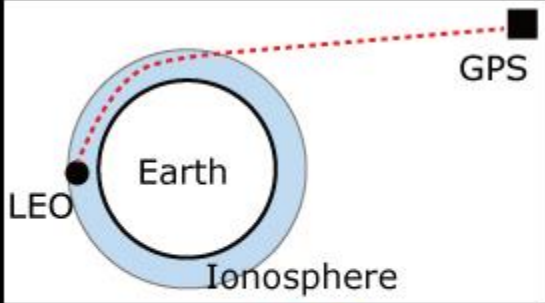




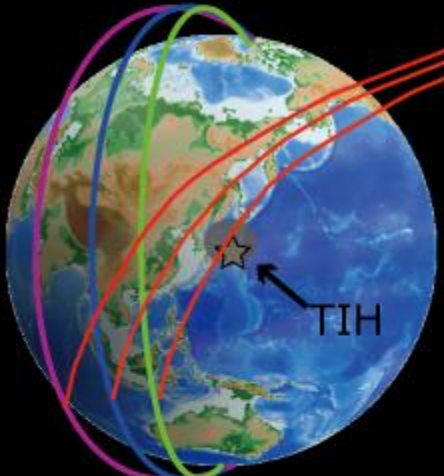








Occulting GPS satellite



Occulting LEO satellites

